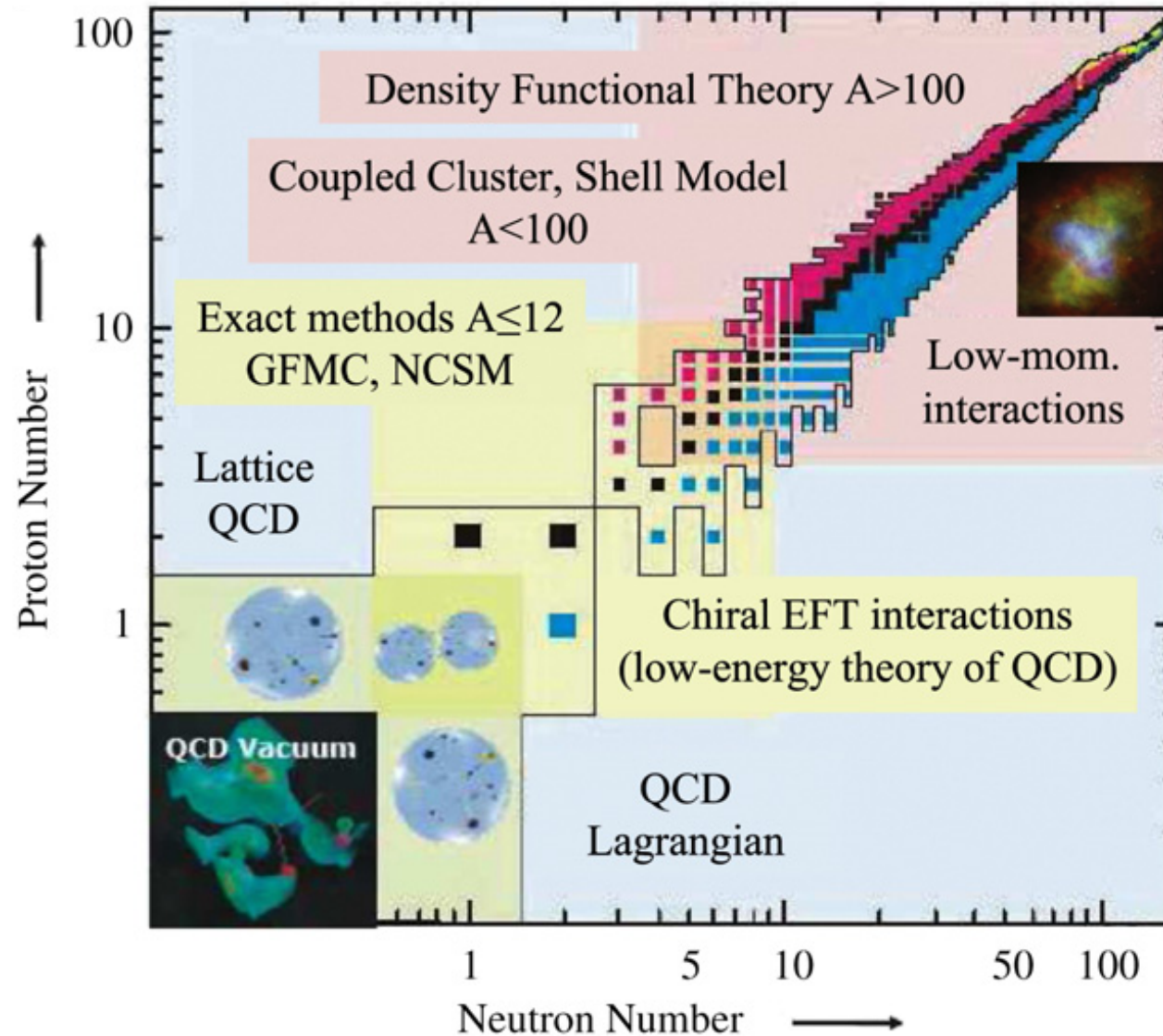


The Challenge of Ab Initio Nuclear Theory

To understand the properties of complex nuclei from first principles



Two significant issues:

Interaction

Not well understood

Not obtainable from QCD

Too “hard” to be useful

Multiple energy scales

Many-body Problem

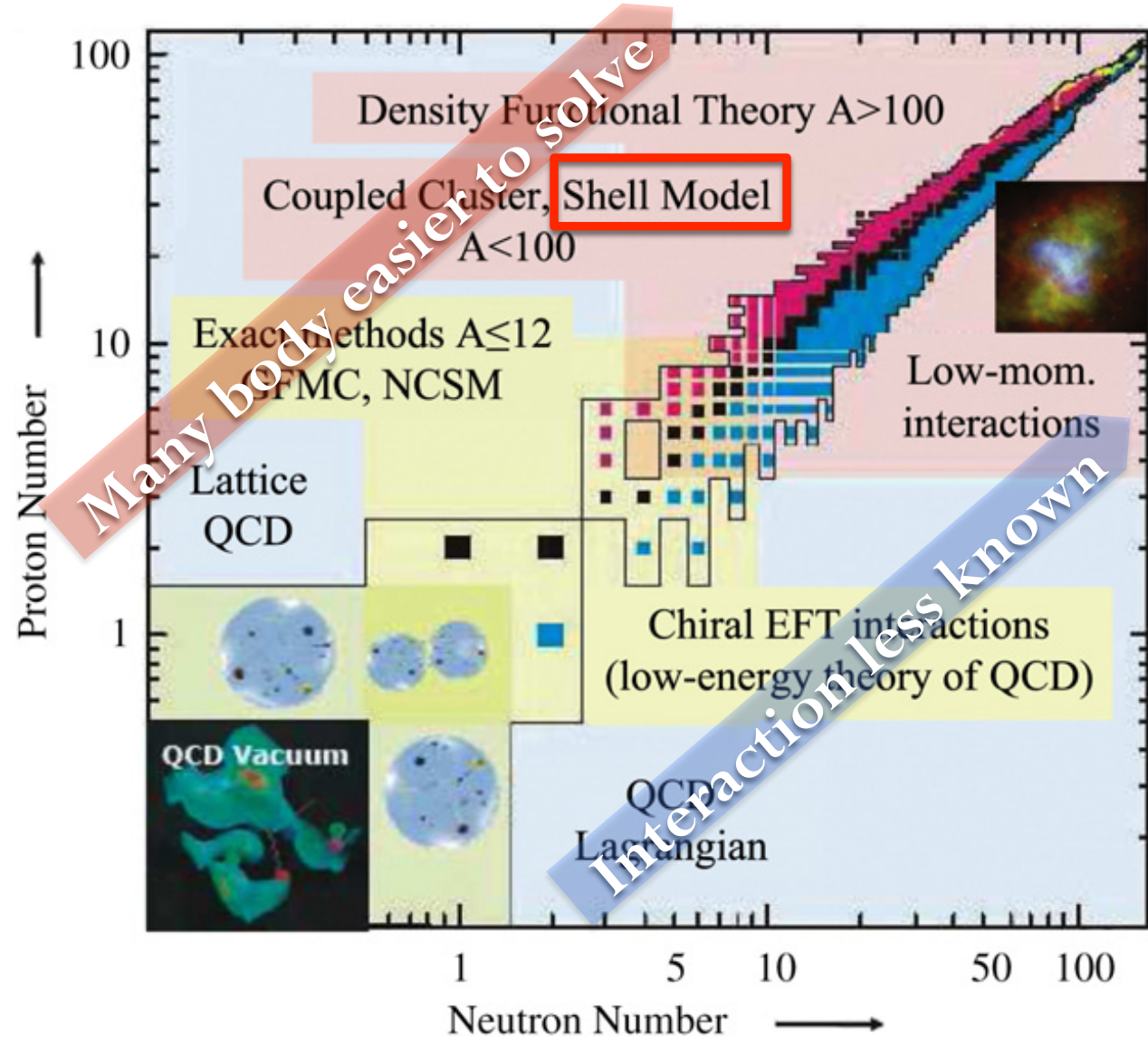
Not ‘exactly’ solvable above

$$A \sim 20$$

Here we focus on shell model

The Challenge of Ab Initio Nuclear Theory

To understand the properties of complex nuclei from first principles



Two significant issues:

Interaction

Not well understood

Not obtainable from QCD

Too "hard" to be useful

Multiple energy scales

Many-body Problem

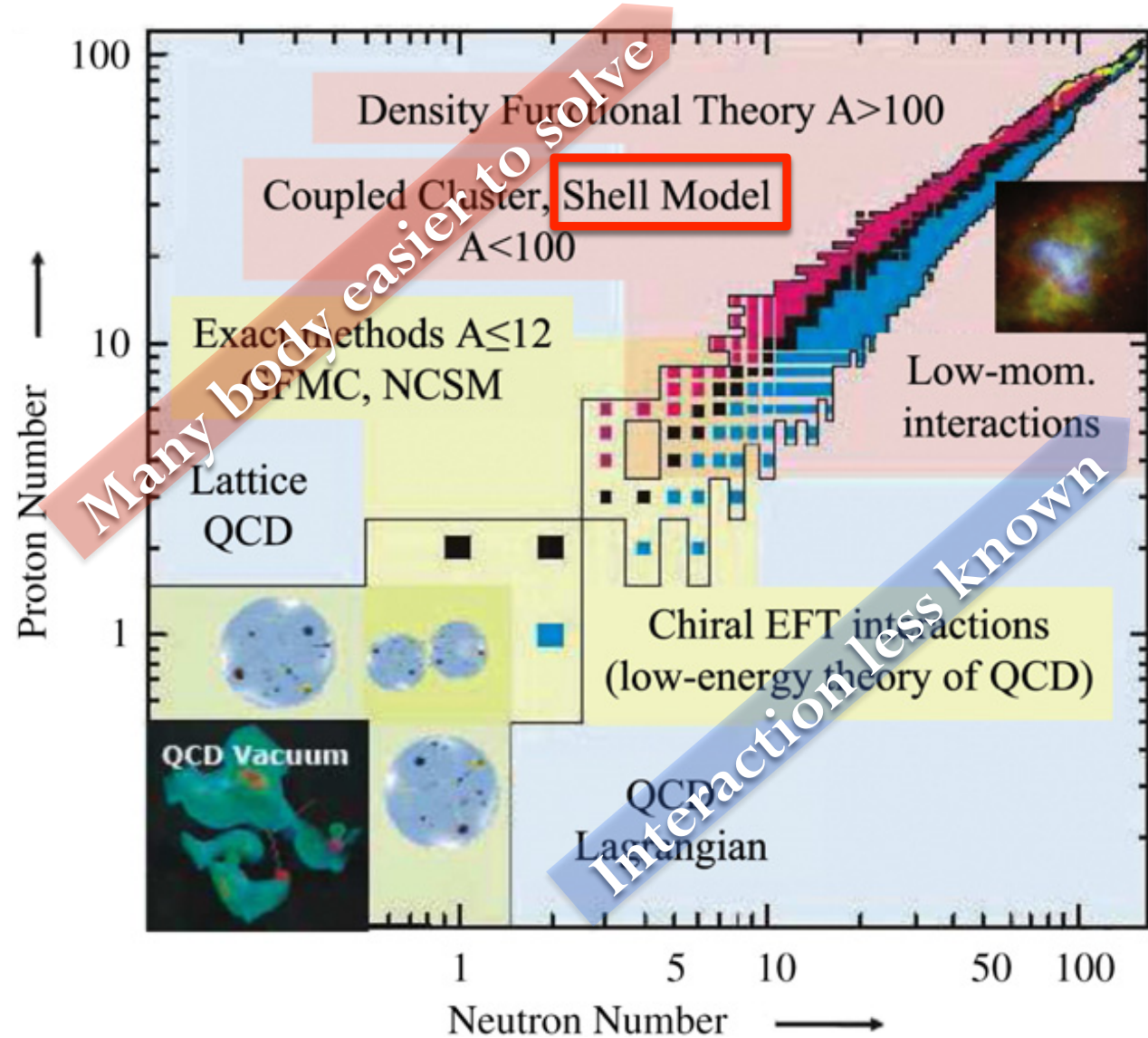
Not 'exactly' solvable above

$$A \sim 20$$

Here we focus on shell model

The Challenge of Ab Initio Nuclear Theory

To understand the properties of complex nuclei from first principles



Two significant issues:

Interaction

Not well understood
Not obtainable from QCD
Too “hard” to be useful
Multiple energy scales

Many-body Problem

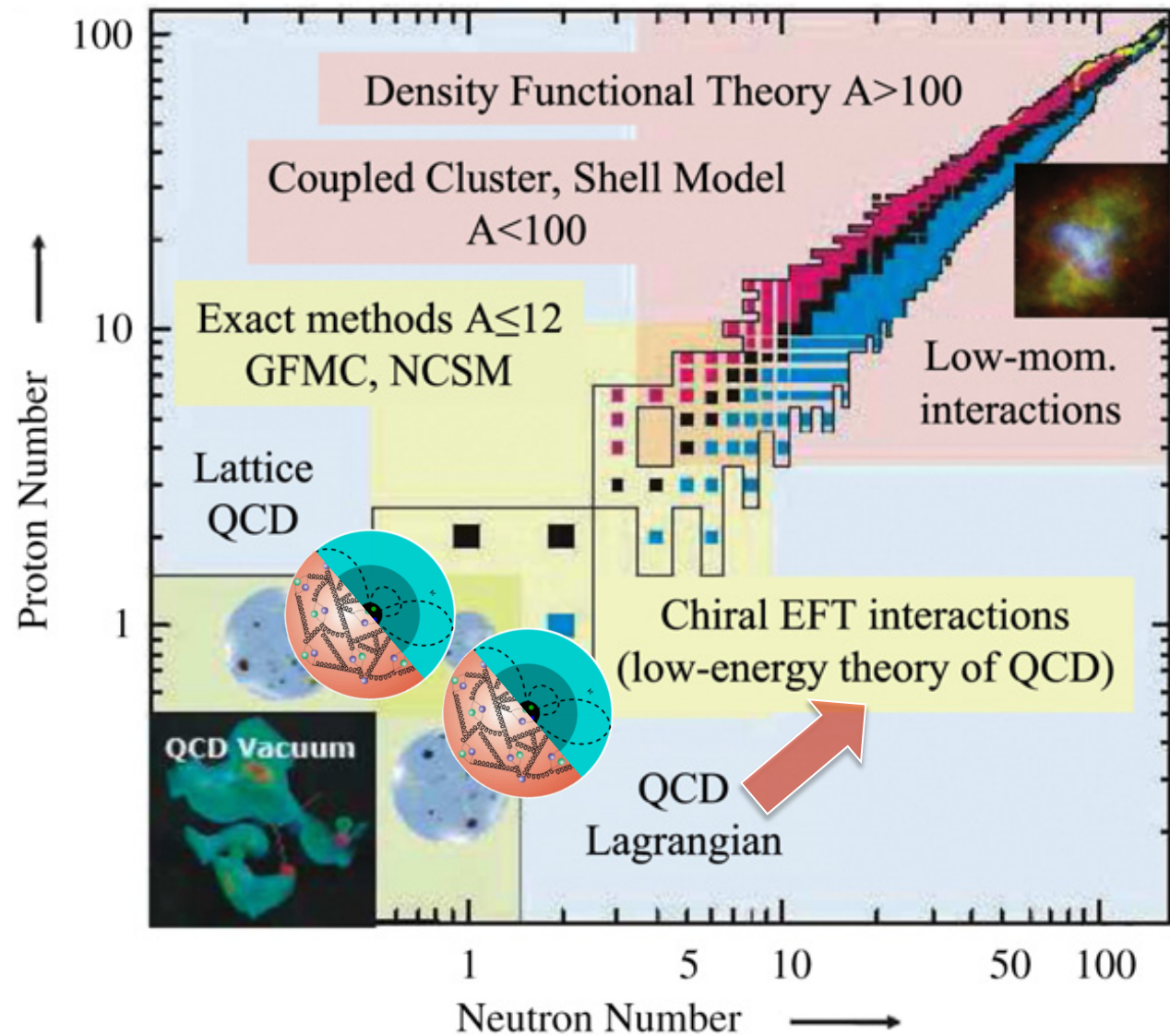
Not ‘exactly’ solvable above
 $A \sim 20$
Here we focus on shell model

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow “Solve” many-body problem \rightarrow Predictions

The Challenge of Ab Initio Nuclear Theory

To understand the properties of complex nuclei from first principles



Nucleon-nucleon interaction

Some history

Anatomy of an NN interaction

Construction from QCD?

Ideas of Effective Field Theory

Chiral EFT for nuclear forces

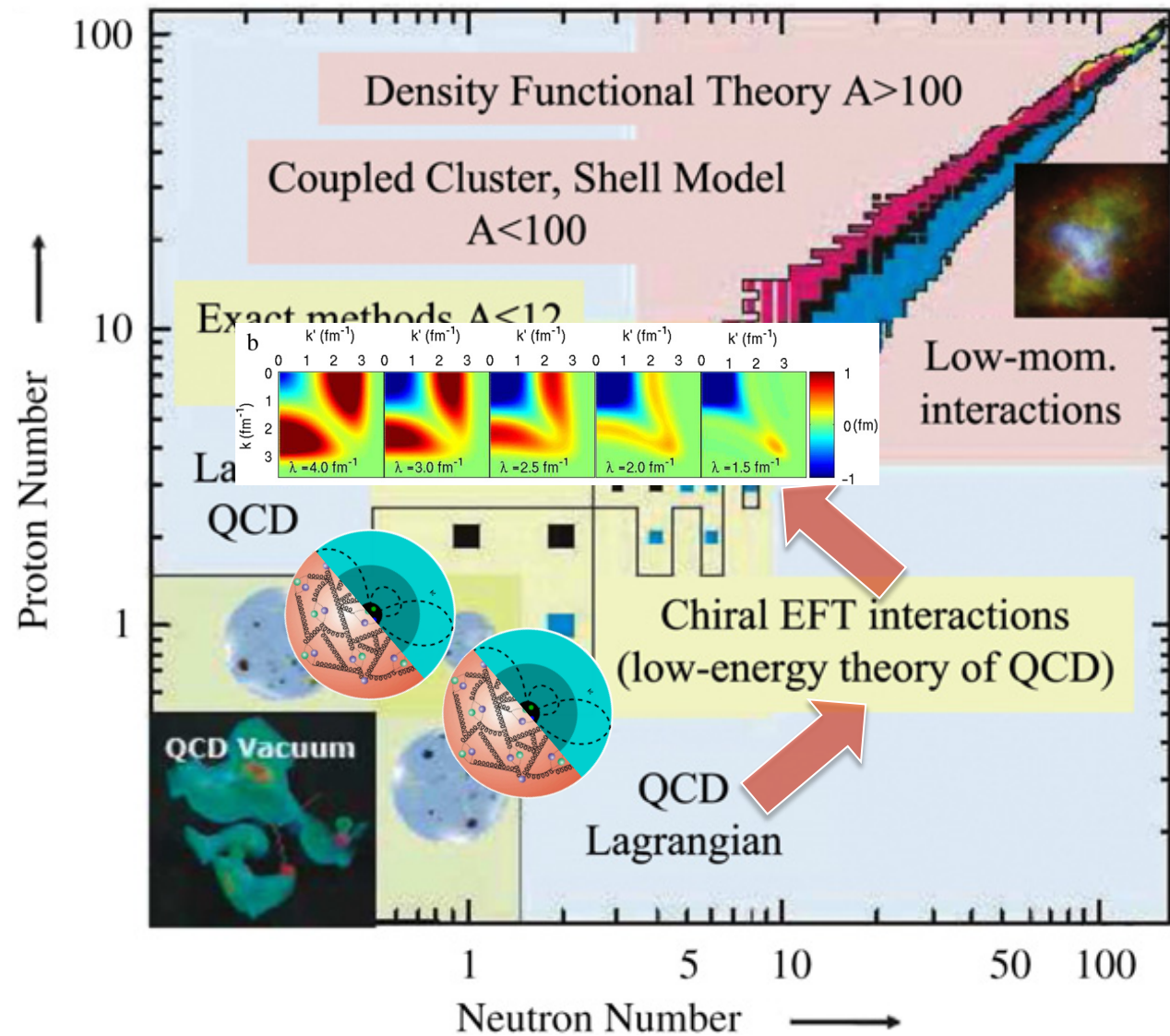
Constraint by data

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow "Solve" many-body problem \rightarrow Predictions

The Challenge of Ab Initio Nuclear Theory

To understand the properties of complex nuclei from first principles



Renormalizing NN Interactions

Basic ideas of RG

Low-momentum interactions

Similarity RG interactions

Benefits of low cutoffs

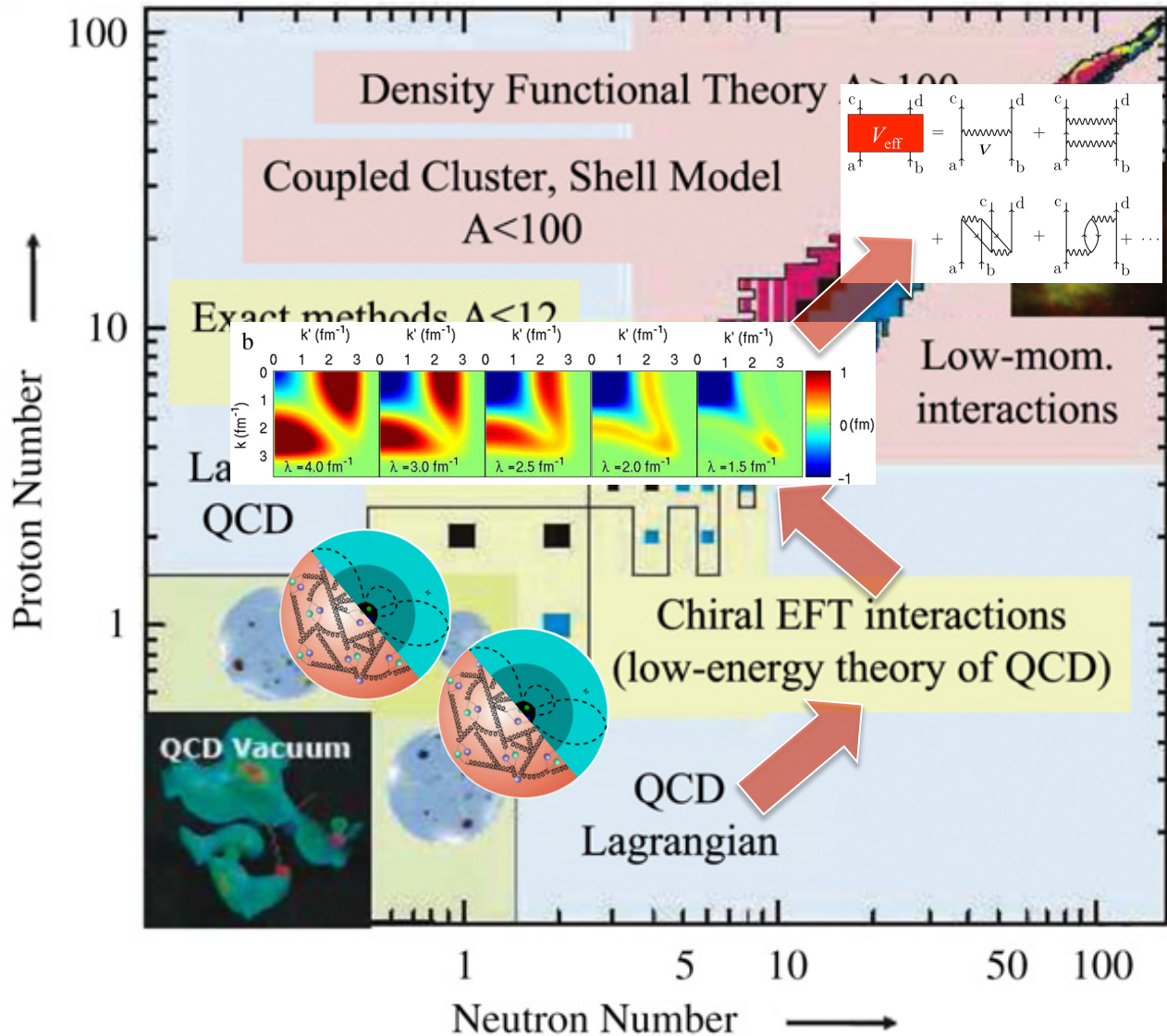
G-matrix renormalization

How will we approach this problem:

QCD → NN (3N) forces → Renormalize → “Solve” many-body problem → Predictions

The Challenge of Ab Initio Nuclear Theory

To understand the properties of complex nuclei from first principles



Microscopic Valence-Space Interactions

Model spaces

Many-body perturbation theory (MBPT)

Calculating effective interaction

In-medium Similarity RG

Monopole part of interaction

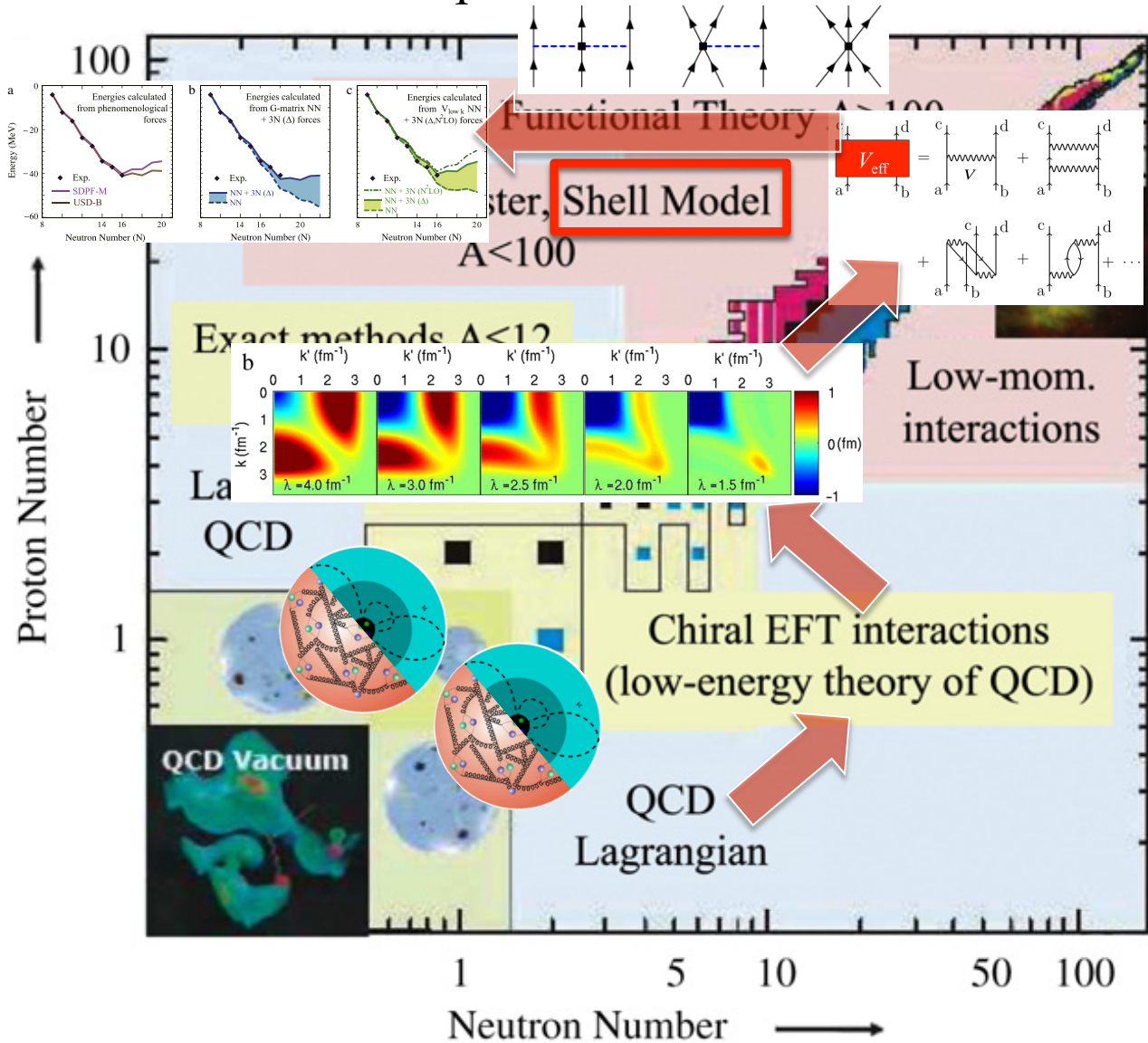
Deficiencies of this approach

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow "Solve" many-body problem \rightarrow Predictions

The Challenge of Ab Initio Nuclear Theory

To understand the properties of complex nuclei from first principles



Three-Nucleon Forces

Basic ideas – why needed?

3N from chiral EFT

Implementing in shell model

Relation to monopoles

Predictions/new discoveries

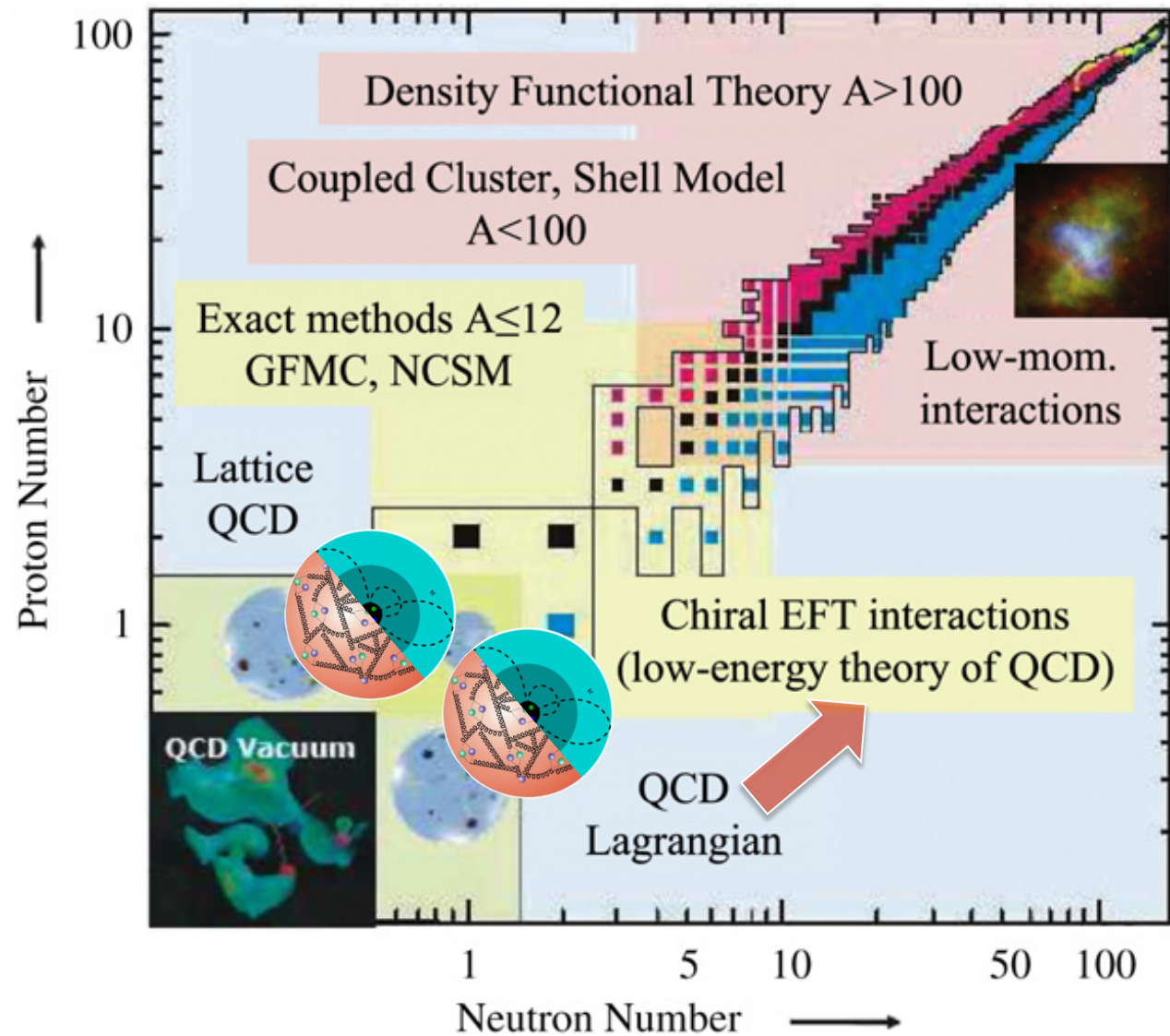
Connections beyond structure

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow “Solve” many-body problem \rightarrow Predictions

Part I: The Nucleon-Nucleon Interaction

To understand the properties of complex nuclei from first principles



Nucleon-nucleon interaction

Some history

Anatomy of an NN interaction

Construction from QCD?

Ideas of Effective Field Theory

Chiral EFT for nuclear forces

Constraint by data

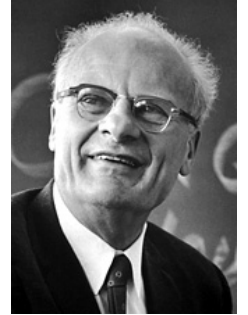
How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow Solve many-body problem \rightarrow Predictions

Interaction Between Two Nucleons

“In the past quarter century physicists have devoted a huge amount of experimentation and mental labor to this problem – probably more man-hours than have been given to any other scientific question in the history of mankind.”

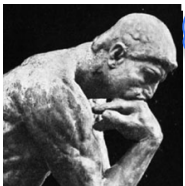
–*H. Bethe*



So let's burn a few more man-hours of mental labor on this!

To start, think to yourself what this should look like, and write it down...

Ok, the nuclear potential as
a function of the distance
between nucleons... Got it!



Meson-Exchange Potentials: Yukawa

- First field-theoretical model of nucleon interaction proposed by **Yukawa** 1935
- Postulated nuclear force mediated by (**NEW!**) particle exchange
- Short range ($\sim 1\text{fm}$) of nuclear force \implies



Meson-Exchange Potentials: Yukawa

- First field-theoretical model of nucleon interaction proposed by **Yukawa** 1935
- Postulated nuclear force mediated by (**NEW!**) particle exchange
- Short range ($\sim 1\text{fm}$) of nuclear force \implies

New particle must be massive: $r \sim 1/m$; $m = ?$

Hint: $\hbar c \approx 197 \text{ MeV} \cdot \text{fm}$



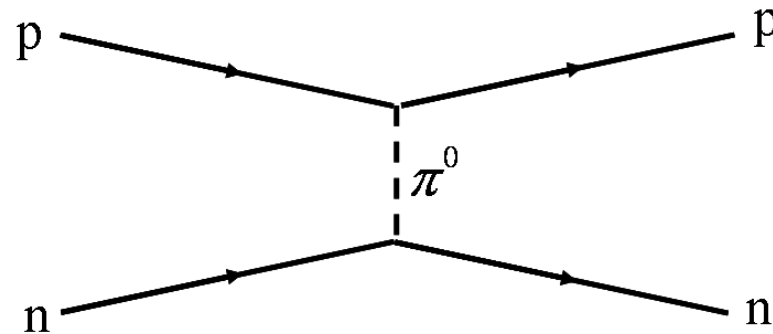
Meson-Exchange Potentials: Yukawa

- First field-theoretical model of nucleon interaction proposed by **Yukawa** 1935
- Postulated nuclear force mediated by (**NEW!**) particle exchange
- Short range ($\sim 1\text{fm}$) of nuclear force \implies

New particle must be massive: $r \sim 1/m$; $m = ?$



- Pion discovered 1947!



$$V(\vec{r}) = -\frac{f_\pi^2}{m_\pi^2} \left\{ \vec{\sigma}_1 \cdot \vec{\sigma}_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(r) \right\} \frac{e^{-m_\pi r}}{m_\pi r}$$

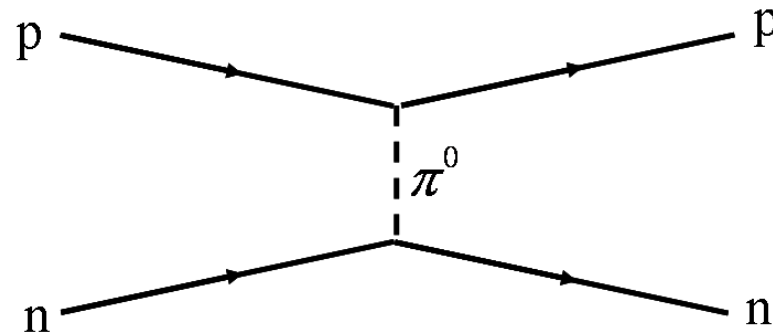
Meson-Exchange Potentials: Yukawa

- First field-theoretical model of nucleon interaction proposed by **Yukawa** 1935
- Postulated nuclear force mediated by (**NEW!**) particle exchange
- Short range ($\sim 1\text{fm}$) of nuclear force \implies

New particle must be massive: $r \sim 1/m$; $m = ?$



- Pion discovered 1947!



$$V(\vec{r}) = -\frac{f_\pi^2}{m_\pi^2} \left\{ \vec{\sigma}_1 \cdot \vec{\sigma}_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(r) \right\} \frac{e^{-m_\pi r}}{m_\pi r}$$

- **Attractive, “long” range**

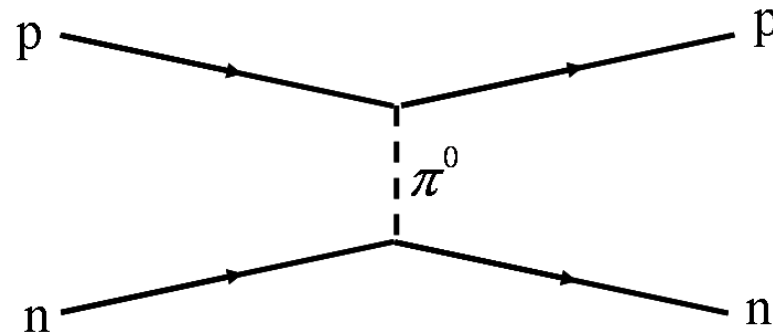
Meson-Exchange Potentials: Yukawa

- First field-theoretical model of nucleon interaction proposed by **Yukawa** 1935
- Postulated nuclear force mediated by (**NEW!**) particle exchange
- Short range ($\sim 1\text{fm}$) of nuclear force \implies

New particle must be massive: $r \sim 1/m$; $m = ?$



- Pion discovered 1947!



$$V(\vec{r}) = -\frac{f_\pi^2}{m_\pi^2} \left\{ \vec{\sigma}_1 \cdot \vec{\sigma}_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(r) \right\} \frac{e^{-m_\pi r}}{m_\pi r}$$

- Attractive, “long” range, **spin dependent**

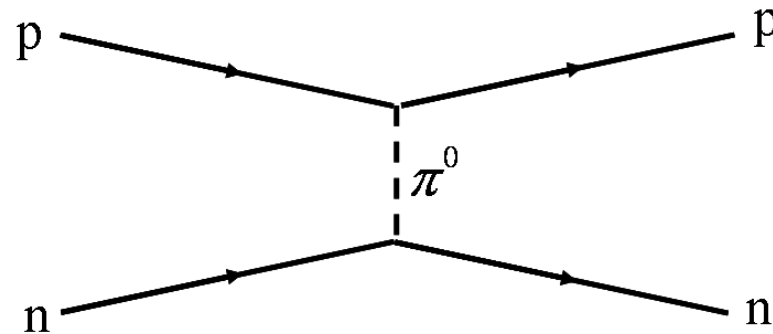
Meson-Exchange Potentials: Yukawa

- First field-theoretical model of nucleon interaction proposed by **Yukawa** 1935
- Postulated nuclear force mediated by (**NEW!**) particle exchange
- Short range ($\sim 1\text{fm}$) of nuclear force \implies

New particle must be massive: $r \sim 1/m$; $m = ?$



- Pion discovered 1947!



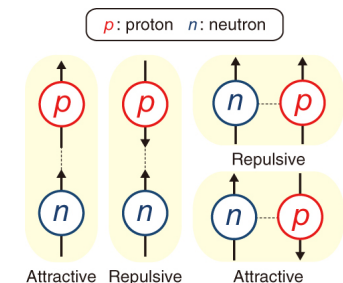
$$V(\vec{r}) = -\frac{f_\pi^2}{m_\pi^2} \left\{ \vec{\sigma}_1 \cdot \vec{\sigma}_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(r) \right\} \frac{e^{-m_\pi r}}{m_\pi r}$$

- Attractive, “long” range, spin dependent, **non-central (tensor) part**

Depends on spin, isospin, orientation of nucleons

Does not conserve L^2 , S^2 , but does conserve parity

\implies Mixes different L states (but only differing by 2 units)



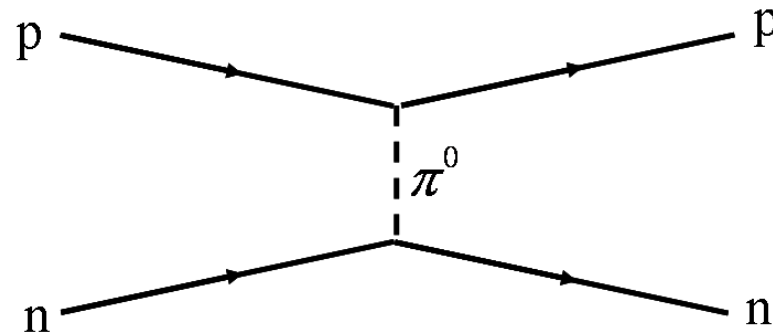
Meson-Exchange Potentials: Yukawa

- First field-theoretical model of nucleon interaction proposed by **Yukawa** 1935
- Postulated nuclear force mediated by (**NEW!**) particle exchange
- Short range ($\sim 1\text{fm}$) of nuclear force \implies

New particle must be massive: $r \sim 1/m$; $m = ?$



- Pion discovered 1947!

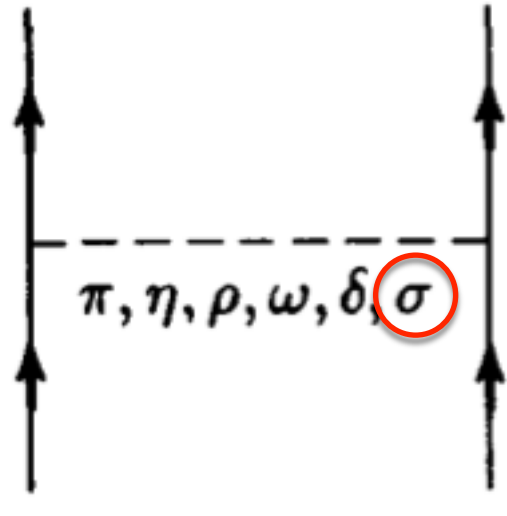


$$V(\vec{r}) = -\frac{f_\pi^2}{m_\pi^2} \left\{ \vec{\sigma}_1 \cdot \vec{\sigma}_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(r) \right\} \frac{e^{-m_\pi r}}{m_\pi r}$$

- **Attractive, “long” range, spin dependent, non-central (tensor) part**
- Successful in explaining scattering data, deuteron
- One pion is good, therefore more pions are better...
- Advanced to multi-pion theories in 1950's – **FAILED! Now what??**

One-Boson Exchange Potentials

- Heavy mesons discovered in late 1950s – formed basis for new theories
- Intermediate range – **attractive central, spin-orbit**



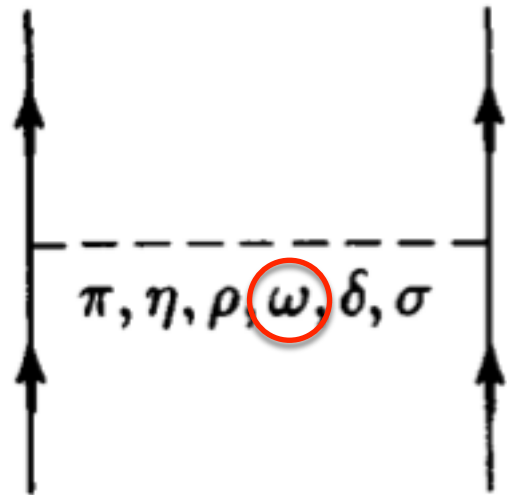
$$V^\sigma = g_{\sigma NN}^2 \frac{1}{\mathbf{k}^2 + m_\sigma^2} \left(-1 + \frac{\mathbf{q}^2}{2M_N^2} - \frac{\mathbf{k}^2}{8M_N^2} - \frac{\vec{L} \cdot \vec{S}}{2M_N^2} \right)$$

$$\vec{q}_i \equiv \vec{p}'_i - \vec{p}_i \quad \vec{k}_i \equiv \frac{1}{2} (\vec{p}'_i + \vec{p}_i)$$

Baryons	Mass (MeV)	Mesons	Mass (MeV)
p, n	938.926	π	138.03
Λ	1116.0	η	548.8
Σ	1197.3	σ	≈ 550.0
Δ	1232.0	ρ	770
Σ^*	1385.0	ω	782.6
		δ	983.0
		K	495.8
		K*	895.0

One-Boson Exchange Potentials

- Heavy mesons discovered in late 1950s – formed basis for new theories
- Short range; **repulsive central force, attractive spin-orbit**

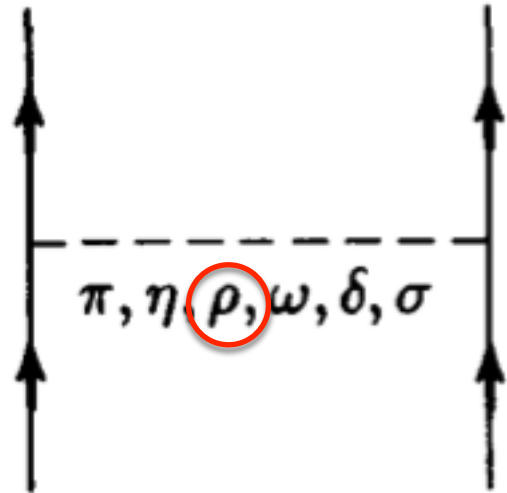


$$V^\omega = g_{\omega NN}^2 \frac{1}{\mathbf{k}^2 + m_\omega^2} \left(1 - 3 \frac{\vec{L} \cdot \vec{S}}{2M_N^2} \right)$$

Baryons	Mass (MeV)	Mesons	Mass (MeV)
p, n	938.926	π	138.03
Λ	1116.0	η	548.8
Σ	1197.3	σ	\approx 550.0
Δ	1232.0	ρ	770
Σ^*	1385.0	ω	782.6
		δ	983.0
		K	495.8
		K*	895.0

One-Boson Exchange Potentials

- Heavy mesons discovered in late 1950s – formed basis for new theories
- Short range; **tensor force opposite sign of one-pion exchange**



$$V^\rho = g_{\rho NN}^2 \frac{\mathbf{k}^2}{\mathbf{k}^2 + m_\rho^2} \left(-2 \vec{\sigma}_1 \cdot \vec{\sigma}_2 + S_{12}(\hat{k}) \right) \vec{\tau}_1 \cdot \vec{\tau}_2$$

Baryons	Mass (MeV)	Mesons	Mass (MeV)
p, n	938.926	π	138.03
Λ	1116.0	η	548.8
Σ	1197.3	σ	≈ 550.0
Δ	1232.0	ρ	770
Σ^*	1385.0	ω	782.6
		δ	983.0
		K	495.8
		K^*	895.0

Parameterizing the NN Interaction

Starting from any NN-interaction, first solve:

Lipmann-Schwinger scattering T-matrix equation:

$$T_{ll'}^\alpha(k, k'; K) = V_{ll'}^\alpha(k, k') + \frac{2}{\pi} \sum_{l''} \int_0^\infty q^2 dq V_{ll''}^\alpha(k, q) \frac{q}{k^2 - q^2 + i\varepsilon} T_{l''l'}^\alpha(q, k'; K)$$

where $T_{ll'}^\alpha(k, k'; K) = \langle kK, lL; \text{JST} | T | k'K, l'L; \text{JST} \rangle$

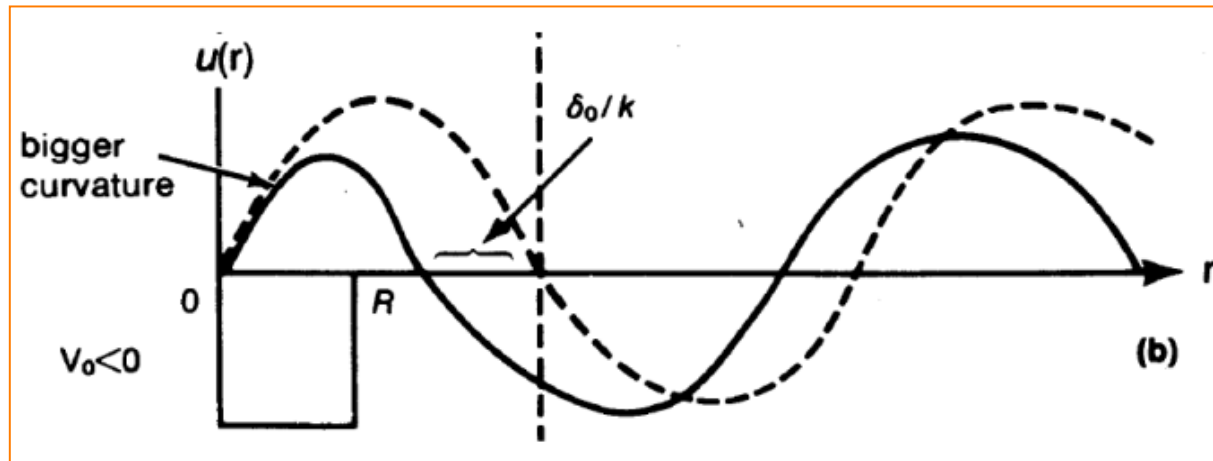
Parameterized in partial waves α in **relative/center of mass frame (K,L)**

$$\tan \delta(k) = -kT(k, k)$$

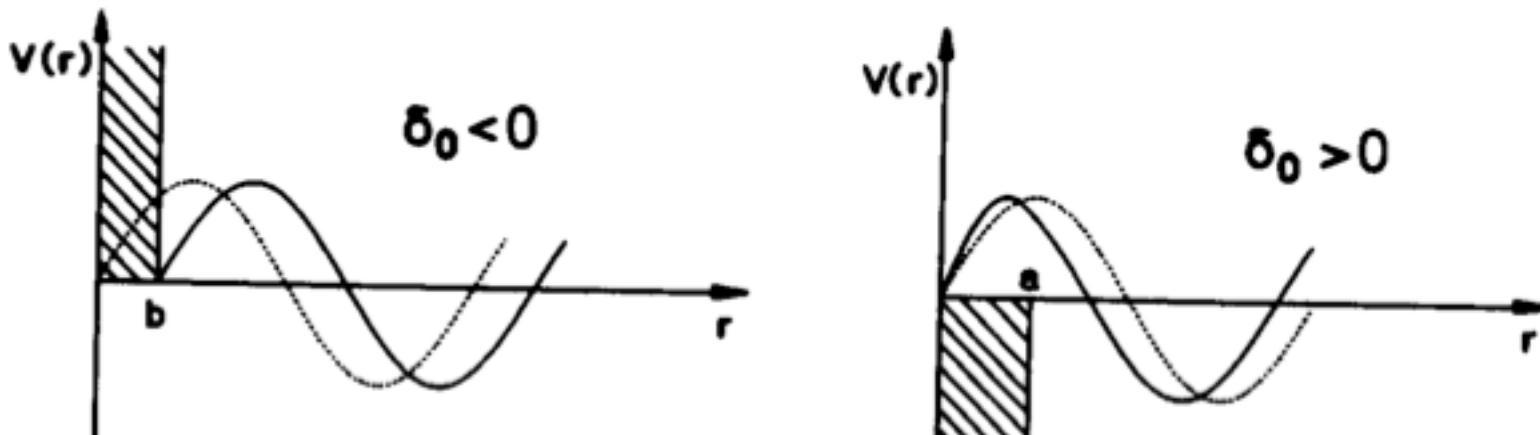
Fully-on-shell T-matrix directly related to experimental data

Constraining NN Scattering Phase Shifts

Phase shift is a function of relative momentum k ; Figure shows s -wave Scattering from an attractive well potential



Scattering from repulsive core: phase shift opposite sign



Parameterizing the NN Interaction

Starting from any NN-interaction, first solve:

Lipmann-Schwinger scattering T-matrix equation:

$$T_{ll'}^\alpha(k, k'; K) = V_{ll'}^\alpha(k, k') + \frac{2}{\pi} \sum_{l''} \int_0^\infty q^2 dq V_{ll''}^\alpha(k, q) \frac{q}{k^2 - q^2 + i\varepsilon} T_{l''l'}^\alpha(q, k'; K)$$

where $T_{ll'}^\alpha(k, k'; K) = \langle kK, lL; JST | T | k'K, l'L; JST \rangle$

Parameterized in partial waves α – in relative/center of mass frame (K,L)

$$\tan \delta(k) = -kT(k, k)$$

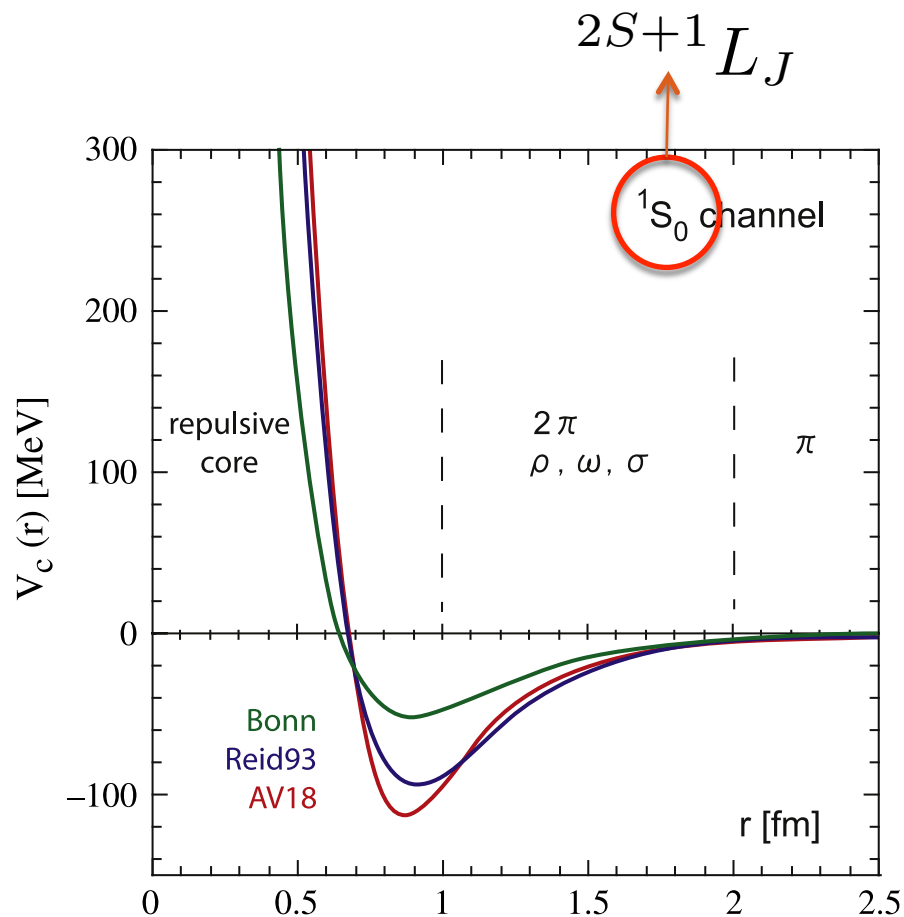
Fully-on-shell T -matrix directly related to experimental data

Note intermediate momentum allowed to infinity (but cutoff by regulators)

Coupling of low-to-high momentum in V

Form of NN Interactions

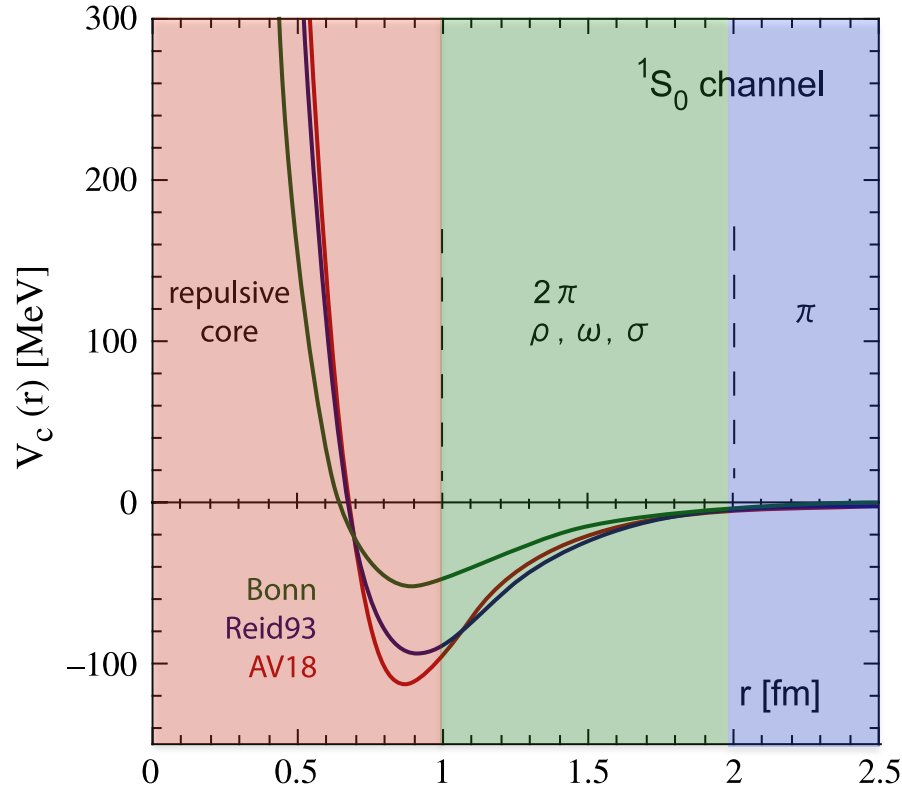
Textbook nuclear potentials in coordinate (\mathbf{r}) space (distance between nucleons)



Form of NN Interactions

Textbook nuclear potentials in coordinate (\mathbf{r}) space (distance between nucleons)

Hard core, **intermediate-range 2π** , **long-range 1π exchange (OPE)**



Form of NN Interactions

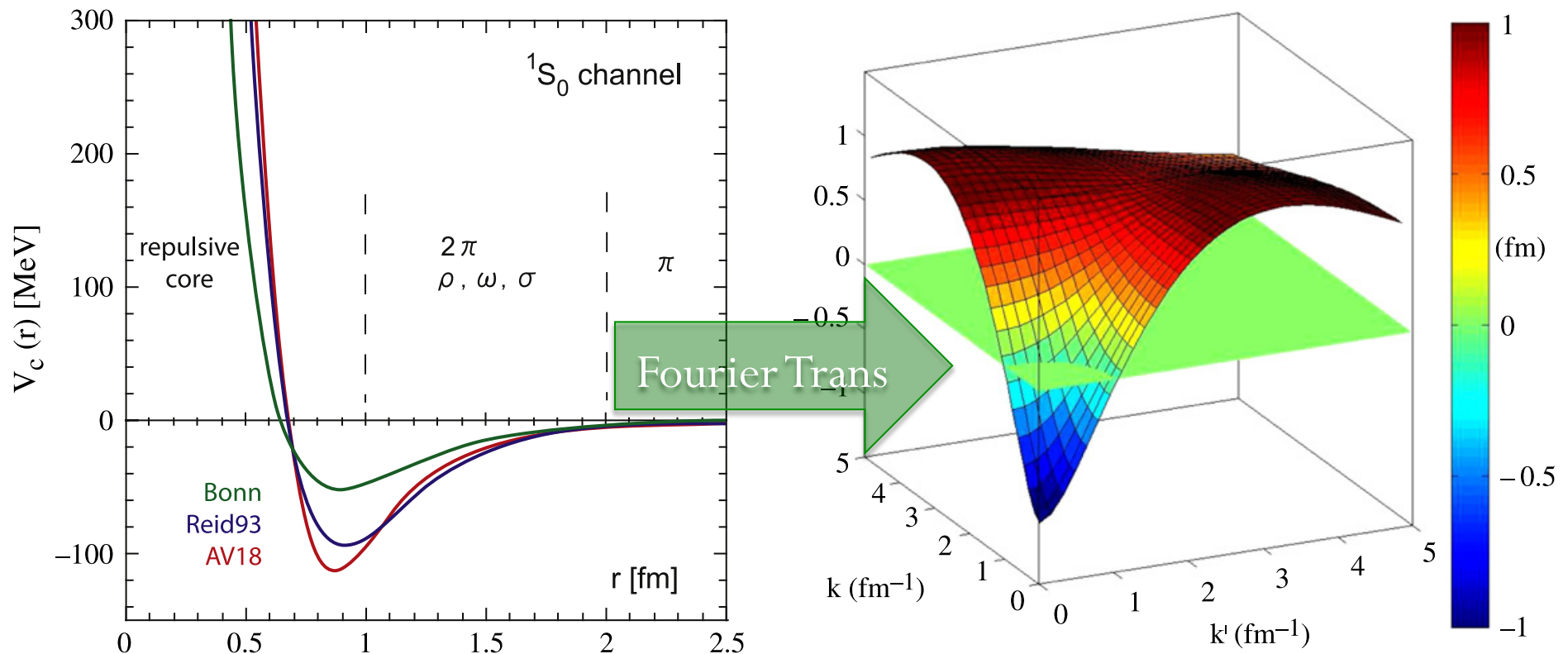
Textbook nuclear potentials in coordinate (\mathbf{r}) space (distance between nucleons)

Hard core, intermediate-range 2π , long-range 1π exchange

Transform to momentum space via **Fourier Transformation**

Strong high-momentum repulsion, **low-momentum attraction**

$$V_l(k, k') = \frac{2}{\pi} \int_0^\infty r^2 dr j_l(kr) V(r) j_l(k'r)$$



Form of NN Interactions

Textbook nuclear potentials in coordinate (\mathbf{r}) space (distance between nucleons)

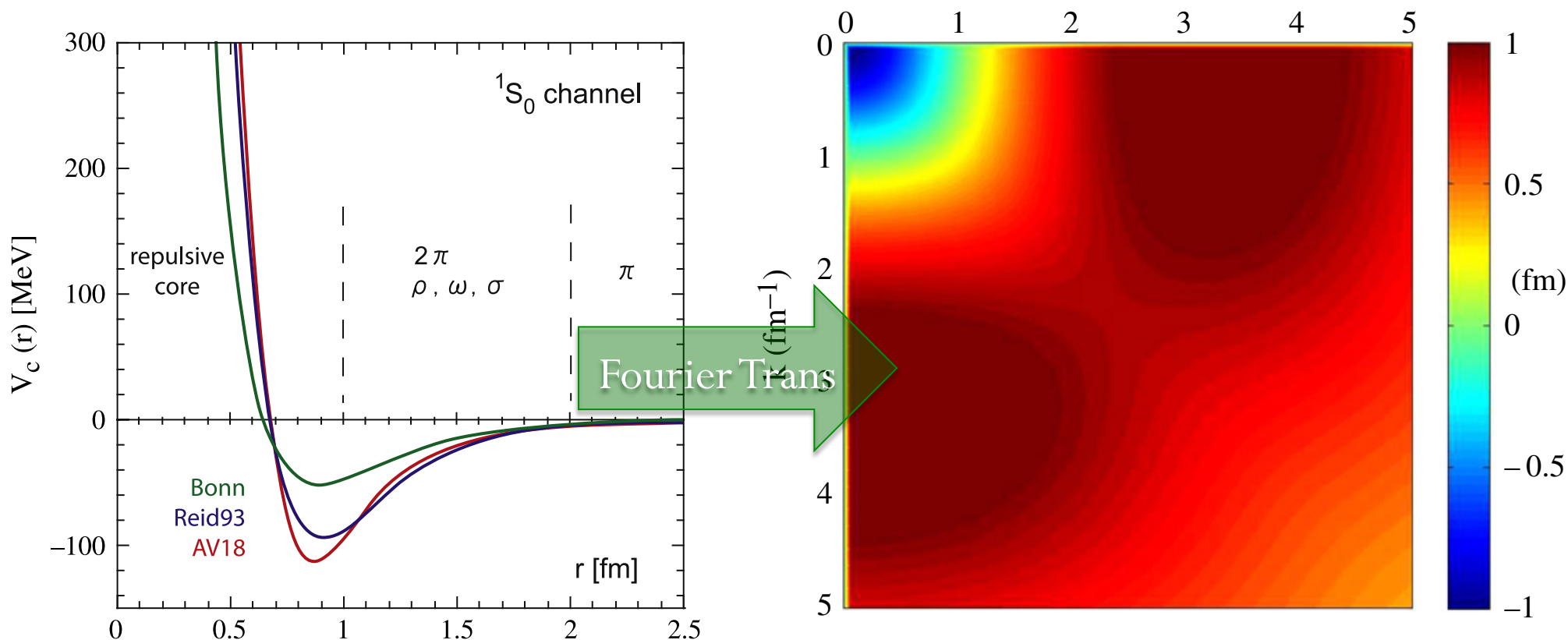
Hard core, intermediate-range 2π , long-range 1π exchange

Transform to momentum space via **Fourier Transformation**

Strong high-momentum repulsion, low-momentum attraction

$$V_l(k, k') = \frac{2}{\pi} \int_0^\infty r^2 dr j_l(kr) V(r) j_l(k'r)$$

$k' \text{ (fm}^{-1}\text{)}$



Form of NN Interactions

Textbook nuclear potentials in coordinate (\mathbf{r}) space (distance between nucleons)

Hard core, intermediate-range 2π , long-range 1π exchange

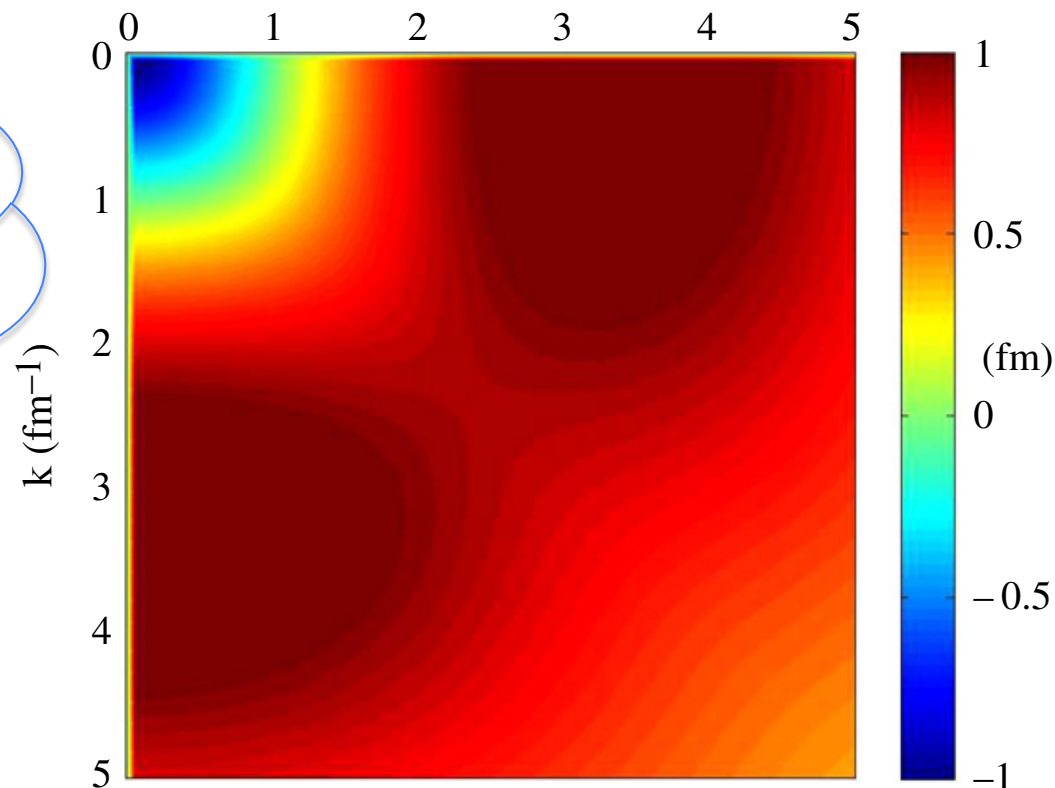
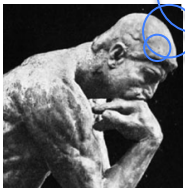
Transform to momentum space via **Fourier Transformation**

Strong high-momentum repulsion, **low-momentum attraction**

$$V_l(k, k') = \frac{2}{\pi} \int_0^\infty r^2 dr j_l(kr) V(r) j_l(k'r)$$

$k' \text{ (fm}^{-1}\text{)}$

Wait a minute... these potentials can't really go to zero range/infinately high energies; that would be QCD?

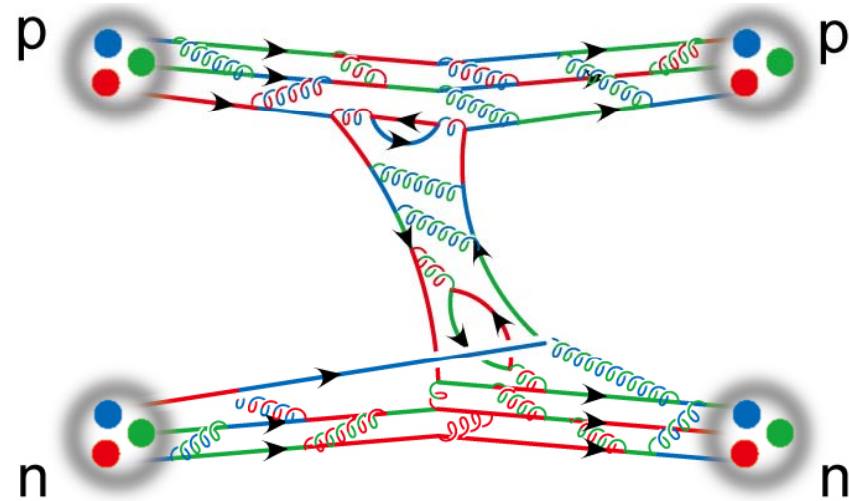
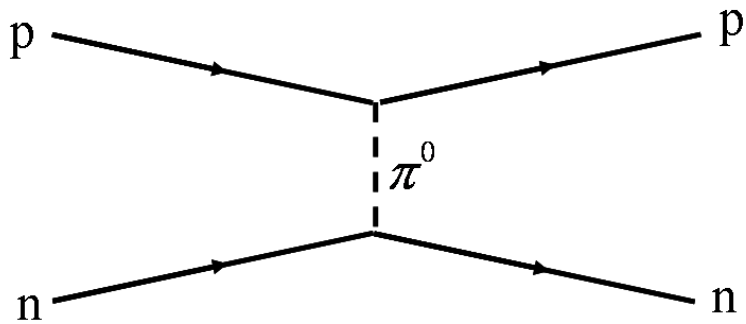


NN Interaction from QCD?

Meson exchange in principle described in Quantum Chromodynamics (QCD)

Low-energy region non-perturbative – treat in the context of **Lattice QCD**

Directly from QCD Lagrangian, solve numerically on discretized space-time

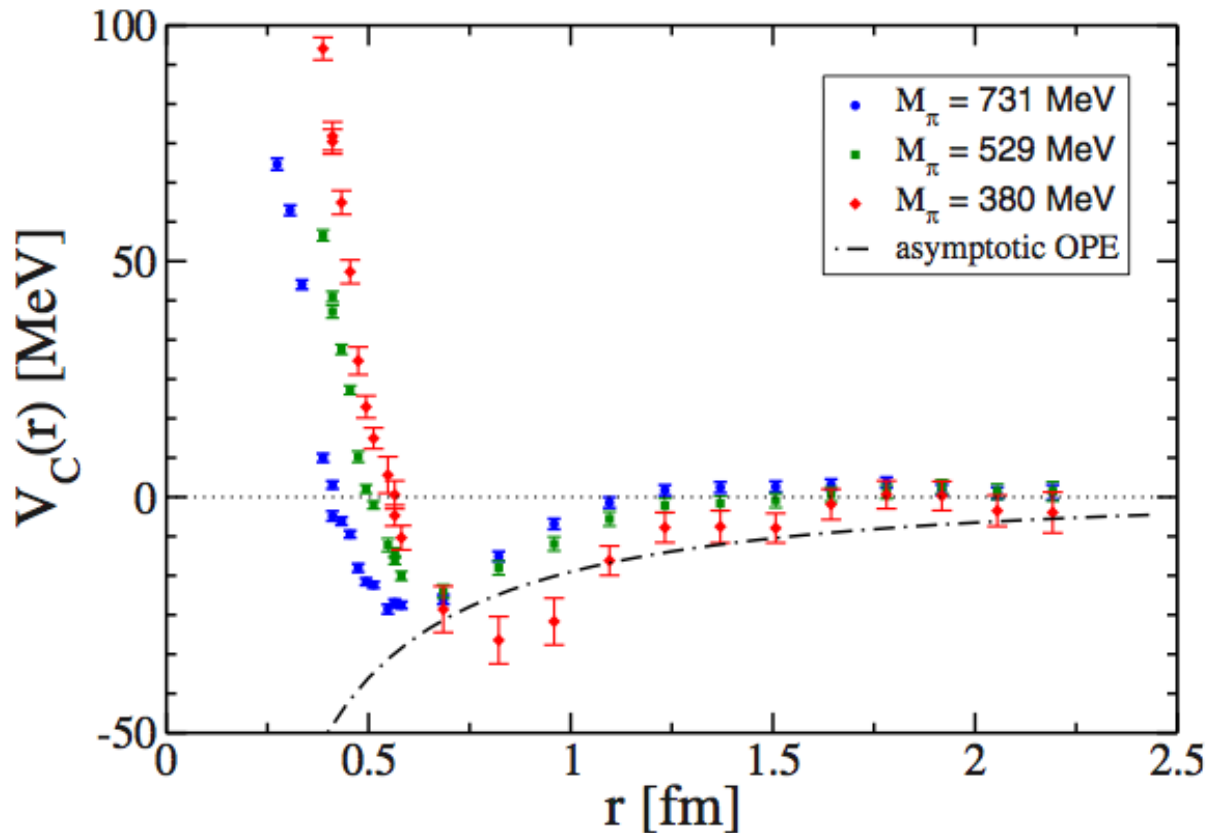


NN Interaction from QCD?

Meson exchange described in QCD

Low-energy region non-perturbative – treat in the context of **Lattice QCD**

Directly from QCD Lagrangian, solve numerically on discretized space-time



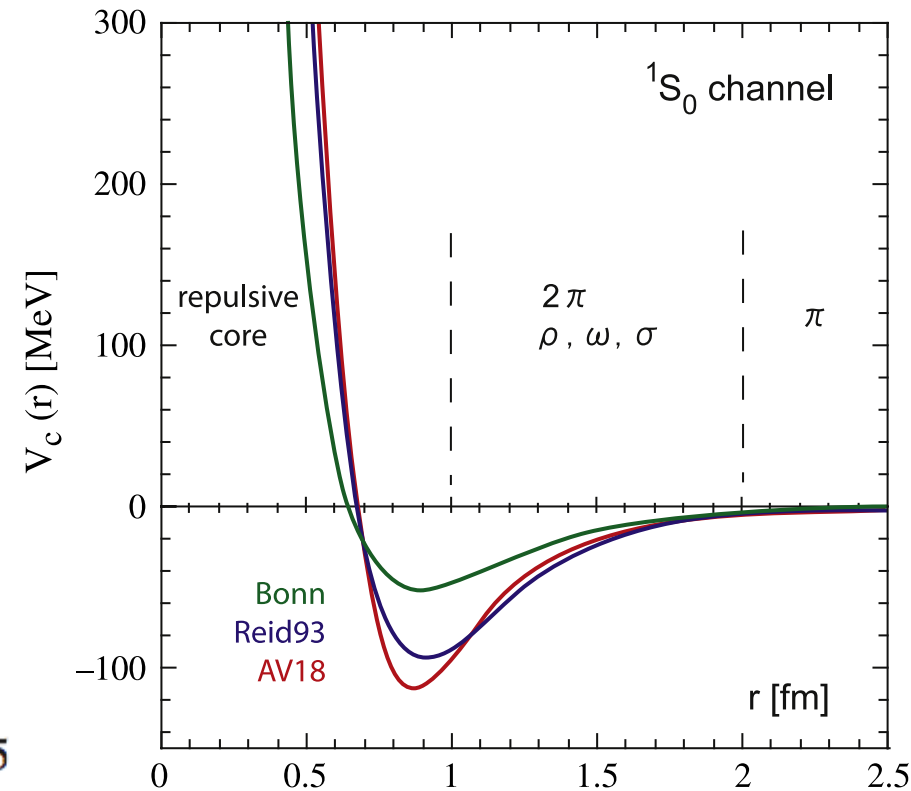
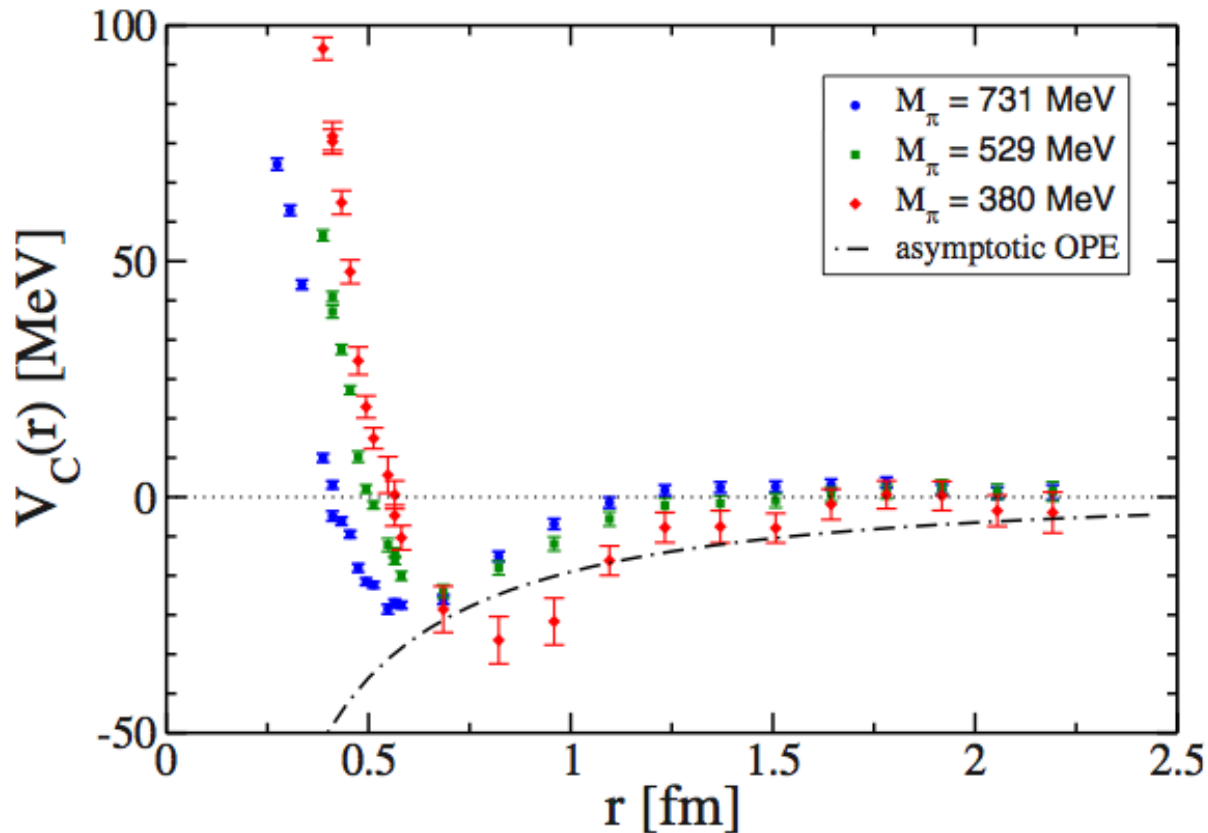
Lattice results give long-range OPE tail, hard core

NN Interaction from QCD?

Meson exchange described in QCD

Low-energy region non-perturbative – treat in the context of **Lattice QCD**

Directly from QCD Lagrangian, solve numerically on discretized space-time



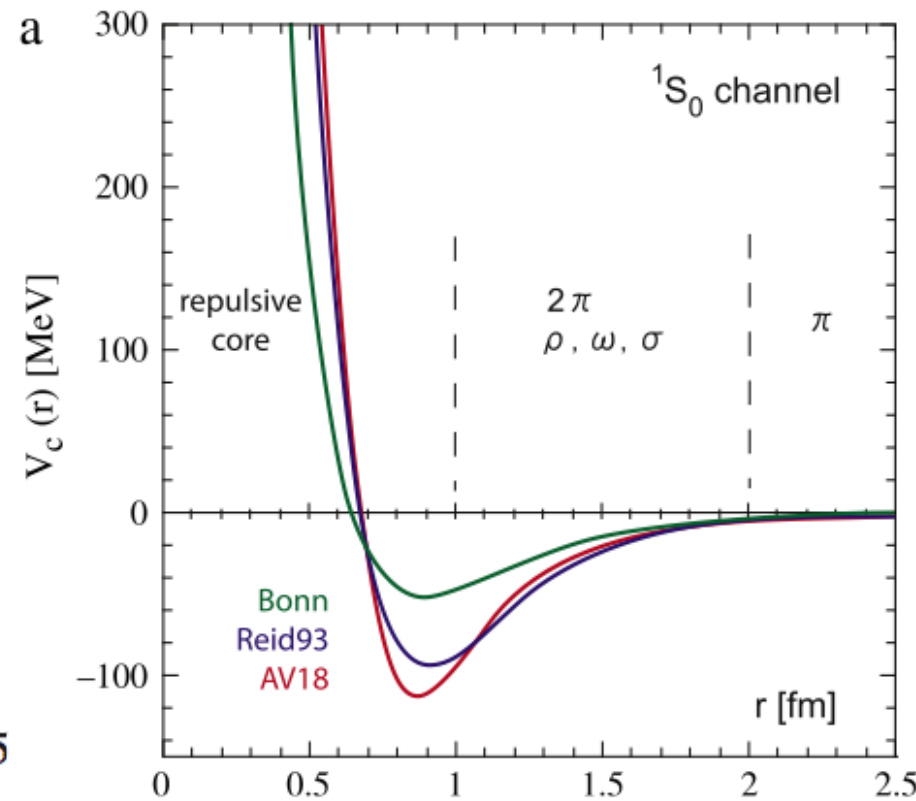
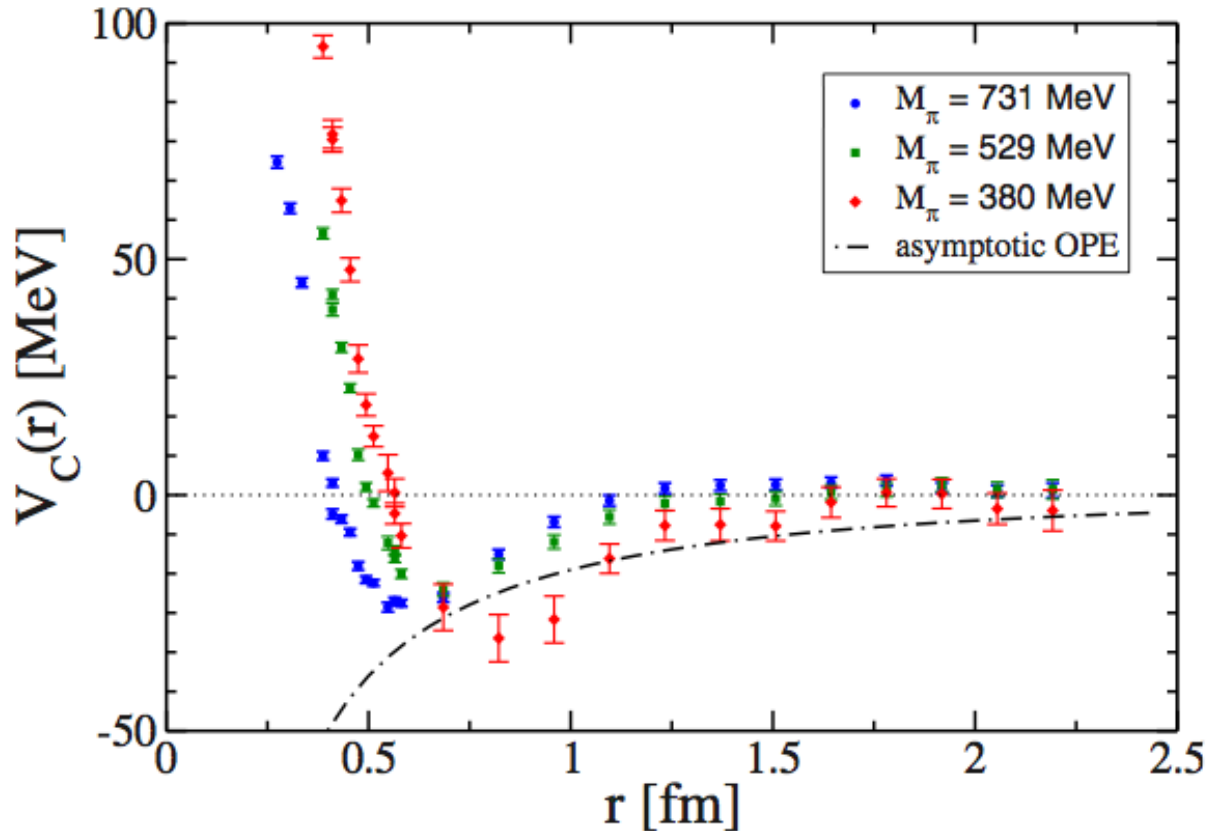
Lattice results give long-range OPE tail, hard core

Not yet to physical pion mass – work in progress – so we're done, right?

Unique NN Potential?

What does this tell us in our quest for an NN-potential?

Expected form seems to be confirmed by QCD



OBE Potentials: Summary/Problems

First generation (1960-1990): Paris, Reid, Bonn-A,B,C $\chi^2/\text{dof} \approx 2$

High-precision potentials (1990s): Focus on precision ~ 40 parameters fit NN data

ArgonneV18, Reid93, Nijmegen, CD-Bonn $\chi^2/\text{dof} \approx 1$

NN problem “solved” !!

OBE Potentials: Summary/Problems

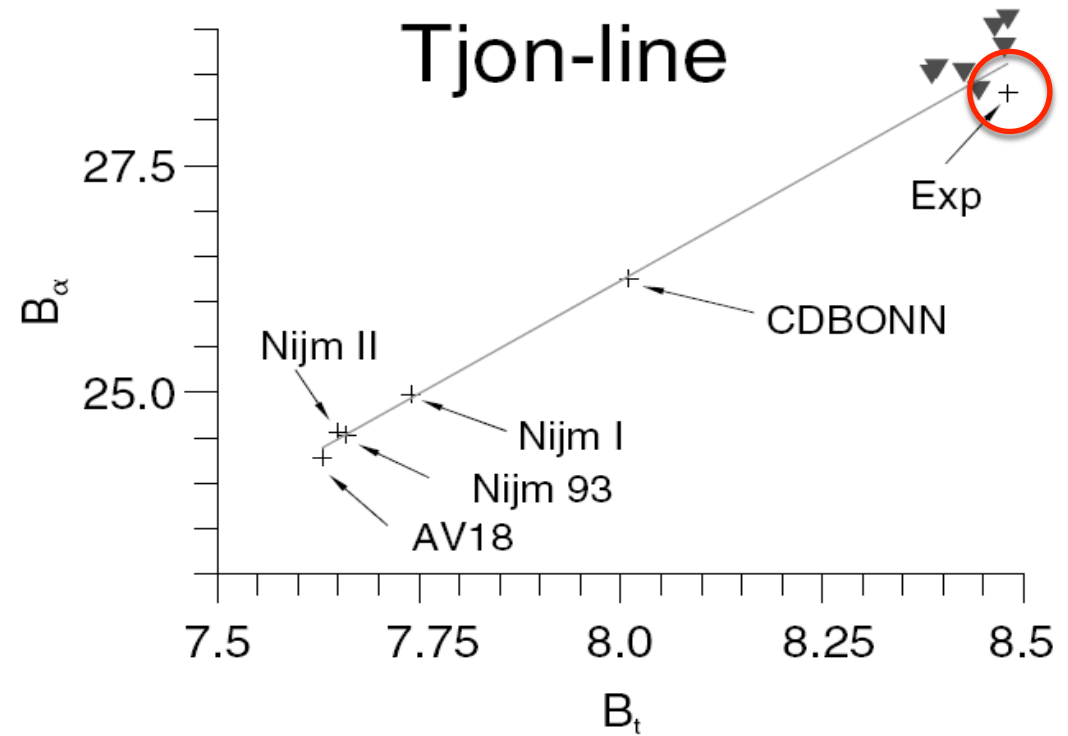
First generation (1960-1990): Paris, Reid, Bonn-A,B,C $\chi^2/\text{dof} \approx 2$

High-precision potentials (1990s): Focus on precision ~ 40 parameters fit NN data

ArgonneV18, Reid93, Nijmegen, CD-Bonn $\chi^2/\text{dof} \approx 1$

NN problem “solved” !!

Hmmm... what if we
were to go beyond
two nucleons?



OBE Potentials: Summary/Problems

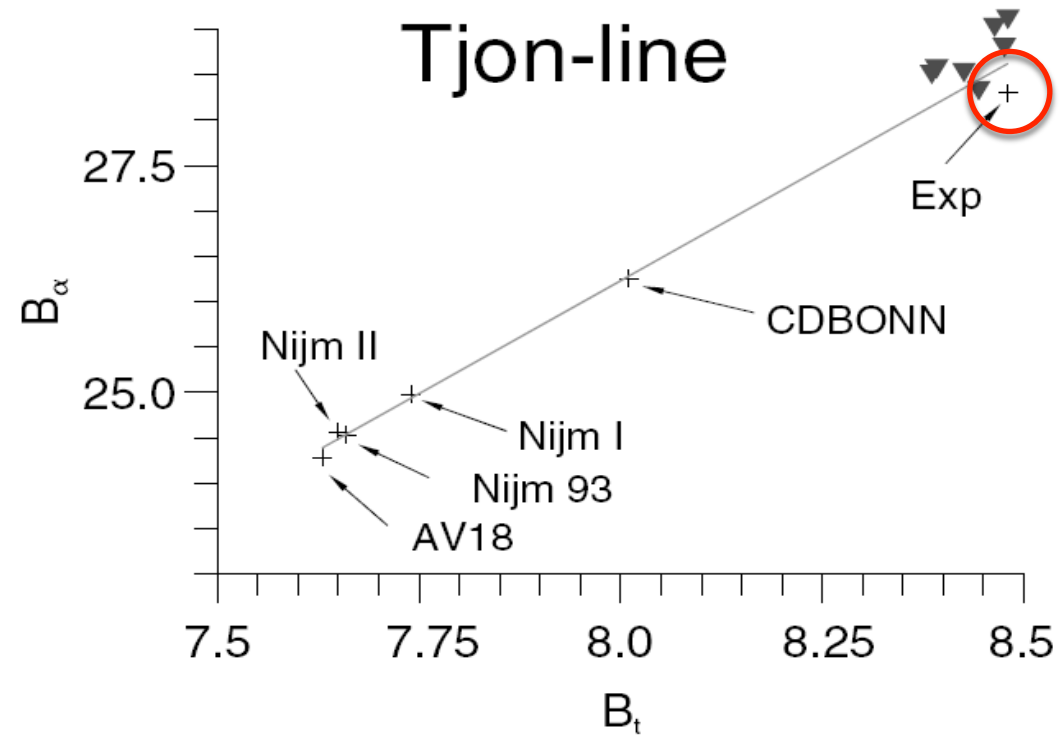
First generation (1960-1990): Paris, Reid, Bonn-A,B,C $\chi^2/\text{dof} \approx 2$

High-precision potentials (1990s): Focus on precision ~ 40 parameters fit NN data

ArgonneV18, Reid93, Nijmegen, CD-Bonn $\chi^2/\text{dof} \approx 1$

NN problem “solved” !!

Hmmm... what if we
were to go beyond
two nucleons?

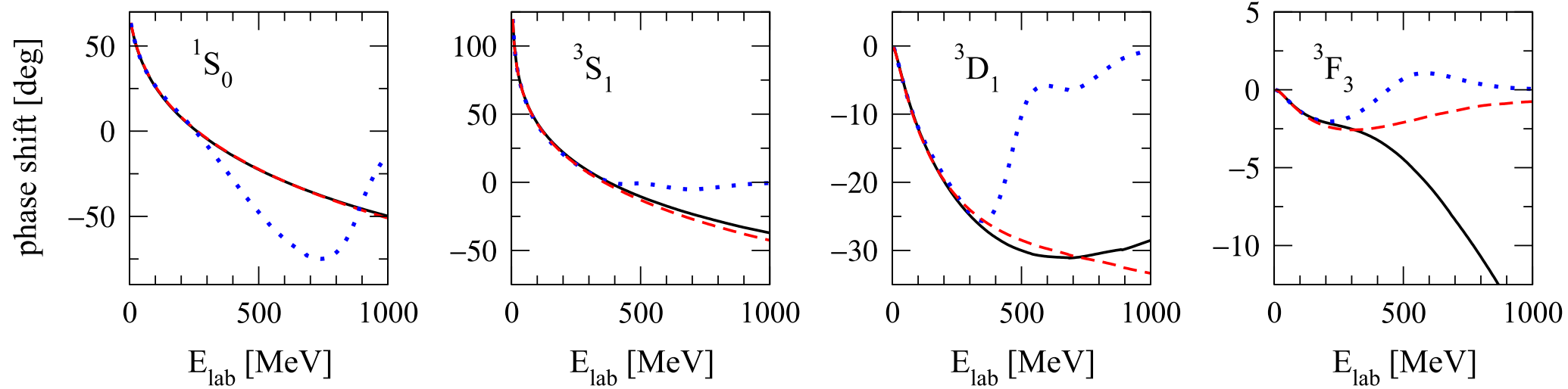


Many successes, but...

- 1) Difficult (impossible) to assign theoretical error
- 2) 3N forces (what are those??) not consistent with NN forces
- 3) No clear connection to QCD
- 4) Clear **model dependence**...

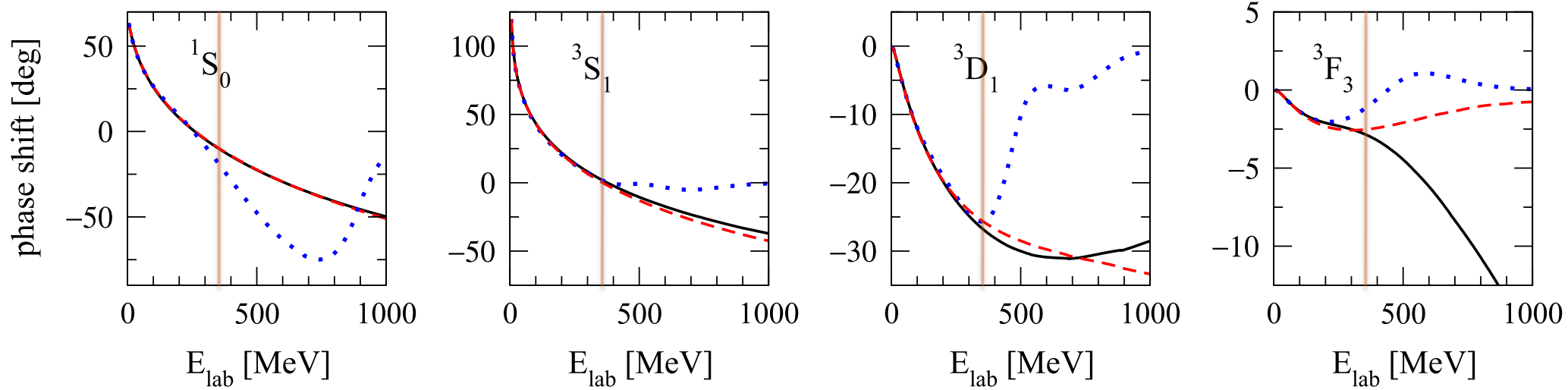
Meson-Exchange Potentials and Phase Shifts

Further model dependence: scattering phase shifts of NN potentials

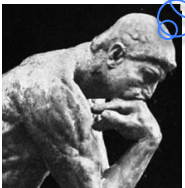


Meson-Exchange Potentials and Phase Shifts

Further model dependence: scattering phase shifts of NN potentials



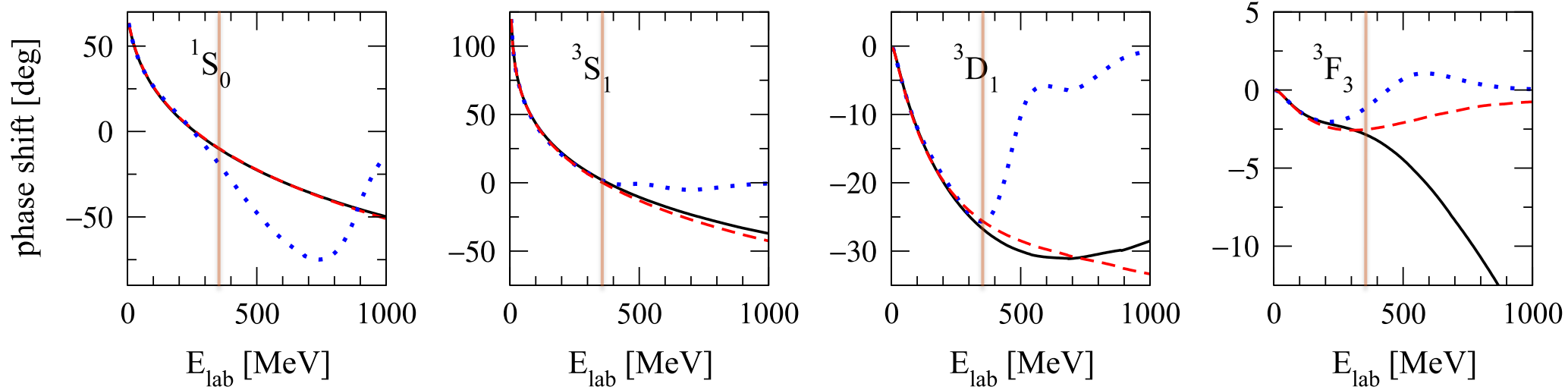
That's strange...
why do they only
agree to 350MeV?



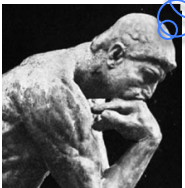
Remember, all have $\chi^2/\text{dof} \approx 1$

Meson-Exchange Potentials and Phase Shifts

Further model dependence: scattering phase shifts of NN potentials



That's strange...
why do they only
agree to 350MeV?

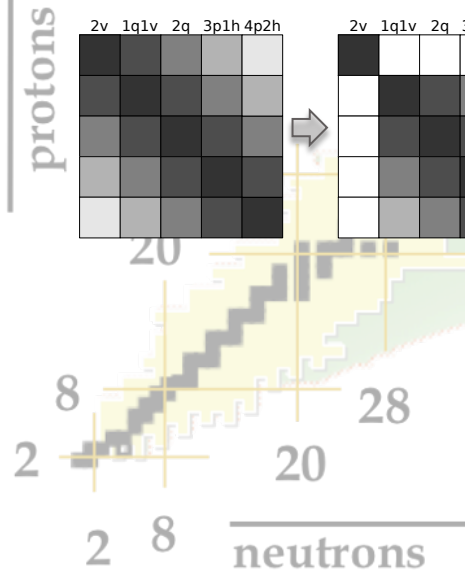
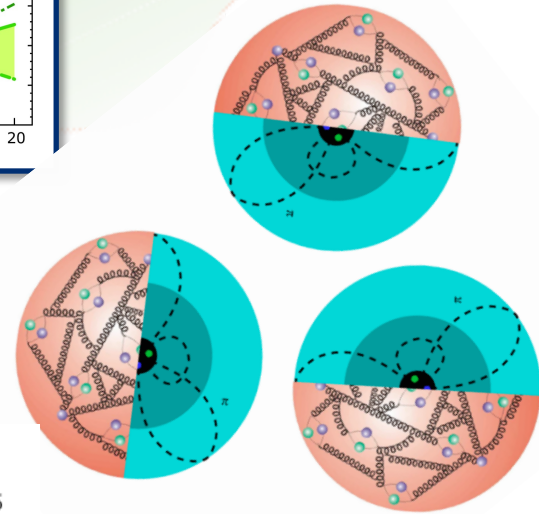
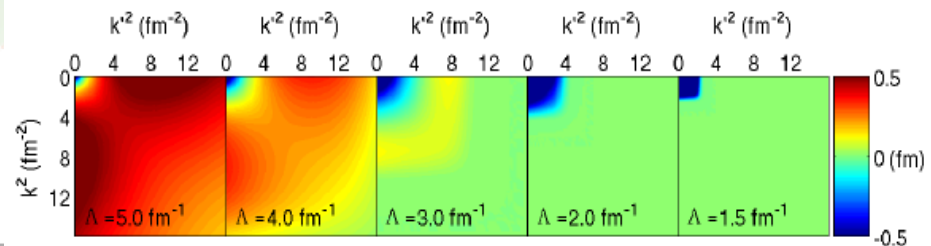
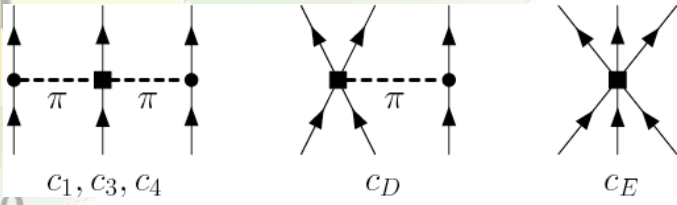
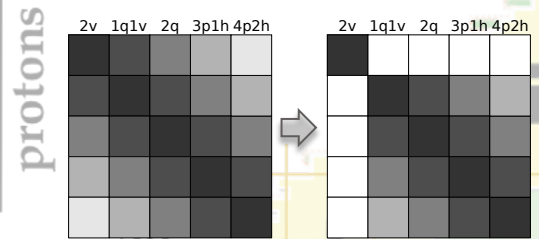
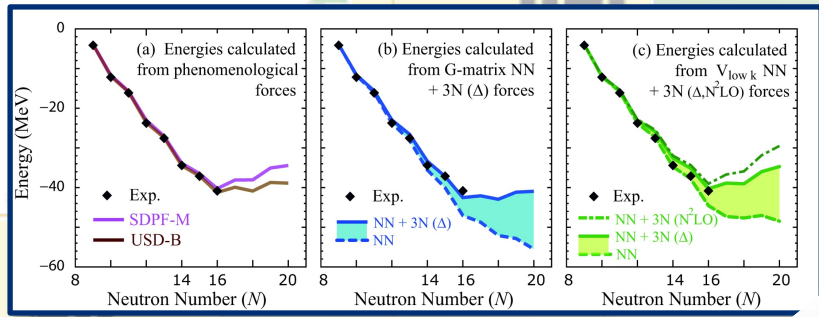
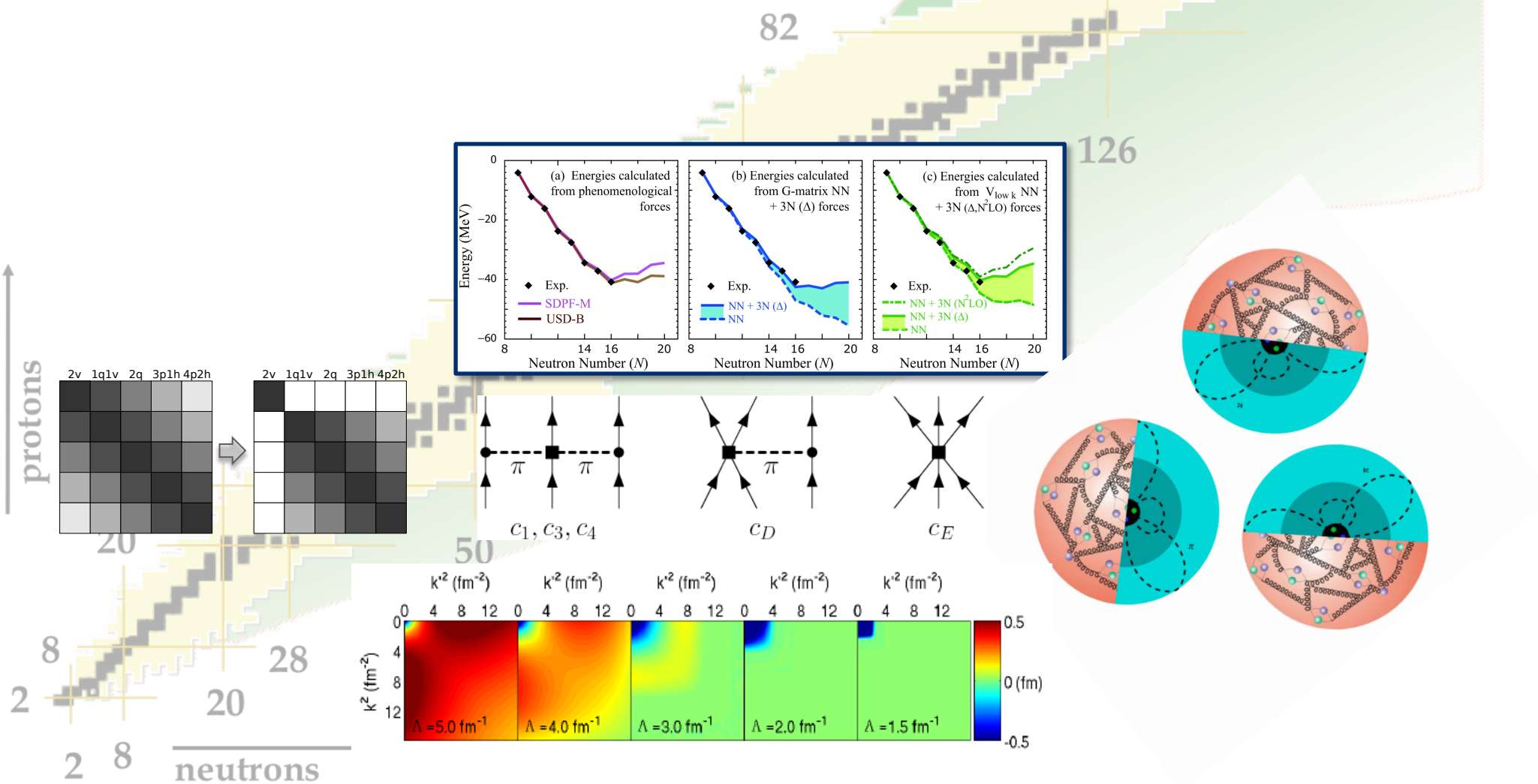


Agree well up to **pion-production threshold** $\sim 350\text{MeV}$

Most models don't fit phase shifts above this energy – **unconstrained**

Day 2: Effective Field Theories

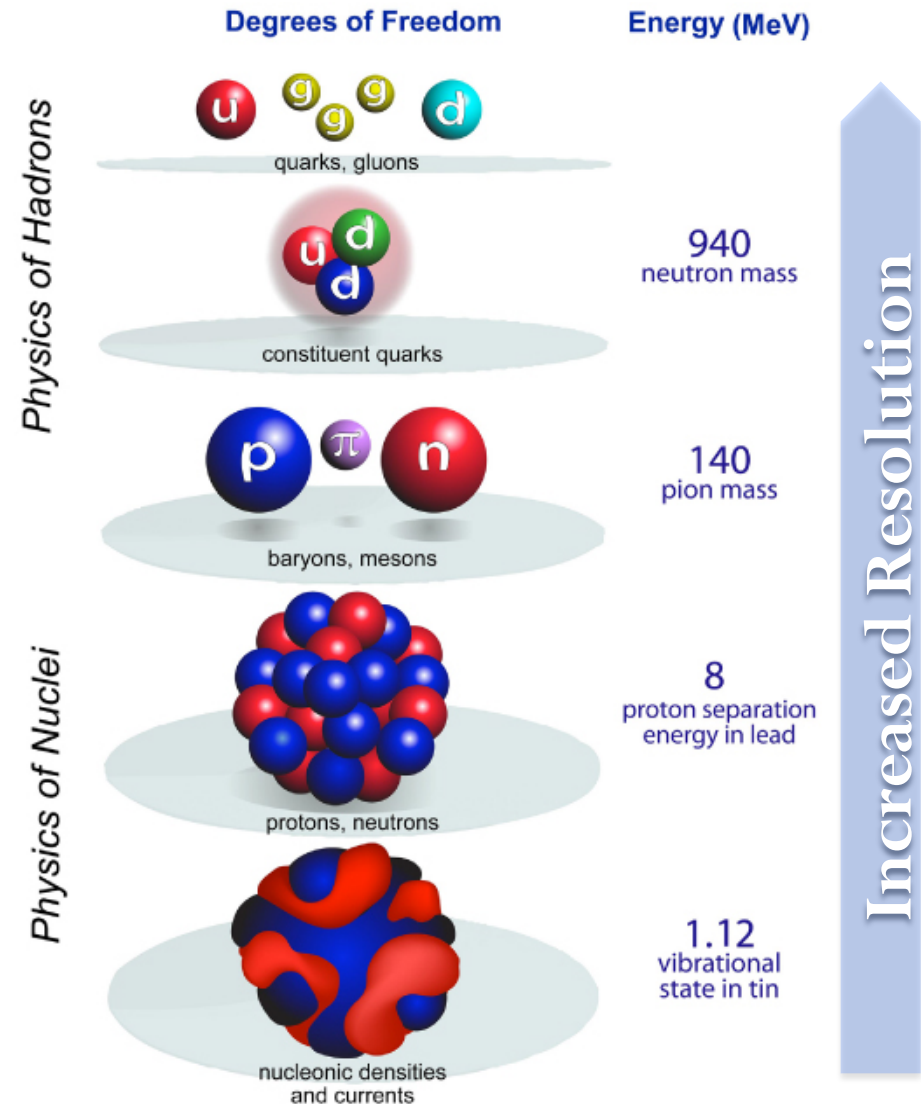
Jason D. Holt



From QCD to Nuclear Interactions

How do we determine interactions between nucleons?

$$H(\Lambda) = T + V_{NN}(\Lambda) + V_{3N}(\Lambda) + V_{4N}(\Lambda) + \dots$$



Old view:

Multiple scales complicate life

No meaningful way to connect them

Modern view:

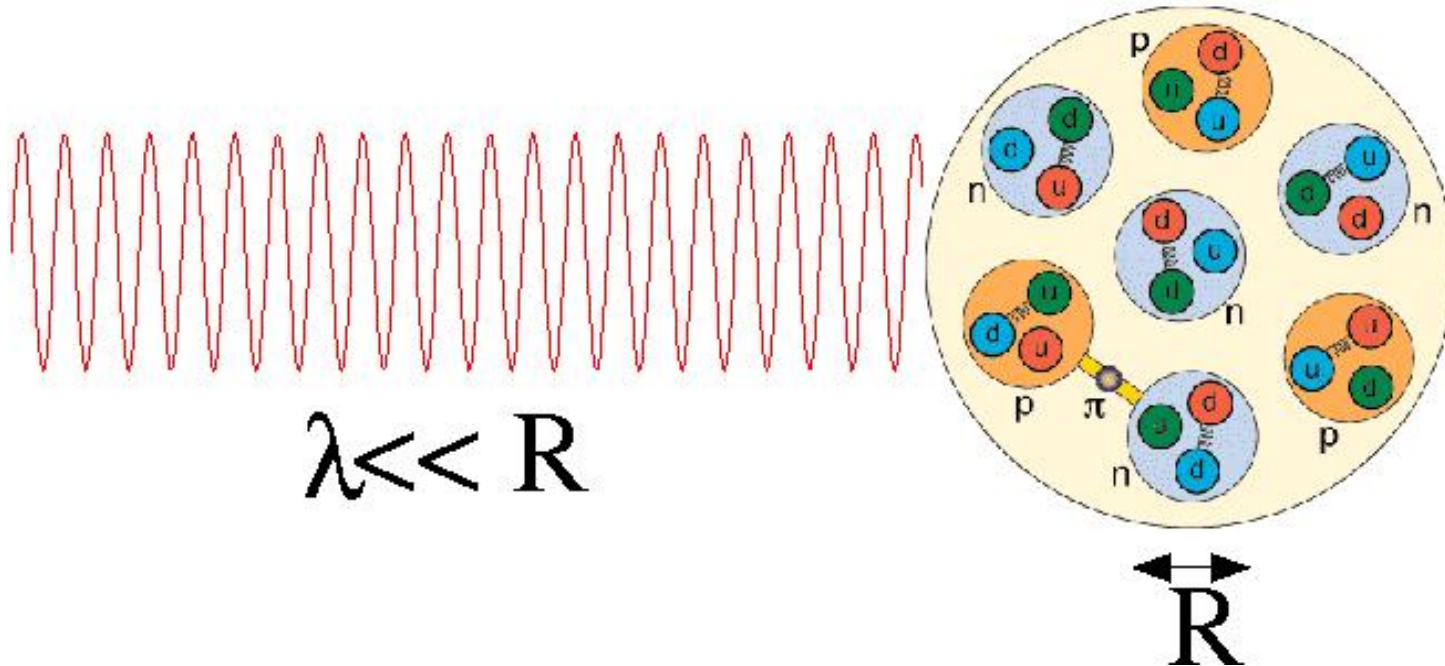
Choose convenient resolution scale

Effective field theory at each scale
connected by RG

Ratio of scales – small parameters

Ideas Behind Effective Theories

Resolution scale and relevant degrees of freedom

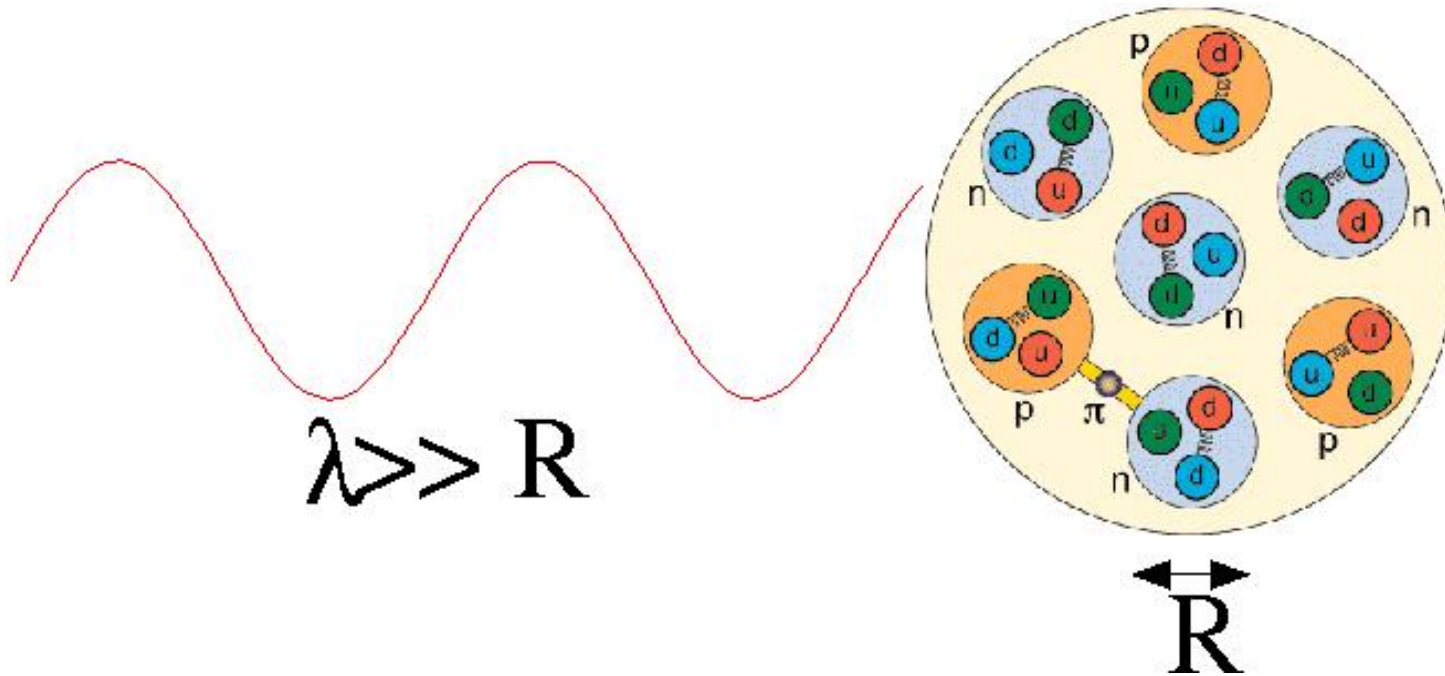


High energy probe resolves fine details

Need high-energy degrees of freedom

Ideas Behind Effective Theories

Resolution scale and relevant degrees of freedom

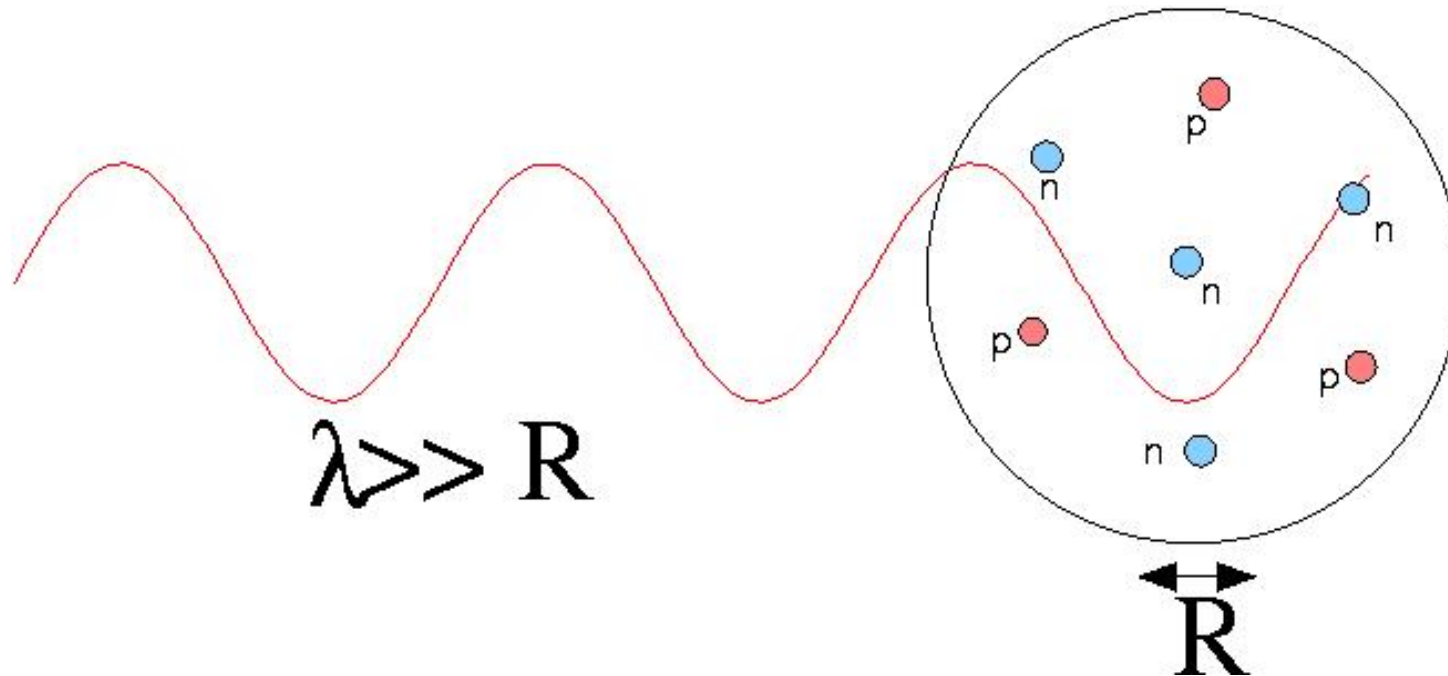


Low-energy probe can't resolve such details

Don't need high-energy degrees of freedom – replace with something simpler

Ideas Behind Effective Theories

Resolution scale and relevant degrees of freedom



Low-energy probe can't resolve such details

Don't need high-energy degrees of freedom – replace with something simpler

Use more convenient dofs, but **preserve low-energy observables!**

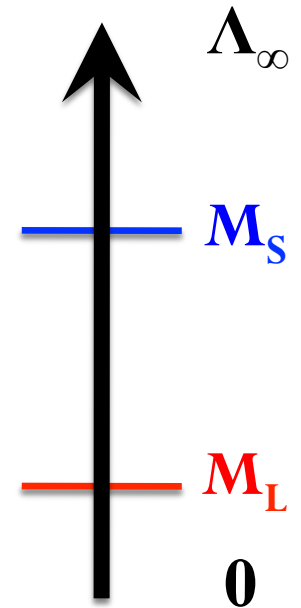
Ideas Behind Effective Theories

Assume underlying theory with cutoff Λ_∞

$$V = V_L + V_S$$

Known **long-distance physics** (like 1π exchange) with some scale M_L

Short-distance physics (ρ, ω exchange) with some scale M_S



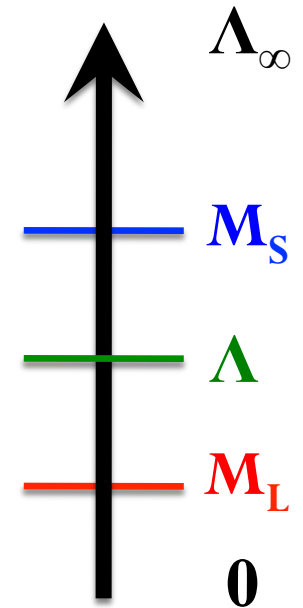
Ideas Behind Effective Theories

Assume underlying theory with cutoff Λ_∞

$$V = V_L + V_S$$

Known **long-distance physics** (like 1π exchange) with some scale M_L

Short-distance physics (ρ, ω exchange) with some scale M_S



And we want a **low-energy effective theory** for physics up to some

$$M_L < \Lambda < M_S$$

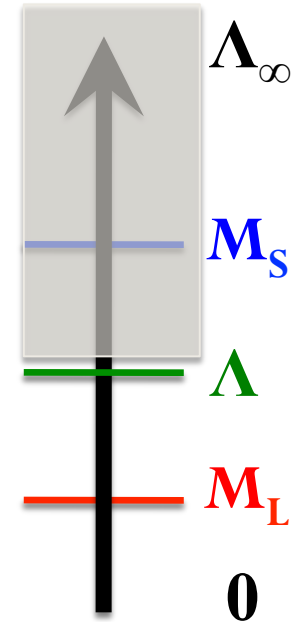
Ideas Behind Effective Theories

Assume underlying theory with cutoff Λ_∞

$$V = V_L + V_S$$

Known **long-distance physics** (like 1π exchange) with some scale M_L

Short-distance physics (ρ, ω exchange) with some scale M_S



And we want a **low-energy effective theory** for physics up to some

$$M_L < \Lambda < M_S$$

Integrate out states above Λ using **Renormalization Group (RG)**

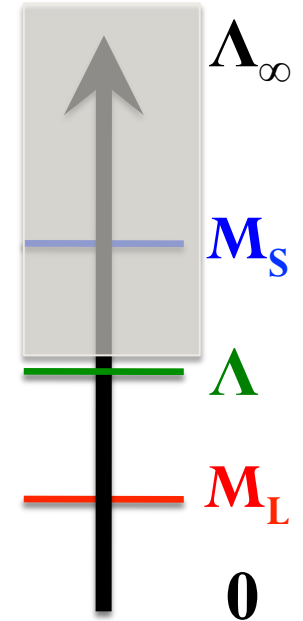
Ideas Behind Effective Theories

Assume underlying theory with cutoff Λ_∞

$$V = V_L + V_S$$

Known **long-distance physics** (like 1π exchange) with some scale M_L

Short-distance physics (ρ, ω exchange) with some scale M_S



And we want a **low-energy effective theory** for physics up to some

$$M_L < \Lambda < M_S$$

Integrate out states above Λ using **Renormalization Group (RG)**

General form of effective theory: $V_{\text{eff}} = V_L + \delta V_{\text{c.t.}}(\Lambda)$

where $\delta V_{\text{c.t.}}(\Lambda) = C_0(\Lambda)\delta^3(\mathbf{r}) + C_2(\Lambda)\nabla^2\delta^3(\mathbf{r}) + \dots$

Also use RG to change resolution scales **within particular EFT**

Ideas Behind Effective Theories

General form of effective theory: $V_{\text{eff}} = V_L + \delta V_{\text{c.t.}}(\Lambda)$

$$\delta V_{\text{c.t.}}(\Lambda) = C_0(\Lambda) \delta^3(\mathbf{r}) + C_2(\Lambda) \nabla^2 \delta^3(\mathbf{r}) + \dots$$

Encodes effects of high-E
dof on low-energy observables

Universal; depends only
on symmetries

TWO choices:

Short distance structure of “true theory” captured in several numbers

- **Calculate from underlying theory**

When short-range physics is unknown or too complicated

- **Extract from low-energy data**

How do we apply these ideas to nuclear forces?

Chiral Effective Field Theory: Philosophy

“At each scale we have different degrees of freedom and different dynamics. Physics at a larger scale (largely) decouples from physics at a smaller scale... thus a theory at a larger scale remembers only finitely many parameters from the theories at smaller scales, and throws the rest of the details away.

More precisely, when we pass from a smaller scale to a larger scale, we average out irrelevant degrees of freedom... The general aim of the RG method is to explain how this decoupling takes place and why exactly information is transmitted from one scale to another through finitely many parameters.”

- *David Gross*

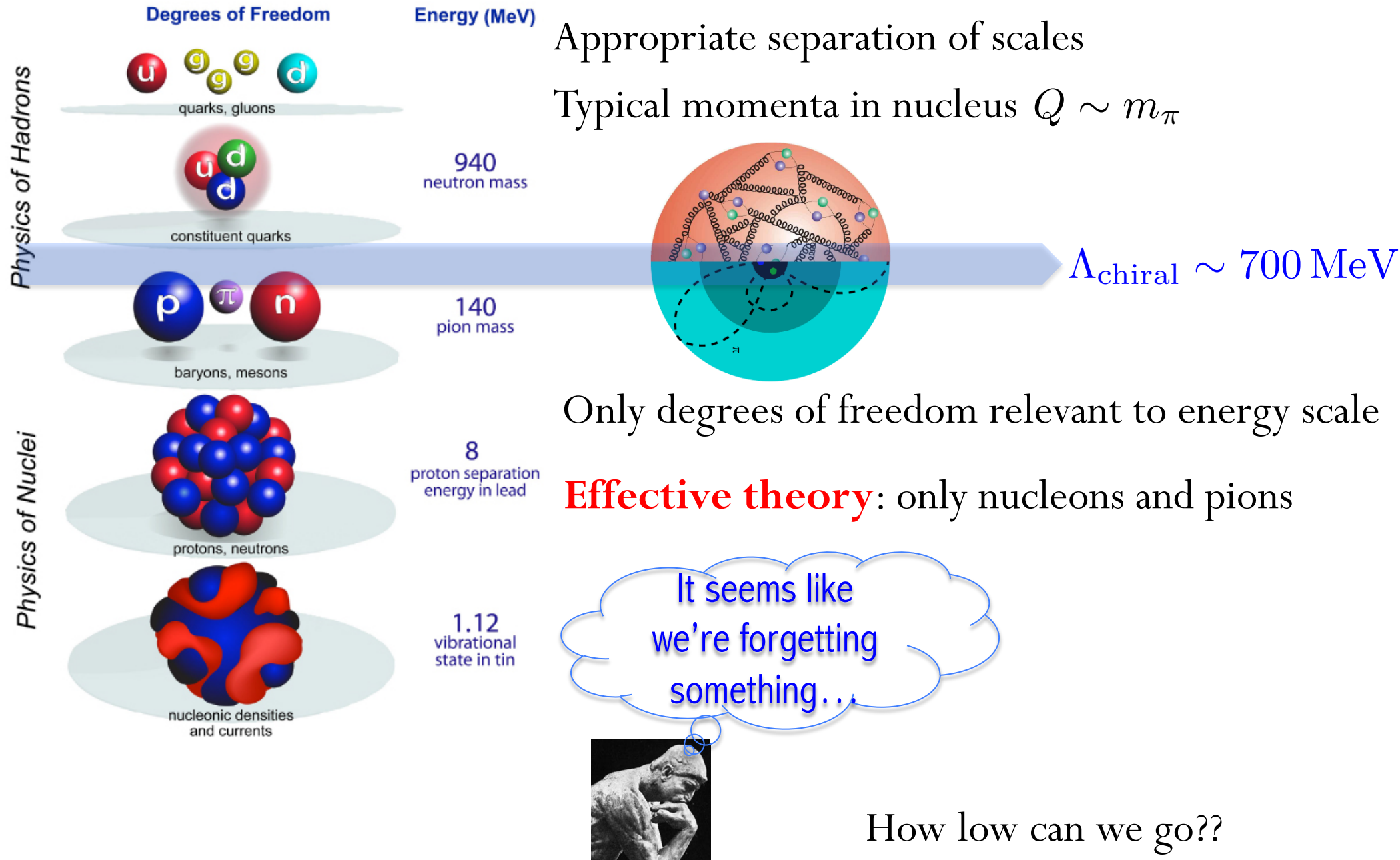
“The method in its most general form can.. be understood as a way to arrange in various theories that the degrees of freedom that you’re talking about are the relevant degrees of freedom for the problem at hand.”

- *Steven Weinberg*

5 Steps to constructing such a theory for nuclear forces

Separation of Scales in Nuclear Physics

Step I: Identify appropriate separation of scales, degrees of freedom



Chiral EFT Symmetries

Step II: Identify relevant symmetries of underlying theory (QCD)

1. SU(3) color symmetry from QCD
(Nucleons and pions are color singlets)
2. Chiral symmetry: u and d quarks are almost massless
 - Left and right-handed (massless) quarks do not mix: $SU(2)_L \times SU(2)_R$ symmetry
 - Explicit symmetry breaking: u and d quarks have a small mass
 - Spontaneous breaking of chiral symmetry (no parity doublets observed in Nature)
 - $SU(2)_L \times SU(2)_R$ symmetry spontaneously broken to $SU(2)_V$
 - Pions are the Nambu-Goldstone bosons of spontaneously broken symmetry
 - Low-energy pion Lagrangian completely determined

Missing ingredient in multi-pion-exchange theories of 50's!

Construct Lagrangian based on these symmetries

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\pi\pi} + \mathcal{L}_{\pi N} + \mathcal{L}_{NN}$$

Chiral EFT Lagrangian

Step III: Construct Lagrangian **based on identified symmetries**

Pion-pion Lagrangian: U is $SU(2)$ matrix parameterized by three pion fields

$$\mathcal{L}_{\pi}^{(0)} = \frac{F^2}{4} \langle \nabla^\mu U \nabla_\mu U^\dagger + \chi_+ \rangle,$$

Leading-order pion-nucleon

$$\mathcal{L}_{\pi N}^{(0)} = \bar{N}(i v \cdot D + \hat{g}_A u \cdot S)N,$$

Leading-order nucleon-nucleon (encodes unknown short-range physics)

$$\mathcal{L}_{NN}^{(0)} = -\frac{1}{2} C_S (\bar{N}N)(\bar{N}N) + 2C_T (\bar{N}S N) \cdot (\bar{N}S N)$$

EFT Power Counting

Step IV: Design an **organized** scheme to distinguish more from less important processes: **Power Counting**

Organize theory in powers of $\left(\frac{Q}{\Lambda_\chi}\right)$ where $Q \sim m_\pi$ typical nuclear momenta

Only valid for small expansion parameters, *i.e.*, low momentum

Irreducible time-ordered diagram has order: $\left(\frac{Q}{\Lambda_\chi}\right)^\nu$

$$\nu = -4 + 2N + 2L + \sum_i V_i \Delta_i \quad \Delta_i = d_i + \frac{1}{2}n_i - 2 \text{ "Chiral dimension"}$$

N = Number of nucleons

L = Number of pion loops

V_i = Number of vertices of type i

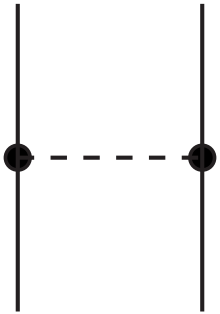
d = Number of derivatives or insertions of m_π

n = Number of nucleon field operators

Chiral EFT: Lowest Order (LO)

Step V: Calculate Feynmann diagrams to the desired accuracy

Leading order (LO) $\nu = 0$



One-pion exchange

$$V_{NN}^{(0)} = - \frac{g_A^2}{4F_\pi^2} \frac{\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q}}{\vec{q}^2 + M_\pi^2} \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2$$

$$g_A = 1.26$$

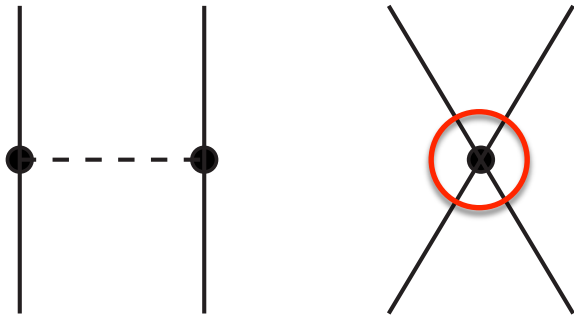
$$F_\pi = 92.4 \text{ MeV}$$

$$\vec{q}_i \equiv \vec{p}'_i - \vec{p}_i \quad \vec{k}_i \equiv \frac{1}{2} (\vec{p}'_i + \vec{p}_i)$$

Chiral EFT: Lowest Order (LO)

Step V: Calculate Feynmann diagrams to the desired accuracy

Leading order (LO) $\nu = 0$



One-pion exchange
NN contact interaction

$$V_{NN}^{(0)} = -\frac{g_A^2}{4F_\pi^2} \frac{\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q}}{\vec{q}^2 + M_\pi^2} \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2 + \underbrace{C_S}_{\text{red circle}} + \underbrace{C_T}_{\text{red circle}} \vec{\sigma}_1 \cdot \vec{\sigma}_2$$

$$g_A = 1.26$$

Two **low-energy constants (LECs)**: C_S, C_T

$$F_\pi = 92.4 \text{ MeV}$$

$$\vec{q}_i \equiv \vec{p}'_i - \vec{p}_i \quad \vec{k}_i \equiv \frac{1}{2} (\vec{p}'_i + \vec{p}_i)$$

Chiral EFT

Step V: Calculate Feynmann diagrams to the desired accuracy

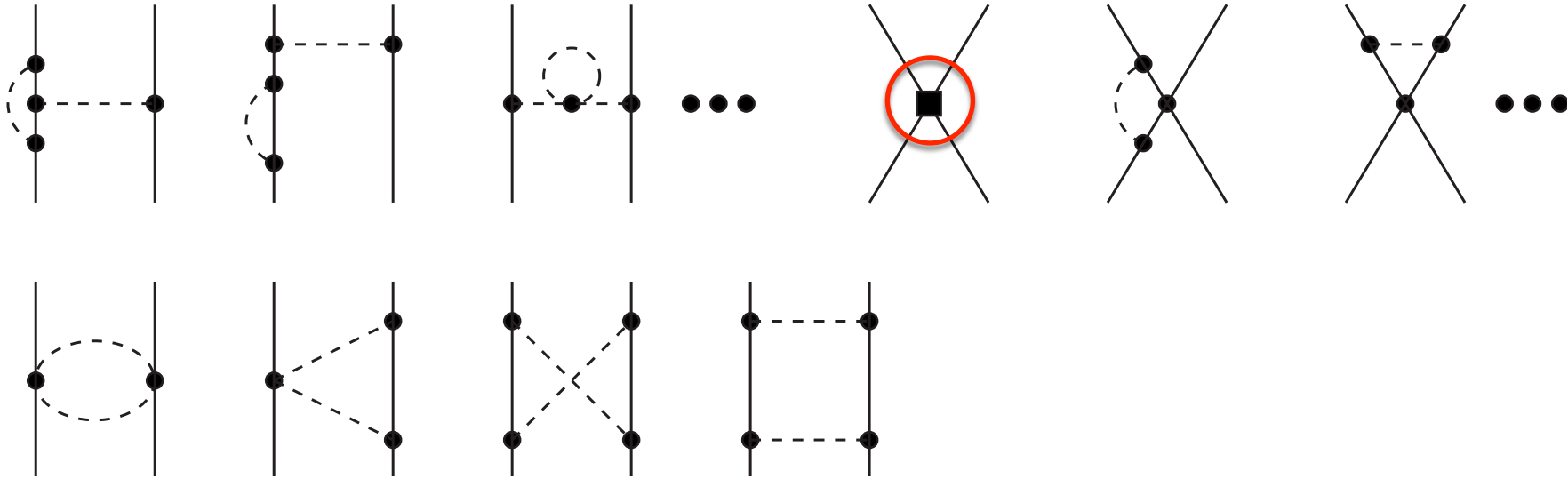
Question: What will $\mathcal{V} = 1$ look like?

Answer: No contribution at this order

Chiral EFT: NLO

Step V: Calculate Feynmann diagrams to the desired accuracy

Next-to-leading order (NLO) $\nu = 2$



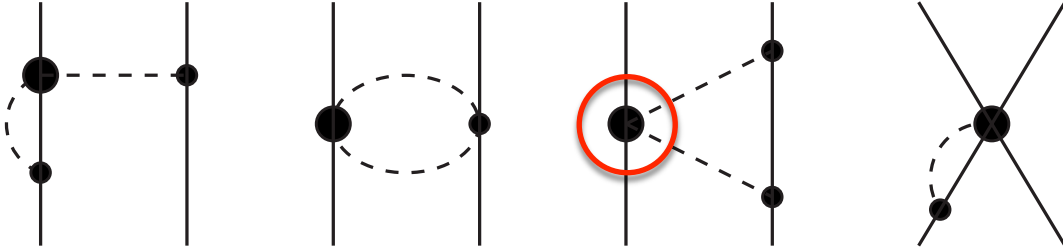
Higher order contact interaction: 7 new LECs, spin-orbit

$$\begin{aligned}
 &+ \textcircled{C_1} \vec{q}^2 + \textcircled{C_2} \vec{k}^2 + (\textcircled{C_3} \vec{q}^2 + \textcircled{C_4} \vec{k}^2) \vec{\sigma}_1 \cdot \vec{\sigma}_2 \\
 &+ i \textcircled{C_5} \frac{1}{2} (\vec{\sigma}_1 + \vec{\sigma}_2) \cdot \vec{q} \times \vec{k} + \textcircled{C_6} \vec{q} \cdot \vec{\sigma}_1 \vec{q} \cdot \vec{\sigma}_2 \\
 &+ \textcircled{C_7} \vec{k} \cdot \vec{\sigma}_1 \vec{k} \cdot \vec{\sigma}_2,
 \end{aligned}$$

Chiral EFT: N²LO

Step V: Calculate Feynmann diagrams to the desired accuracy

Next-to-next-to-leading order (N²LO) $\nu = 3$



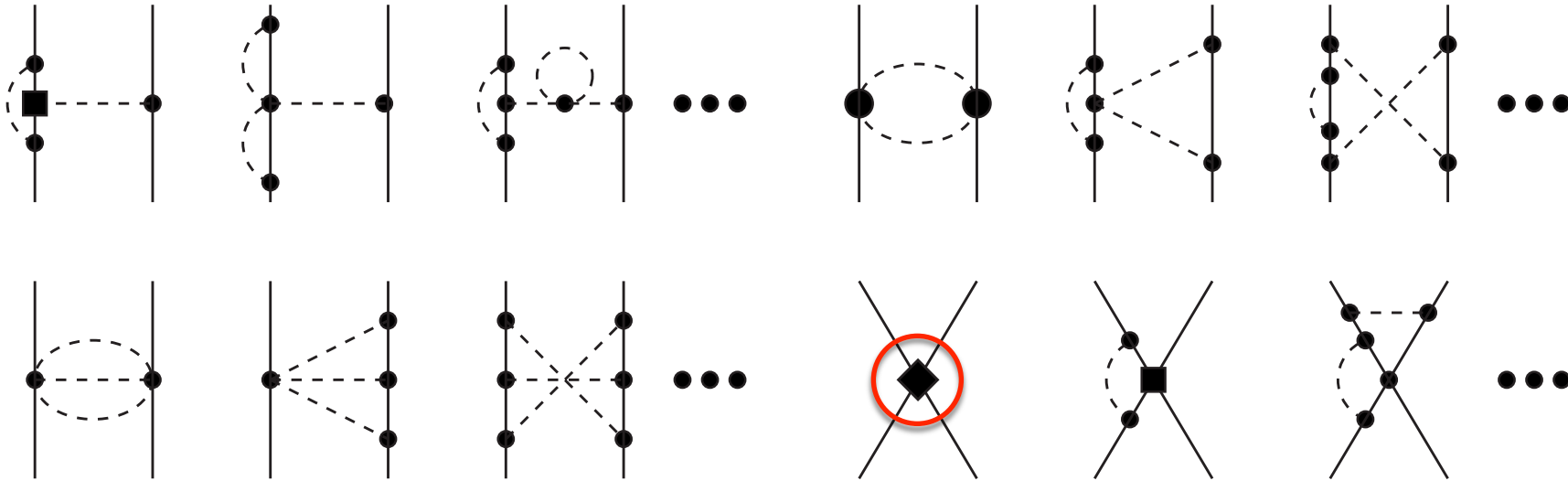
3 new couplings from $\pi\pi NN$ vertex – not LECs!

$$\begin{aligned}
 V_{NN}^{(3)} = & -\frac{3g_A^2}{16\pi F_\pi^4} [2M_\pi^2 (2c_1 - c_3) - c_3 \vec{q}^2] \\
 & \times (2M_\pi^2 + \vec{q}^2) A^{\tilde{\Lambda}}(q) - \frac{g_A^2 c_4}{32\pi F_\pi^4} \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2 (4M_\pi^2 \\
 & + q^2) A^{\tilde{\Lambda}}(q) (\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q} - \vec{q}^2 \vec{\sigma}_1 \cdot \vec{\sigma}_2),
 \end{aligned}$$

Chiral EFT: $N^3\text{LO}$

Step V: Calculate Feynmann diagrams to the desired accuracy

Next-to-next-to-next-to-leading order $\nu = 4$



Higher order contact interaction: 15 new LECs

Regularization of Chiral potentials

Remember: constructing potential involves solving **L-S equation**

All NN potentials cutoff loop momenta at some value $> 1\text{ GeV}$

Impose exponential regulator, Λ , in Chiral EFT potentials – **not in integral**

$$T^\alpha(k, k') = V^\alpha(k, k') + \frac{2}{\pi} \sum_{l''} \int_0^\infty q^2 dq V^\alpha(k, q) \frac{q}{k^2 - q^2 + i\varepsilon} T^\alpha(q, k')$$

$$V(k, k') \rightarrow e^{(-k'/\Lambda)^{2n}} V(k, k') e^{(-k/\Lambda)^{2n}}$$

LECs will depend on regularization approach and Λ

Infinitely many ways to do this

\implies **Infinitely many chiral potentials!**

Indeed, many on the market – some fit well to phase shifts, others not

Chiral EFT: Resulting fits to Phase shifts

Systematic improvement of chiral EFT potentials fit to phase shifts

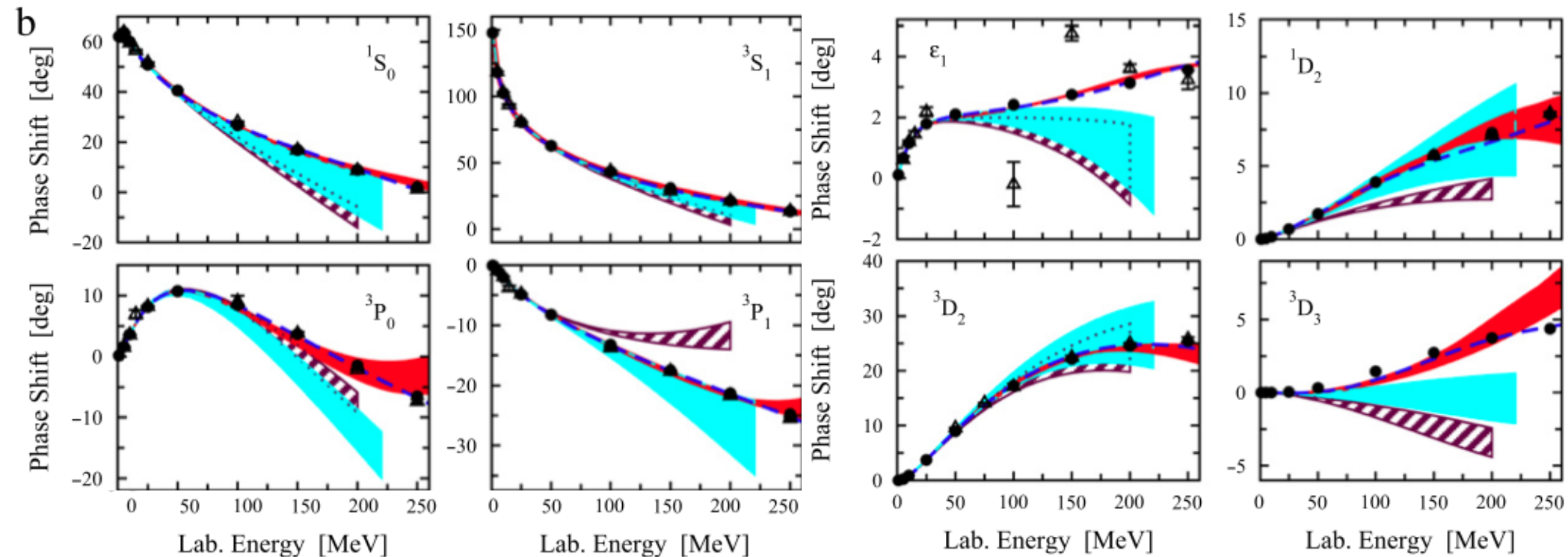
Cutoff variation – information about missing physics

NLO: dashed band  9 Parameters

N²LO: light band  12 Parameters

N³LO: dark band  27 Parameters

Generally decreasing error and increasing accuracy – not entirely... (exercise)



Chiral Effective Field Theory: Nuclear Forces

	2N forces	3N forces	4N forces
LO			
NLO			
N ² LO			
N ³ LO			

Meson exchange potentials were an admirable effort

Using ideas of **effective field theory**:

Nucleons interact via pion exchanges and contact interactions

Lower momentum

Systematic – can assign error

Connected to QCD

Hierarchy: $V_{\text{NN}} > V_{\text{3N}} > \dots$

Consistent treatment of NN, 3N, ... electroweak operators

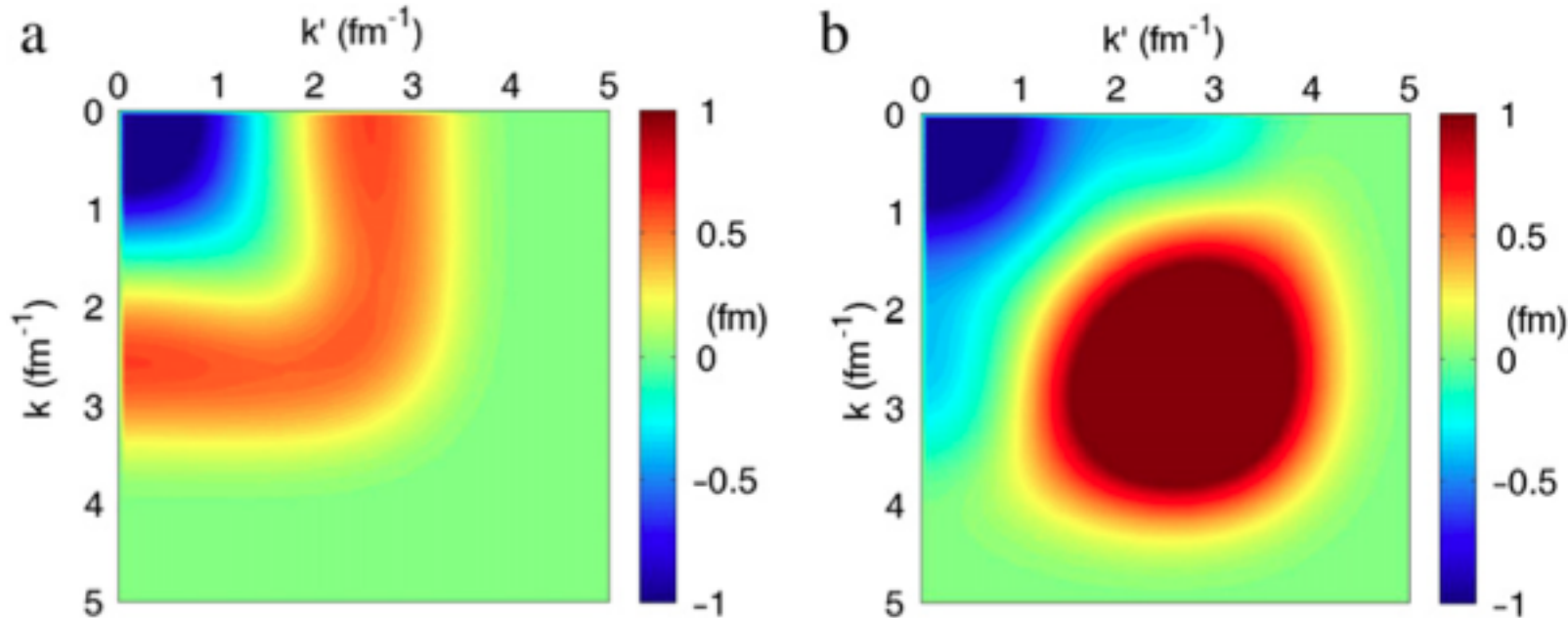
Couplings fit to experiment once

Weinberg, van Kolck, Kaplan, Savage, Wise, Epelbaum, Kaiser, Meissner,...

Chiral NN Potentials

Two chiral potentials with regulators of 500MeV and 600MeV

Still low-to-high momentum coupling: poor convergence, non perturbative, etc.



How do these compare to the potential you drew?

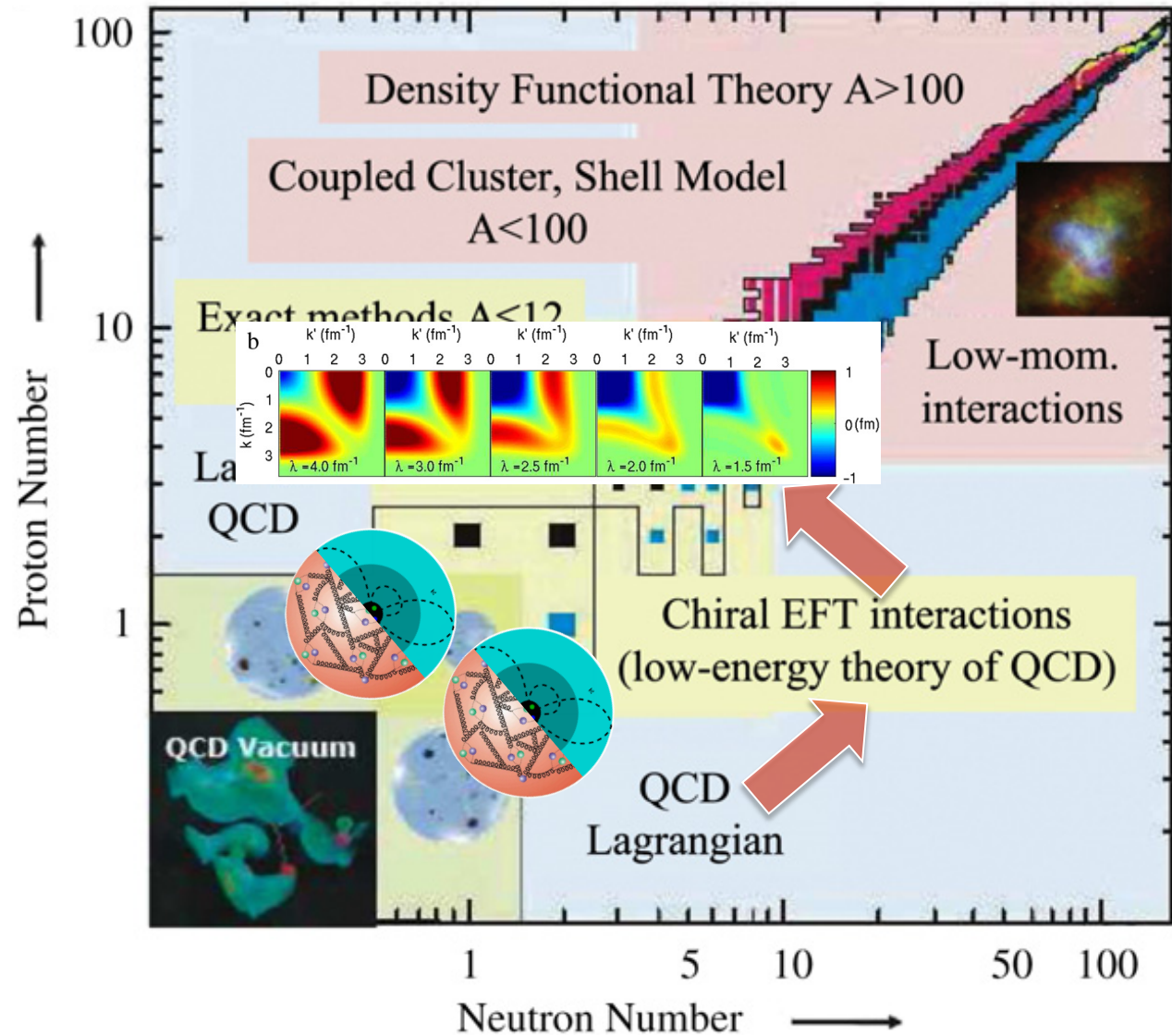
Lesson: Infinitely many phase-shift equivalent potentials

$$E_n = \langle \Psi_n | H | \Psi_n \rangle = \left(\langle \Psi_n | U^\dagger \right) U H U^\dagger (U | \Psi_n \rangle) = \left\langle \tilde{\Psi}_n \left| \tilde{H} \right| \tilde{\Psi}_n \right\rangle$$

NN interaction not observable Low-to-high momentum makes life difficult for low-energy nuclear theorists

Part II: (S)RG and Low-Momentum Interactions

To understand the properties of complex nuclei from first principles



Renormalizing NN Interactions

Basic ideas of RG

Low-momentum interactions

Similarity RG interactions

Benefits of low cutoffs

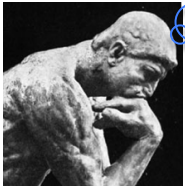
G-matrix renormalization

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow "Solve" many-body problem \rightarrow Predictions

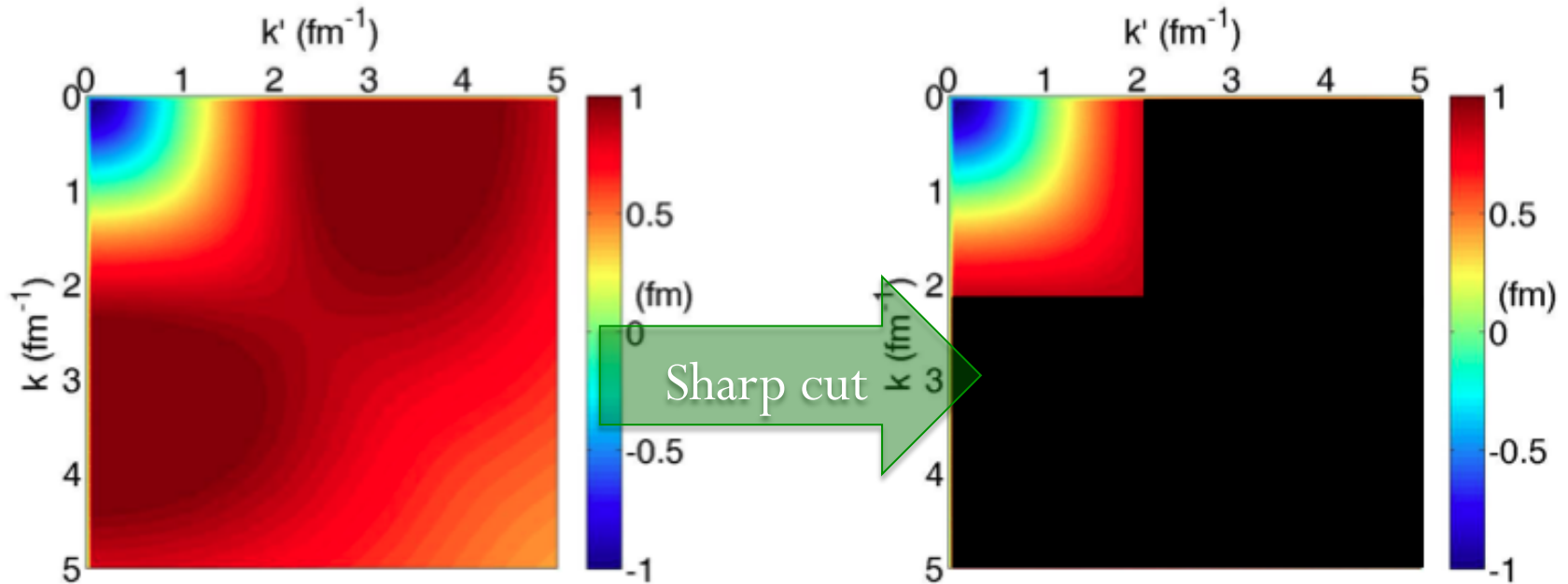
Renormalization of Meson-Exchange Potentials

Ok, high momentum is a pain. I wonder what would happen to low-energy observables...



Low-to-high momentum makes life difficult for low-energy nuclear theorists, so let's get rid of it

Can we just make a sharp cut and see if it works?



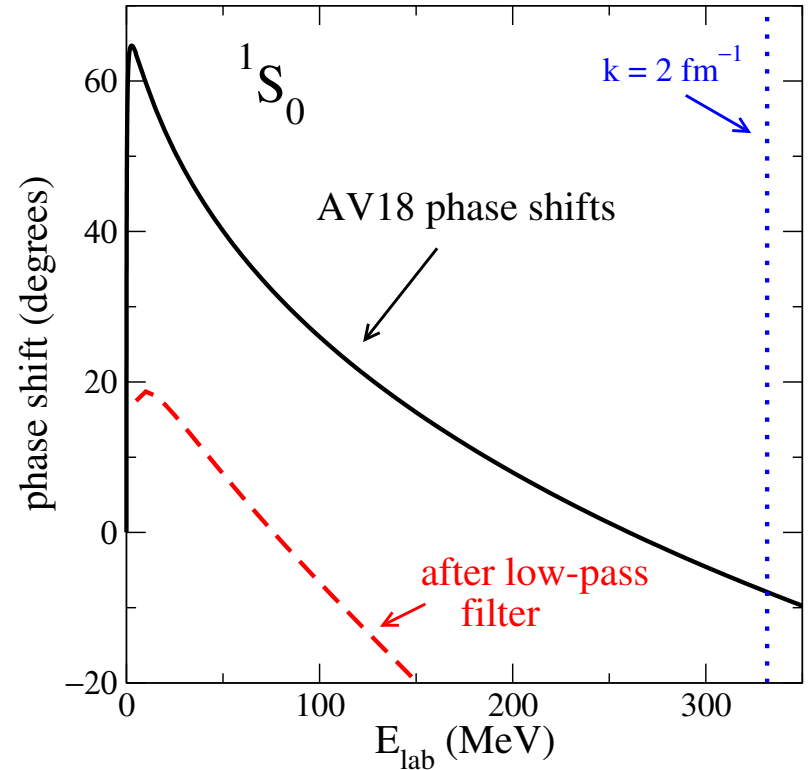
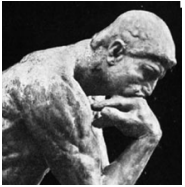
$$V_{\text{filter}}(k', k) \equiv 0; \quad k, k' > 2.2 \text{ MeV}$$

Renormalization of Meson-Exchange Potentials

Can we just make a sharp cut?

Nope! Low-energy physics is not correct

Glad I didn't bet money
on that... I wonder what
went wrong

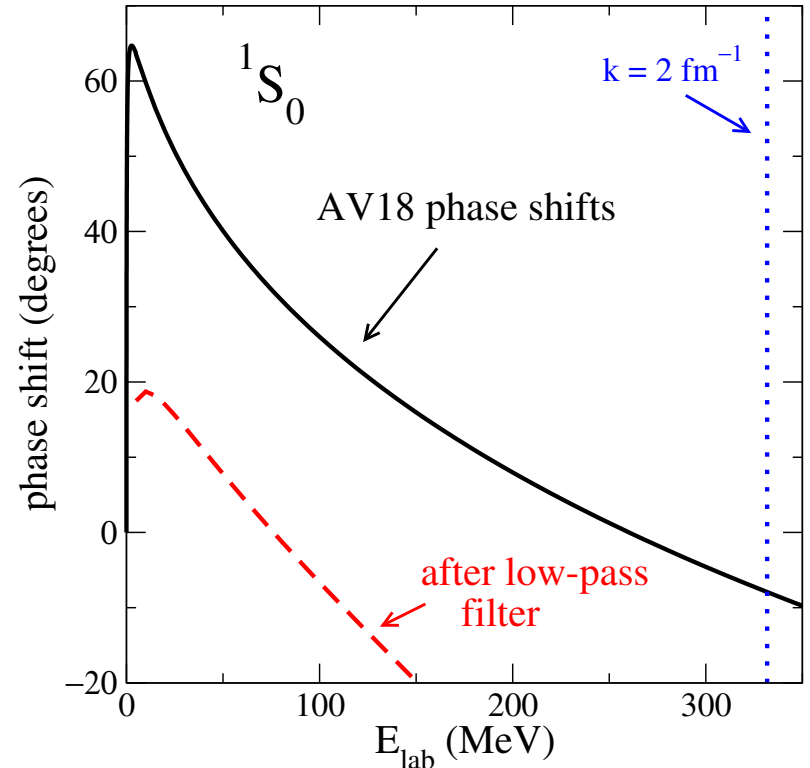
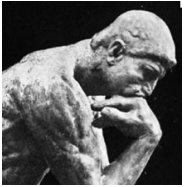


Renormalization of Meson-Exchange Potentials

Can we just make a sharp cut?

Nope! Low-energy physics is not correct

Glad I didn't bet money
on that... I wonder what
went wrong



Phase shifts involve couplings of low-to-high momenta

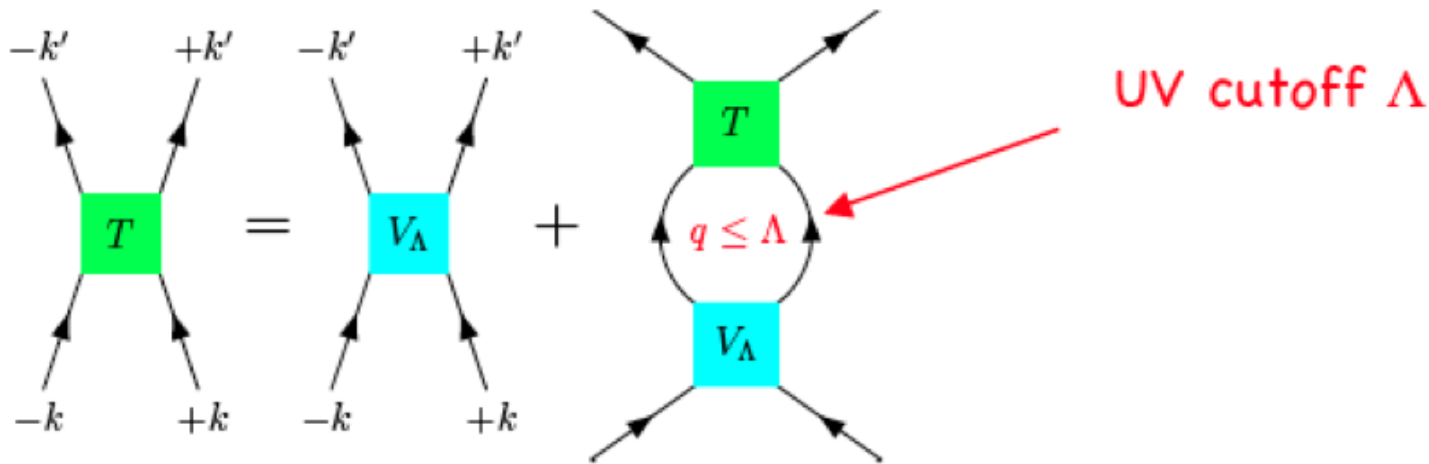
$$\langle k|V|k'\rangle + \sum_{q=0}^{\Lambda} \frac{\langle k|V|q\rangle\langle q|V|k'\rangle}{\epsilon_{k'} - \epsilon_q} + \sum_{q=\Lambda}^{\infty} \frac{\langle k|V|q\rangle\langle q|V|k'\rangle}{\epsilon_{k'} - \epsilon_q}$$

Lesson: Must ensure low-energy physics is preserved!

Renormalization of Meson-Exchange Potentials

To do properly, from T -matrix equation, define **low-momentum** equation:

$$T^\alpha(k, k') = V_{\text{NN}}^\alpha(k, k') + \frac{2}{\pi} \int_0^\infty q^2 dq \frac{V_{\text{NN}}^\alpha(k, q) T^\alpha(q, k')}{k^2 - q^2 + i\varepsilon}$$
$$\rightarrow V_{\text{low } k}^\Lambda(k, k') + \frac{2}{\pi} \int_0^\Lambda q^2 dq \frac{V_{\text{low } k}^\Lambda(k, q) T(q, k')}{k^2 - q^2 + i\varepsilon}$$



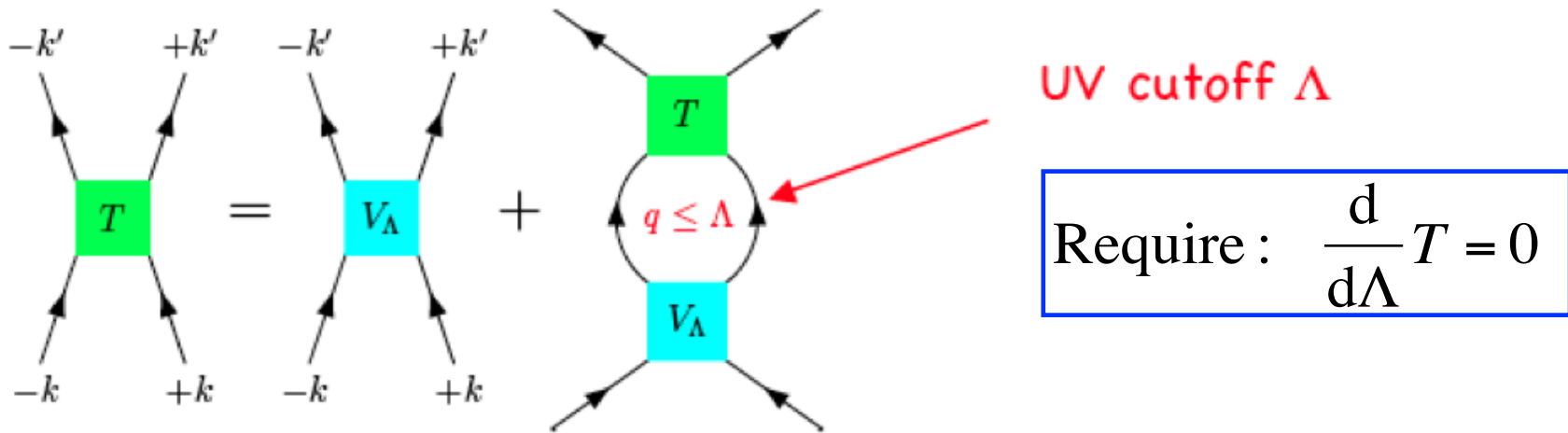
Lower UV cutoff, but preserve low-energy physics!

Renormalization of Meson-Exchange Potentials

To do properly, from T -matrix equation, define **low-momentum** equation:

$$T^\alpha(k, k') = V_{\text{NN}}^\alpha(k, k') + \frac{2}{\pi} \int_0^\infty q^2 dq \frac{V_{\text{NN}}^\alpha(k, q) T^\alpha(q, k')}{k^2 - q^2 + i\varepsilon}$$

$$\rightarrow V_{\text{low } k}^\Lambda(k, k') + \frac{2}{\pi} \int_0^\Lambda q^2 dq \frac{V_{\text{low } k}^\Lambda(k, q) T(q, k')}{k^2 - q^2 + i\varepsilon}$$



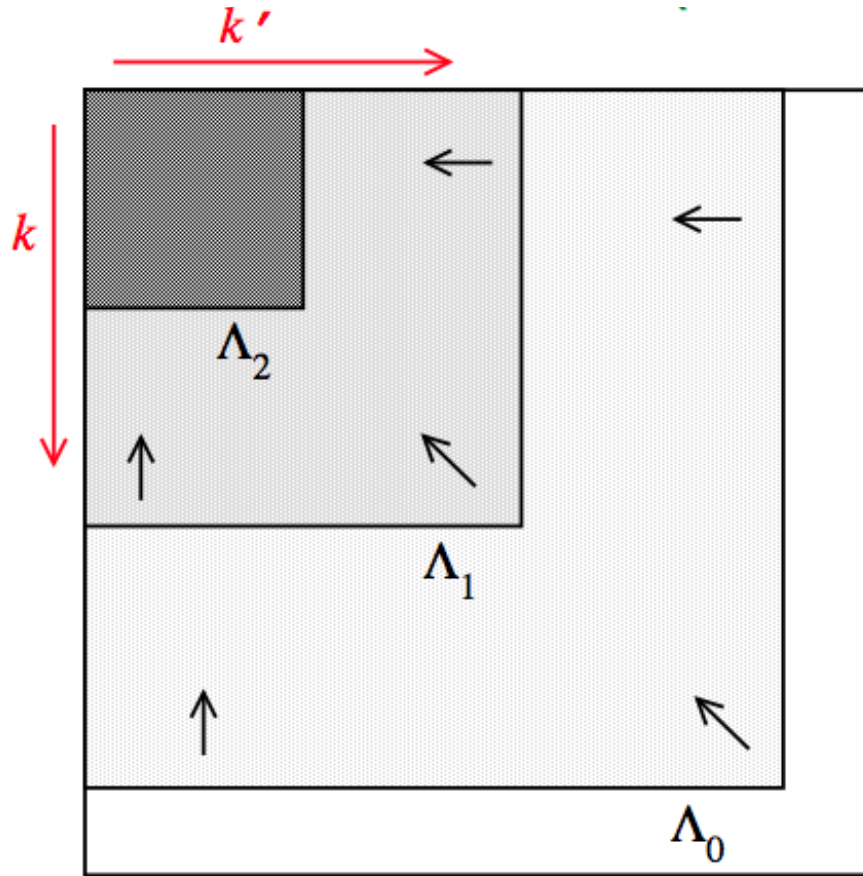
Lower UV cutoff, but preserve low-energy physics!

Leads to **renormalization group equation** for low-momentum interactions

$$\frac{d}{d\Lambda} V_{\text{low } k}^\Lambda(k', k) = \frac{2}{\pi} \frac{V_{\text{low } k}^\Lambda(k', \Lambda) T^\Lambda(\Lambda, k)}{1 - (k/\Lambda)^2}$$

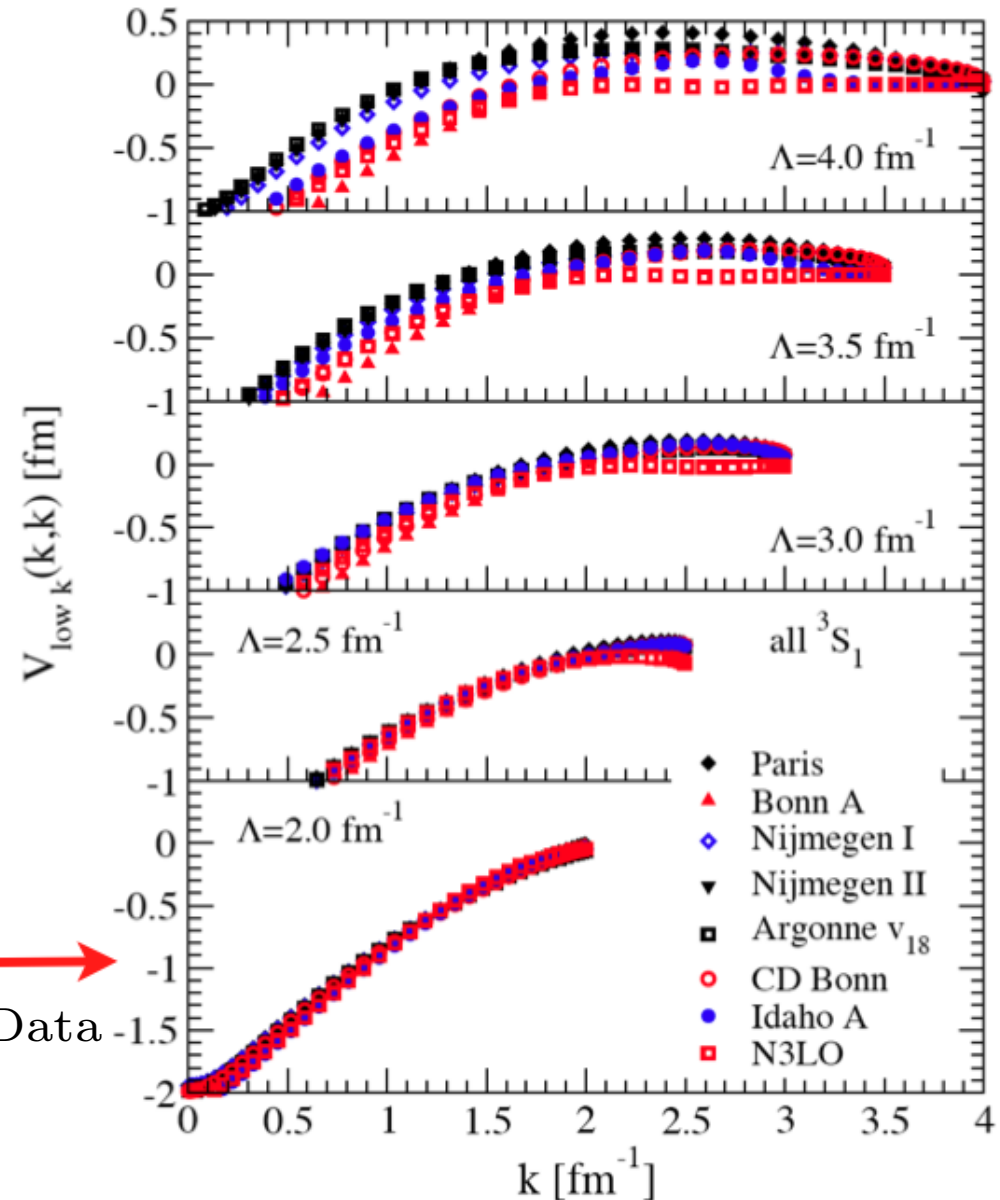
Renormalization of Meson-Exchange Potentials

Run cutoff to lower values – **decouples** high-momentum modes



Start from some initial V_{NN}
at high cutoff Λ_0

$$\Lambda \approx \Lambda_{\text{Data}} \rightarrow$$

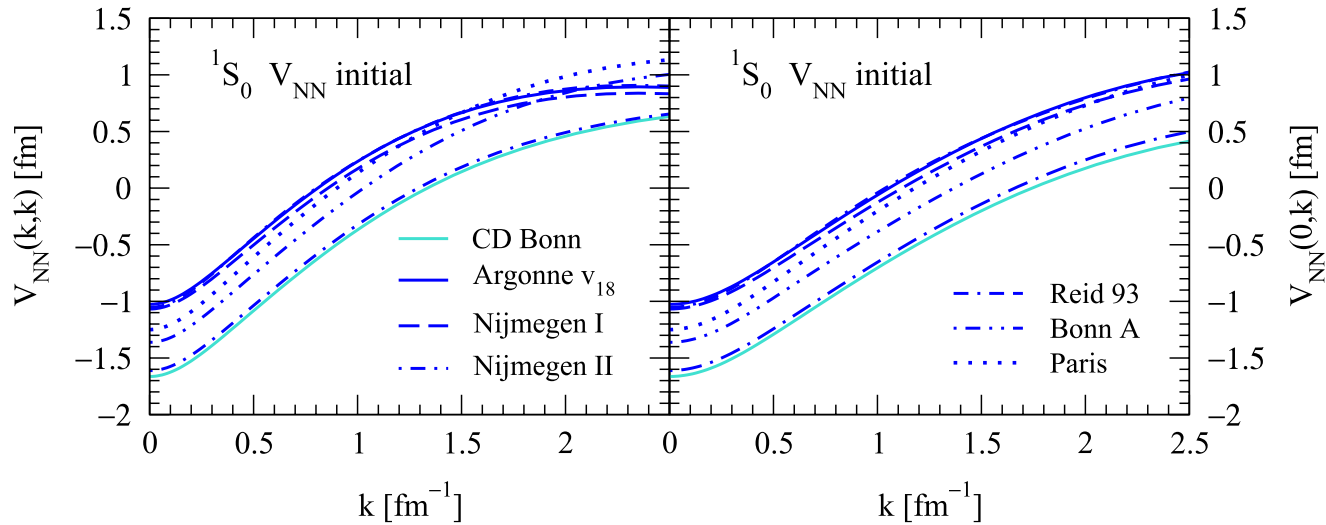


“Universality” at low momentum

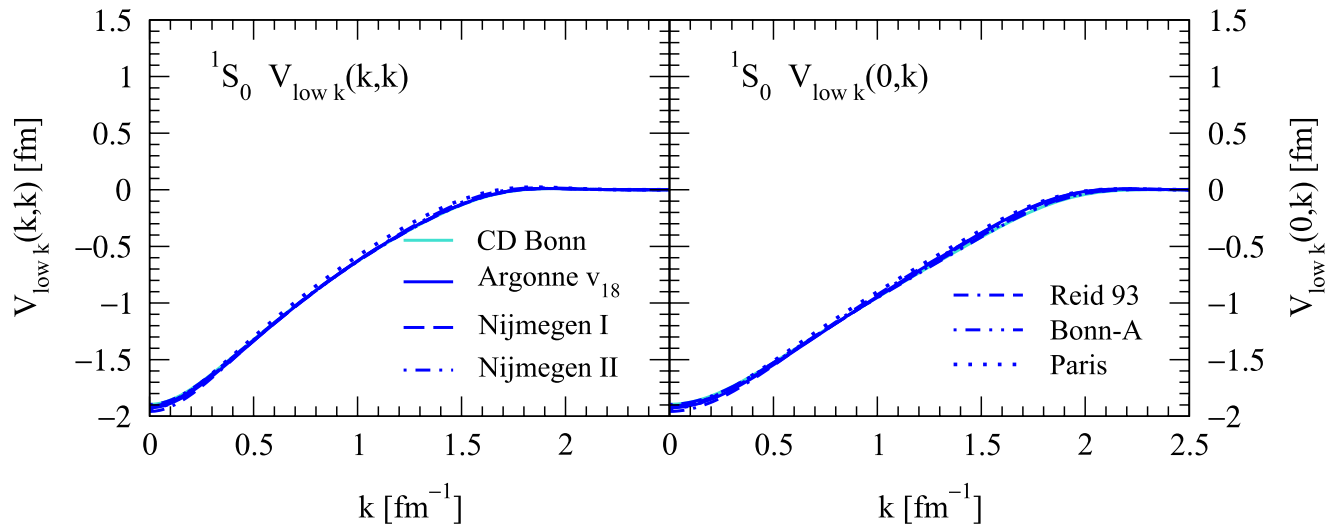
Renormalization of Meson-Exchange Potentials

Diagonal

Off-diagonal



These are all our favorite OBE NN potentials...



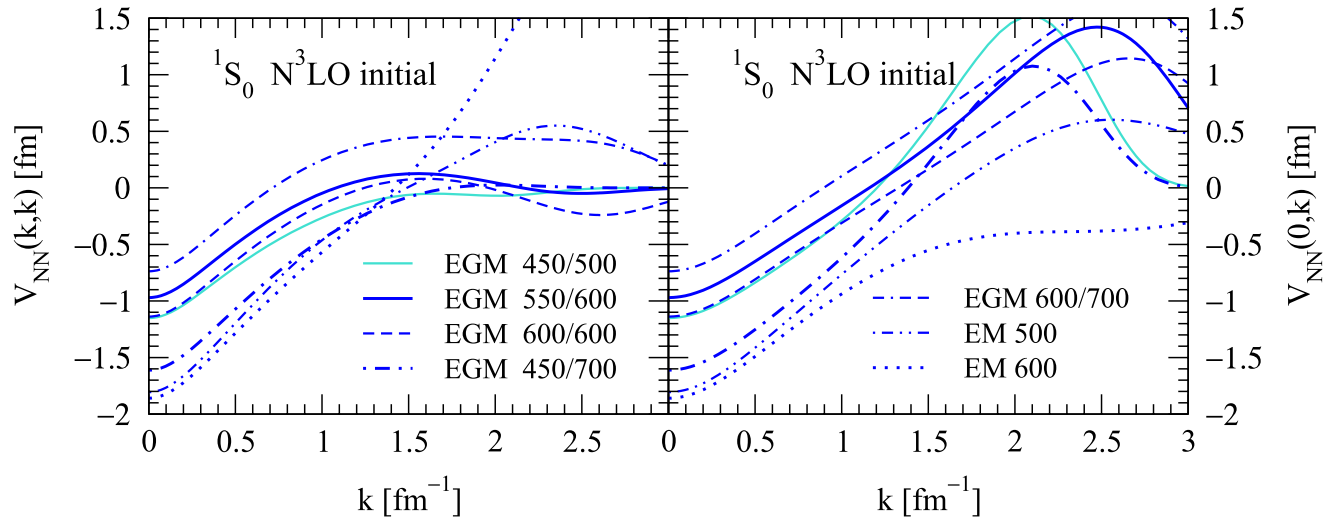
These are all our favorite OBE NN potentials...
at low momentum

Universal collapse in both diagonal/off-diagonal components, most partial waves

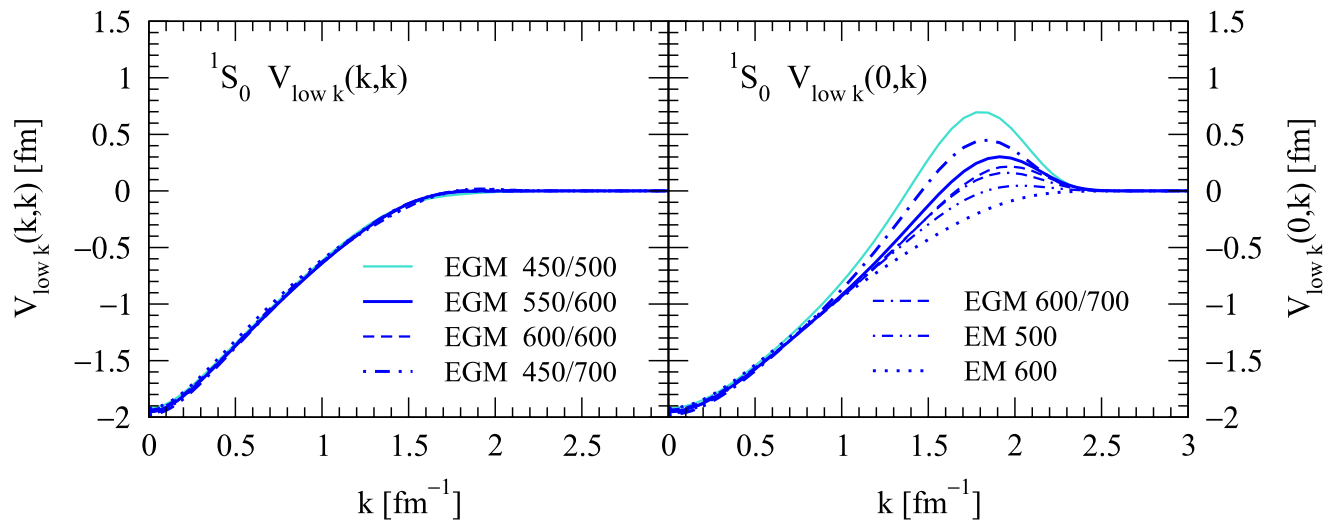
Renormalization of Chiral EFT Potentials

Diagonal

Off-diagonal



These are all our favorite Chiral EFT NN potentials...



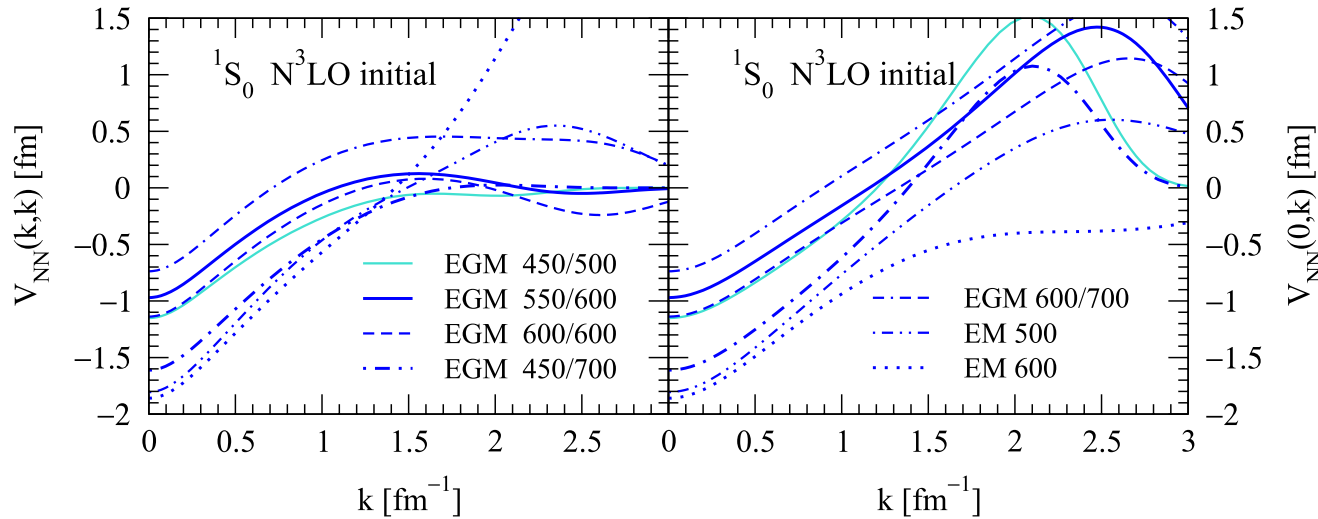
These are all our favorite Chiral EFT NN potentials...
at low momentum

Differences remain in off-diagonal matrix elements. Why?

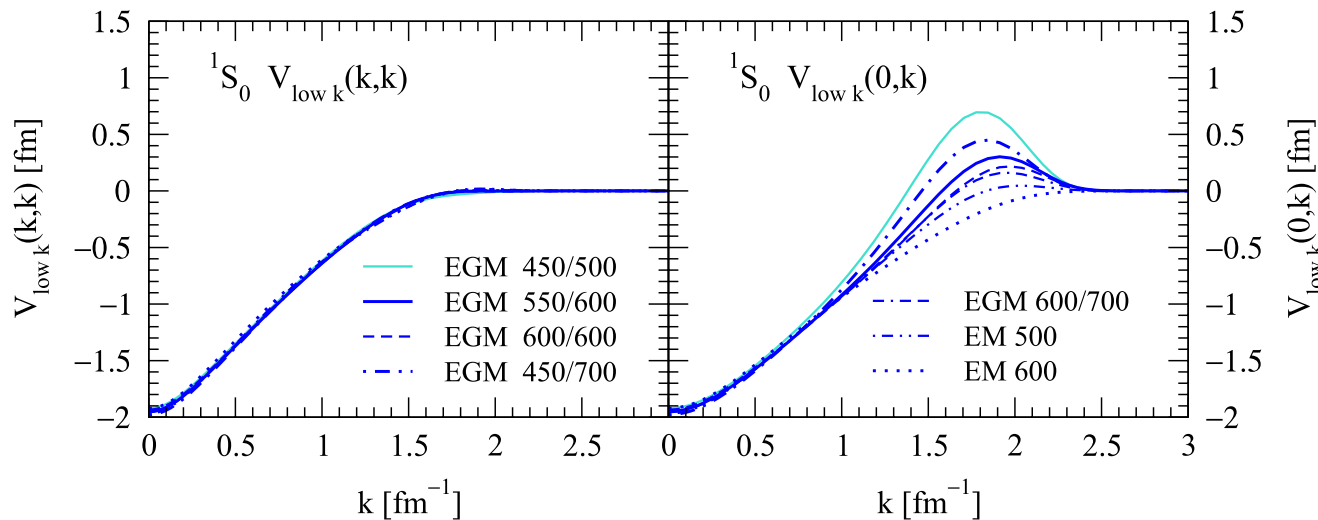
Renormalization of Chiral EFT Potentials

Diagonal

Off-diagonal



These are all our favorite Chiral EFT NN potentials...

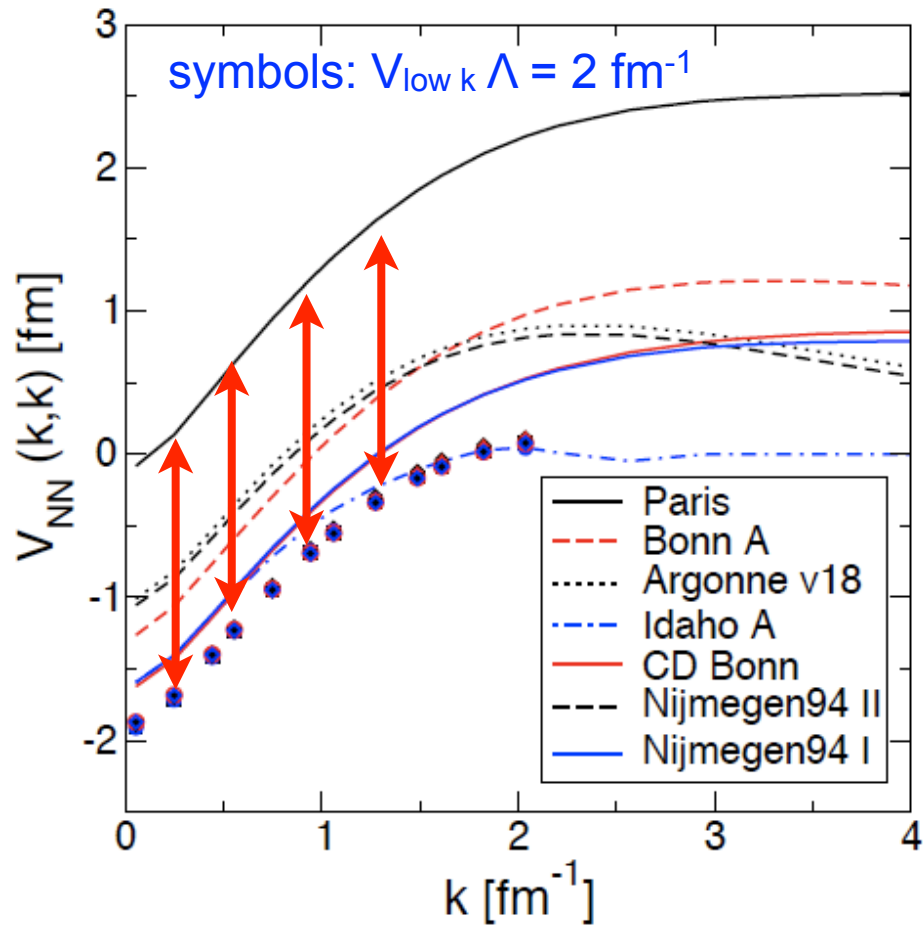


These are all our favorite Chiral EFT NN potentials...
at low momentum

Differences remain in off-diagonal matrix elements

Sensitive to agreement for phase shifts (not all fit perfectly)

Renormalization of NN Potentials



Why is it mostly a shift?

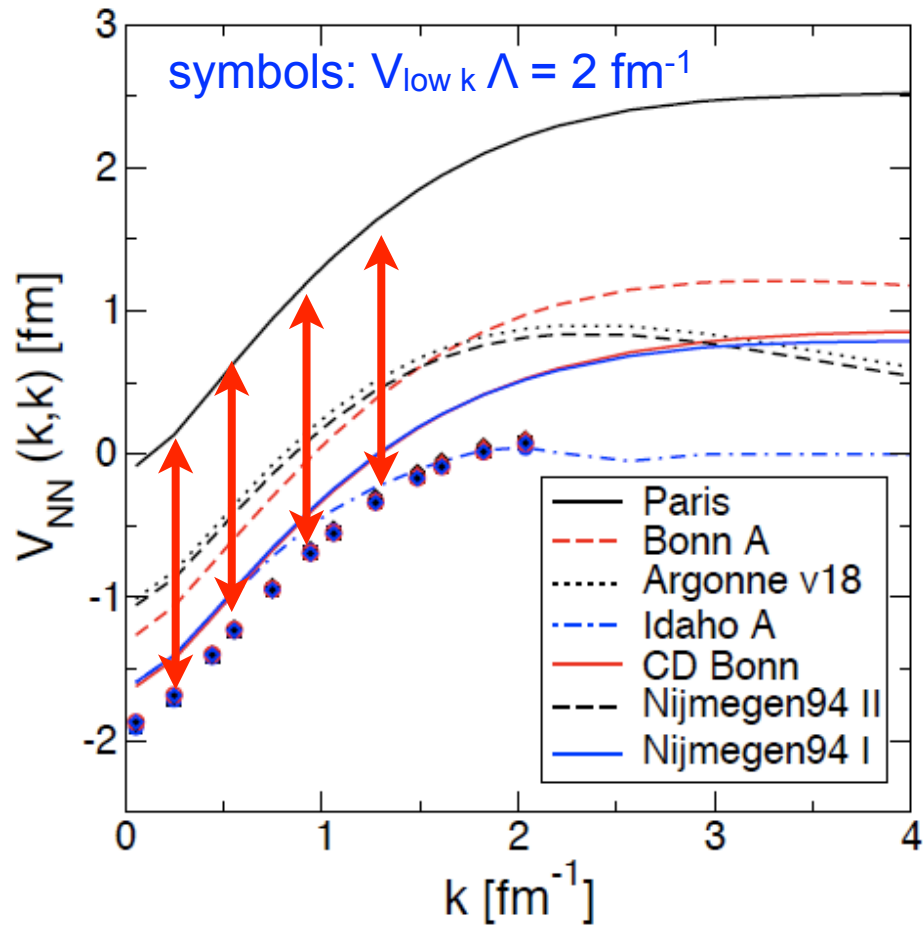


$$V_{\text{eff}} = V_L + \delta V_{\text{c.t.}}(\Lambda)$$

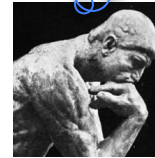
Overall effect of evolving to low momentum

Main effect is shift in momentum space

Renormalization of NN Potentials



Why is it mostly a shift?



$$V_{\text{eff}} = V_L + \delta V_{\text{c.t.}}(\Lambda)$$

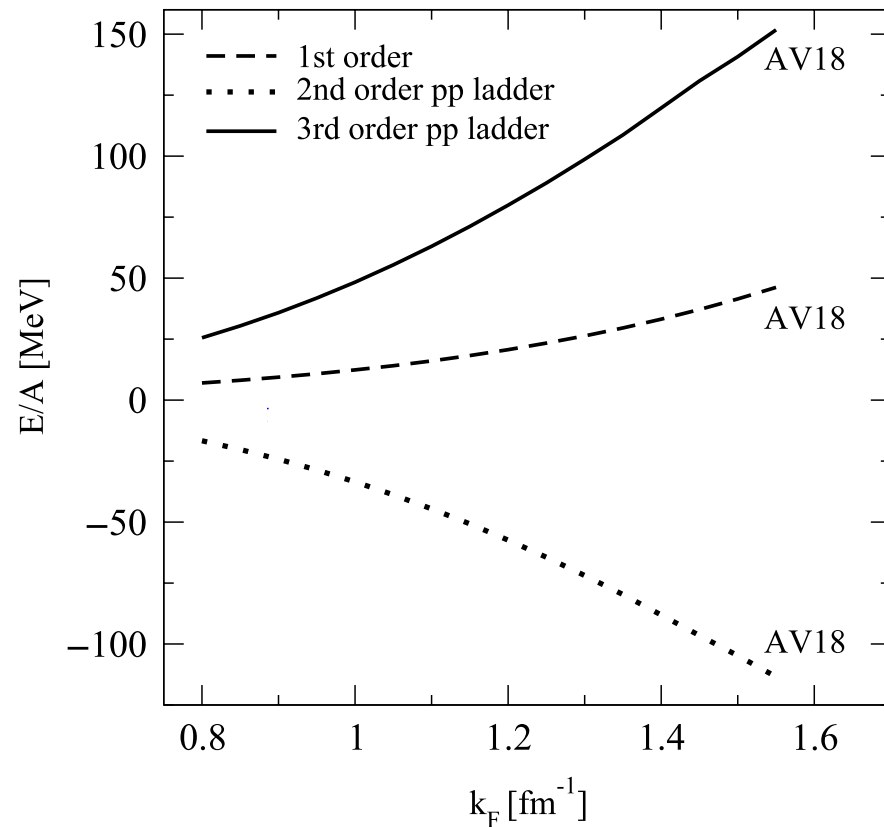
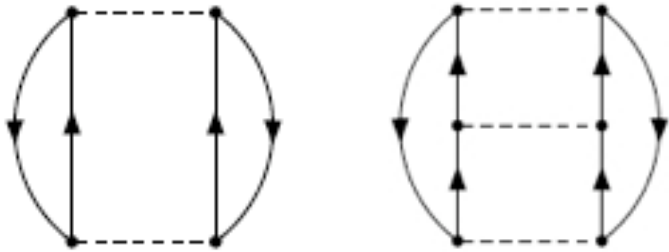
Overall effect of evolving to low momentum

Main effect is shift in momentum space – delta function
Removes hard core (unconstrained short-range physics)!

Improvements in Perturbation Theory

Explore improvements in symmetric infinite matter calculations

Order by order in **many-body perturbation theory (MBPT)**

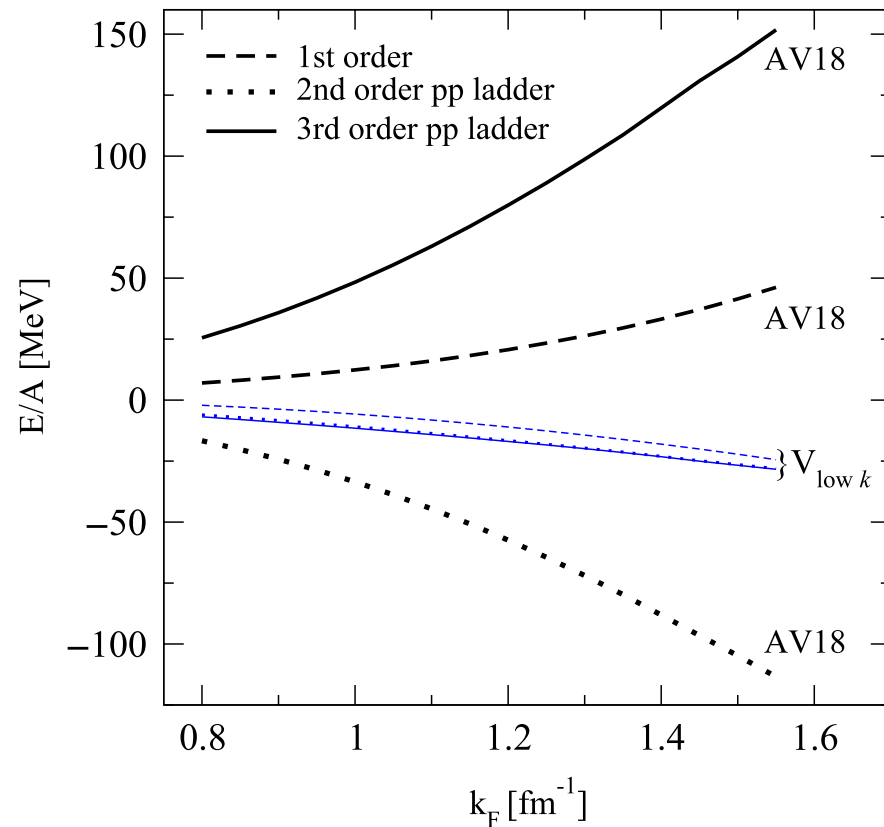
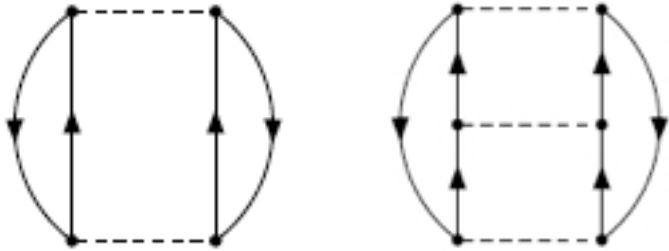


No clear convergence with increasing order in bare potential

Improvements in Perturbation Theory

Explore improvements in symmetric infinite matter calculations

Order by order in **many-body perturbation theory (MBPT)**



No clear convergence with increasing order in bare potential

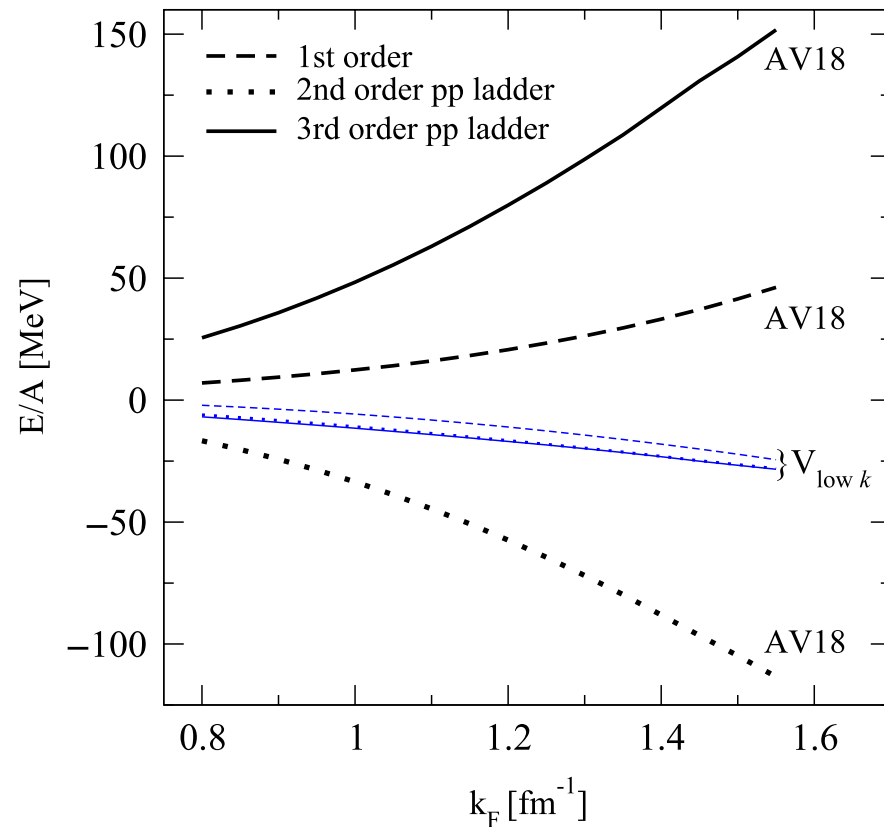
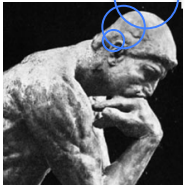
Significant improvement with low-momentum interactions!

Improvements in Perturbation Theory

Explore improvements in symmetric infinite matter calculations

Order by order in **many-body perturbation theory (MBPT)**

Ok, the interactions look perturbative, but something is wrong here...



No clear convergence with increasing order in bare potential

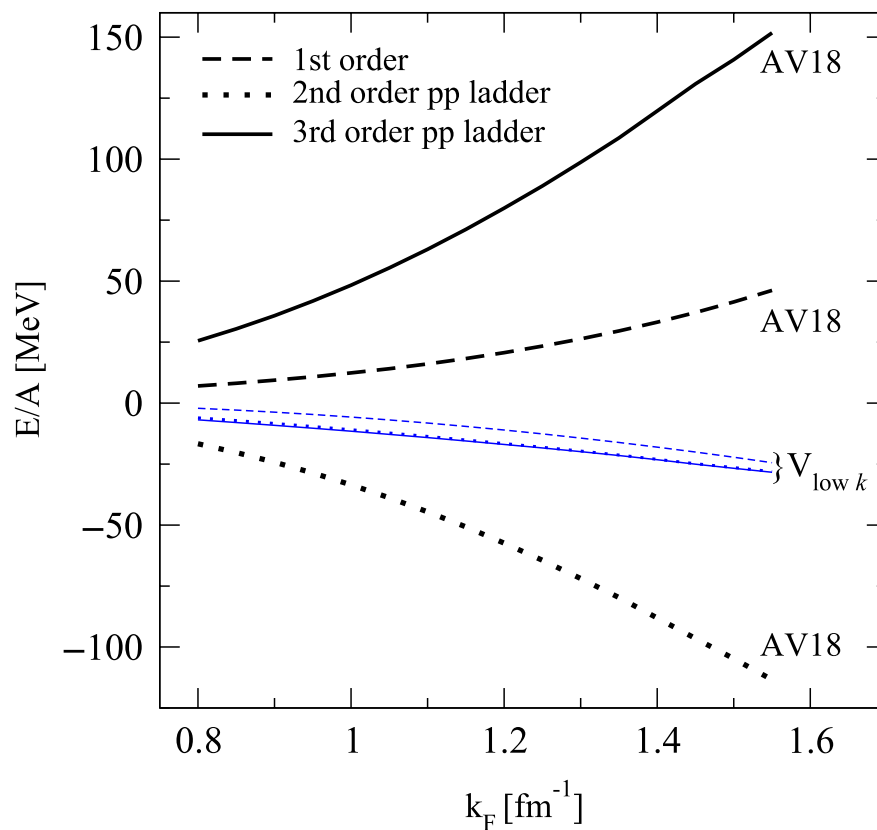
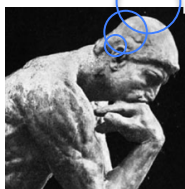
Significant improvement with low-momentum interactions!

Improvements in Perturbation Theory

Explore improvements in symmetric infinite matter calculations

Order by order in **many-body perturbation theory (MBPT)**

Ok, the interactions look perturbative, but something is wrong here...



No clear convergence with increasing order in bare potential

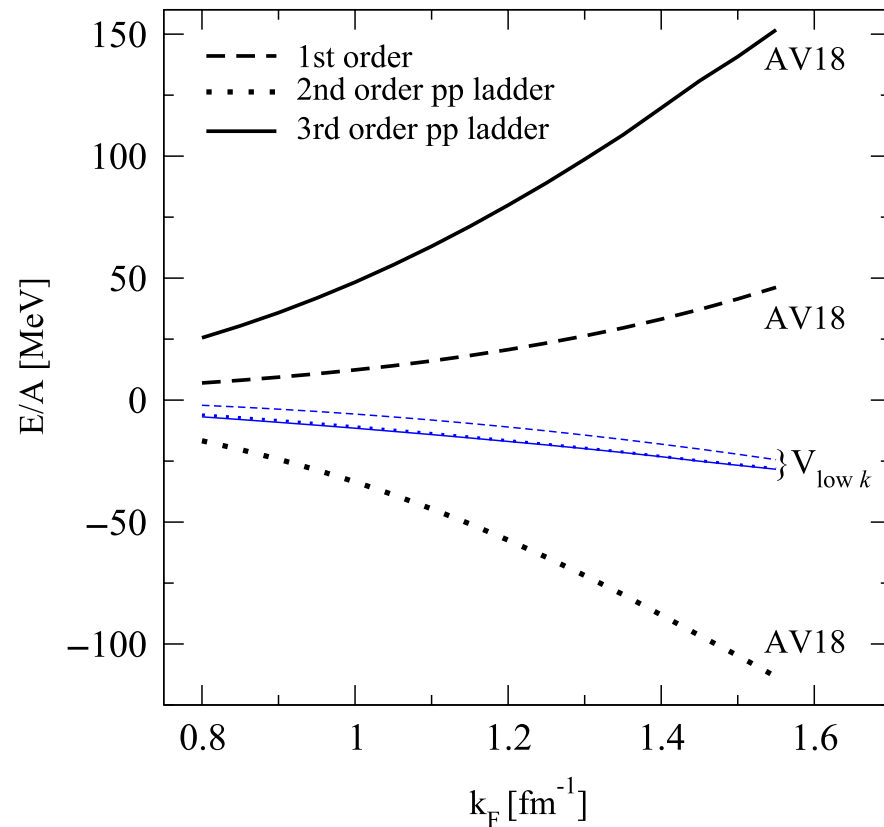
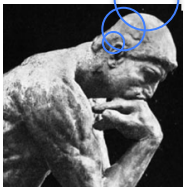
Significant improvement with low-momentum interactions!

Does not saturate – what might be missing?

Improvements in Perturbation Theory

$$H(\Lambda) = T + V_{\text{NN}}(\Lambda) + V_{\text{3N}}(\Lambda) + V_{\text{4N}}(\Lambda) + \dots$$

Ok, the interactions look perturbative, but something is wrong here...



No clear convergence with increasing order in bare potential

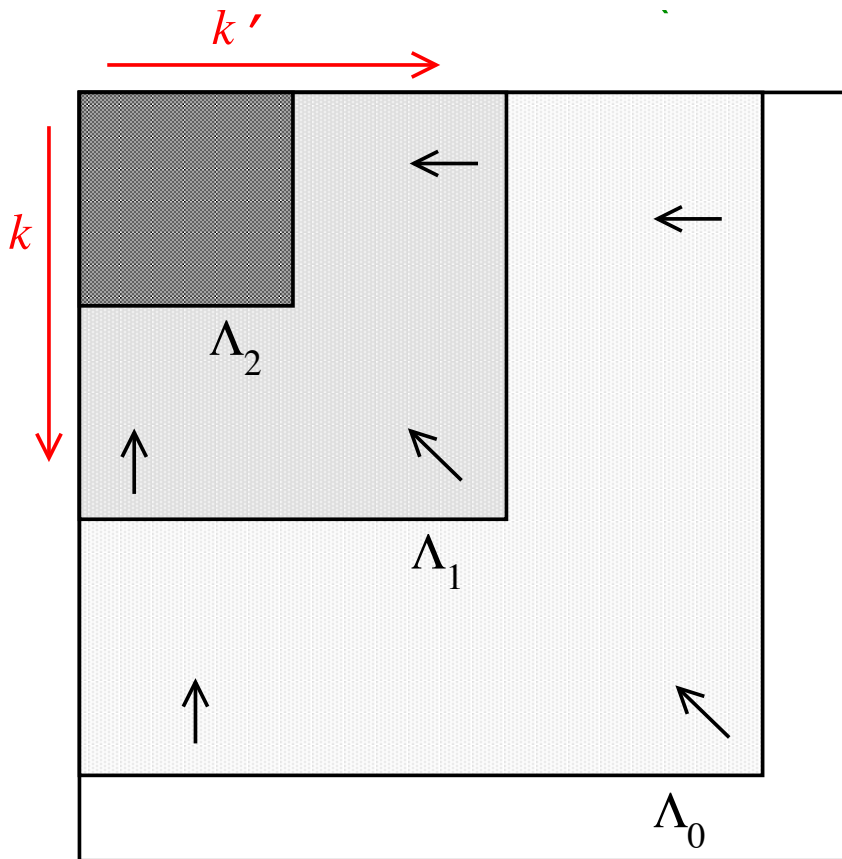
Significant improvement with low-momentum interactions!

Does not saturate – what might be missing?

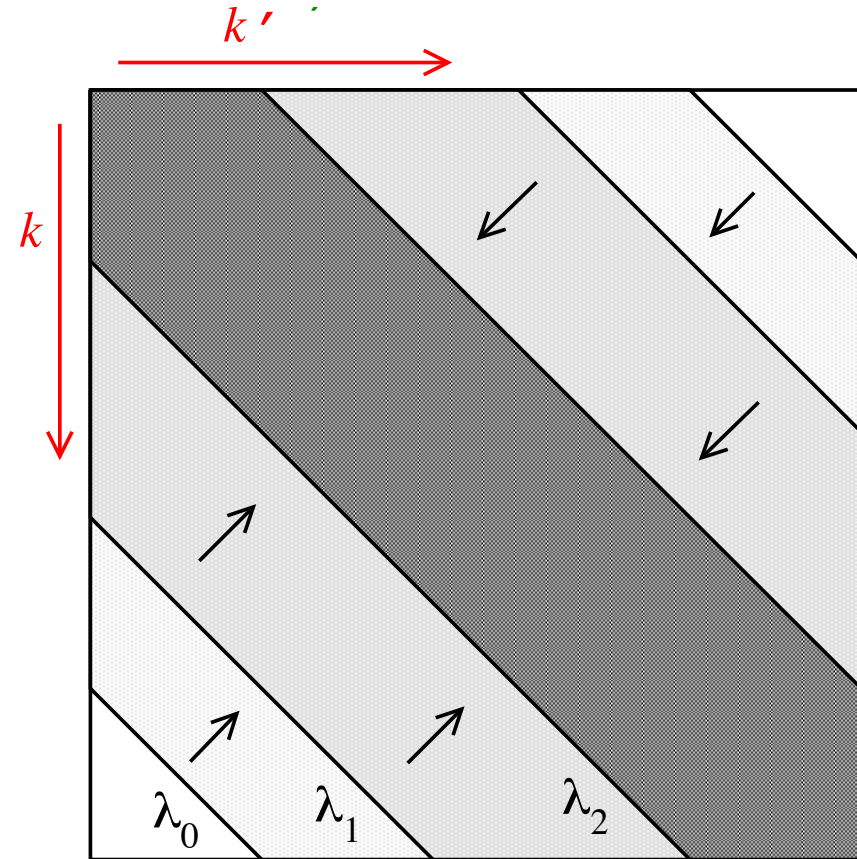
Similarity Renormalization Group

Wegner, Glazek/Wilson (1990s)

Complementary method to decouple low from high momenta



Decouples high-momentum



Similarity Renormalization Group

Drives Hamiltonian to band-diagonal

Similarity Renormalization Group

Wegner, Glazek/Wilson (1990s)

Apply a continuous unitary transformation, parameterized by s :

$$H = T + V \rightarrow H(s) = U(s)HU^\dagger(s) \equiv T + V(s)$$

where differentiating (exercise) yields:

$$\frac{dH(s)}{ds} = [\eta(s), H(s)] \quad \text{where} \quad \eta(s) \equiv \frac{dU(s)}{ds}U^\dagger(s)$$

Never explicitly construct unitary transformation

Instead **choose generator to obtain desired behavior:**

$$\eta(s) = [G(s), H(s)]$$

Many options, e.g.,

$$\eta(s) = [T, H(s)] \quad \text{Drives } H(s) \text{ to band-diagonal form}$$

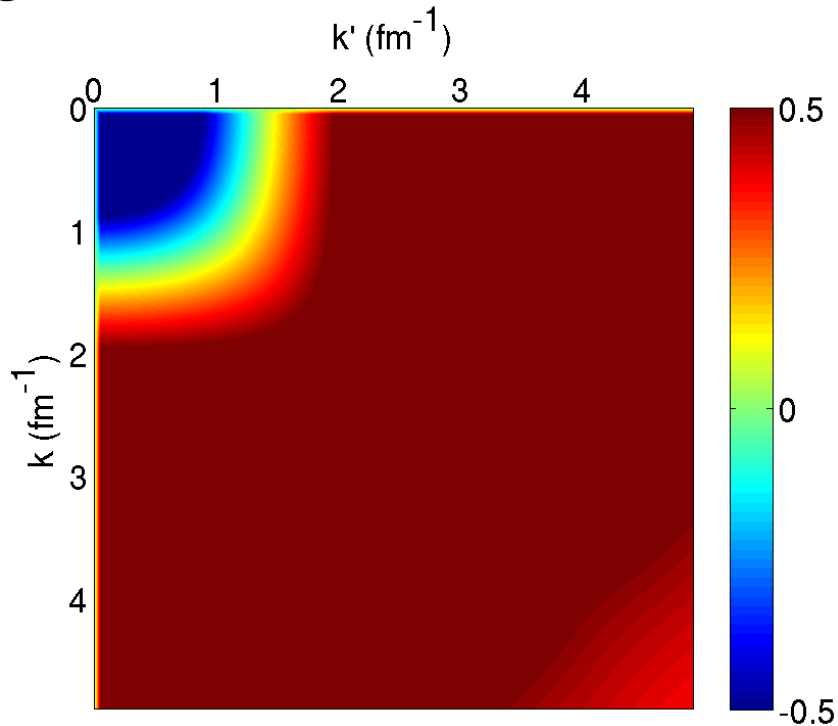
Illustration of SRG Flow

Drive H to band-diagonal form with kinetic-energy generator:

$$\eta(s) = [T, H(s)]$$

With alternate definition of flow parameter: $\lambda^2 = \frac{1}{\sqrt{s}}$

Argonne V_{18} 1S_0



$$\lambda = 8.0 \text{ fm}^{-1}$$

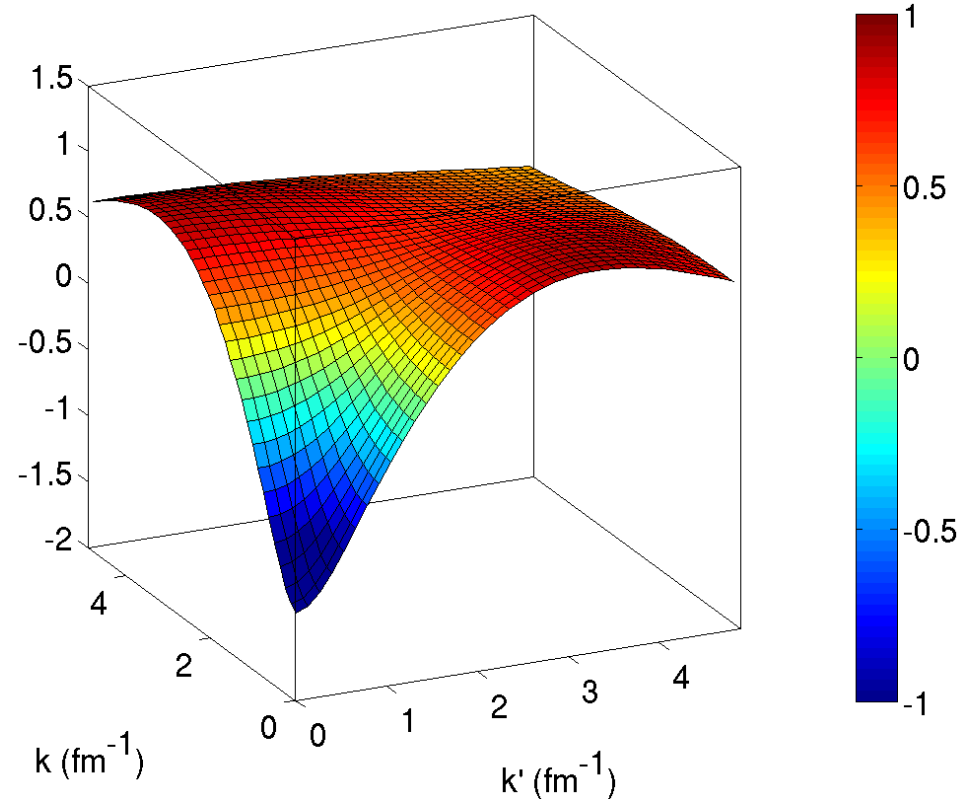


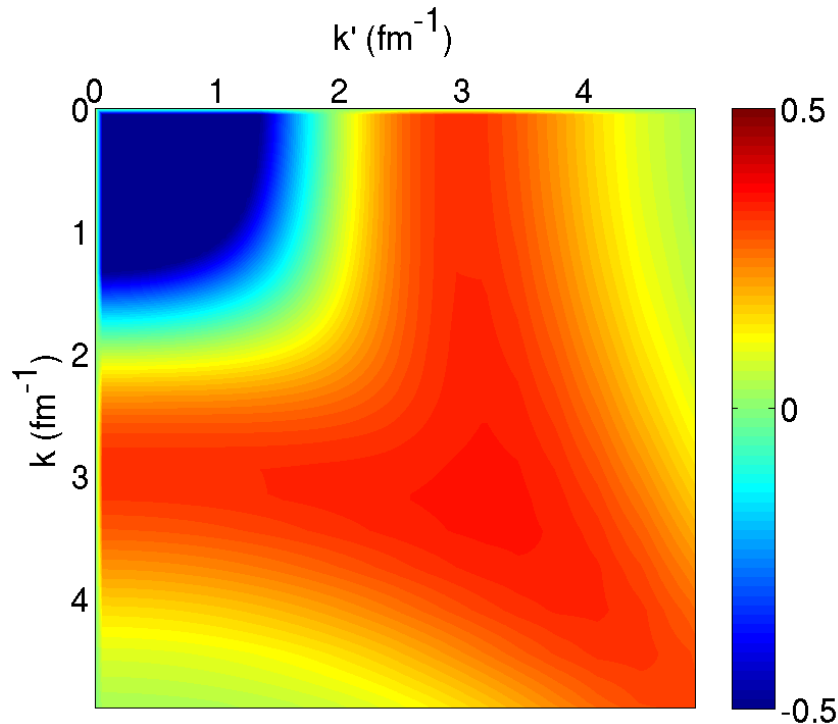
Illustration of SRG Flow

Drive H to band-diagonal form with standard choice:

$$\eta(s) = [T, H(s)]$$

With alternate definition of flow parameter: $\lambda^2 = \frac{1}{\sqrt{s}}$

Argonne V_{18} 1S_0



$$\lambda = 4.0 \text{ fm}^{-1}$$

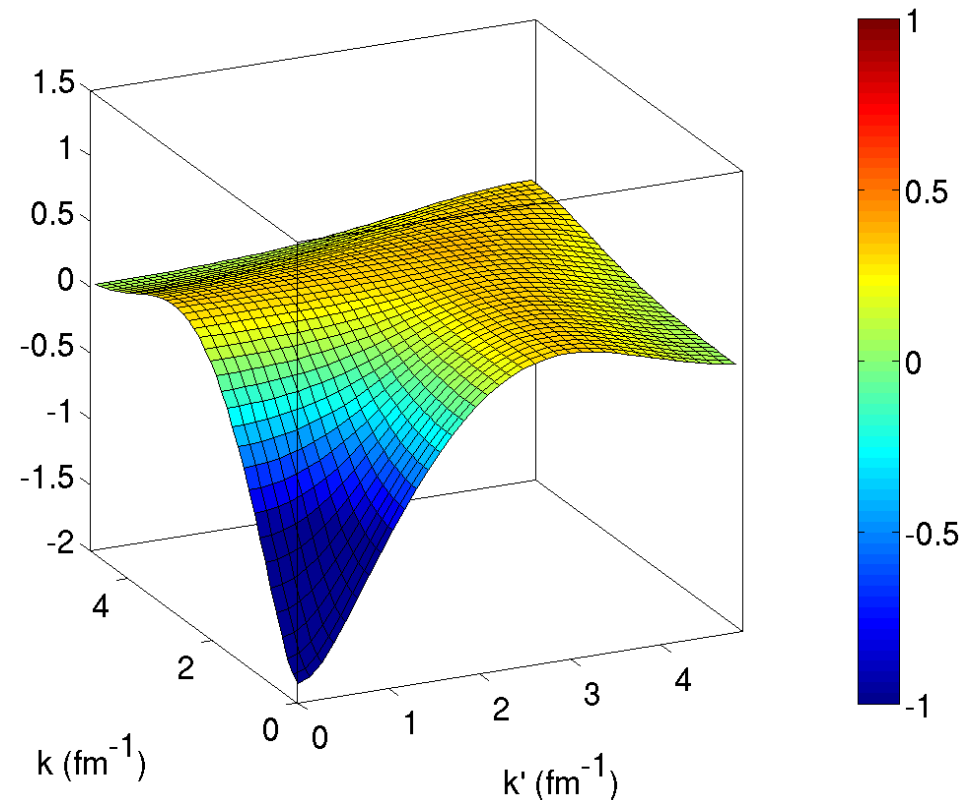


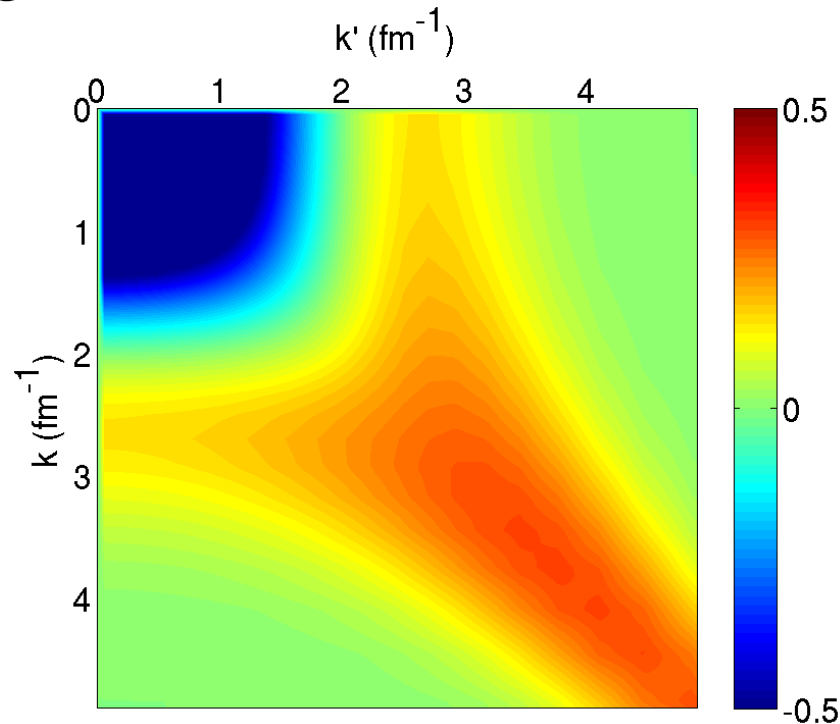
Illustration of SRG Flow

Drive H to band-diagonal form with standard choice:

$$\eta(s) = [T, H(s)]$$

With alternate definition of flow parameter: $\lambda^2 = \frac{1}{\sqrt{s}}$

Argonne V_{18} 1S_0



$$\lambda = 3.0 \text{ fm}^{-1}$$

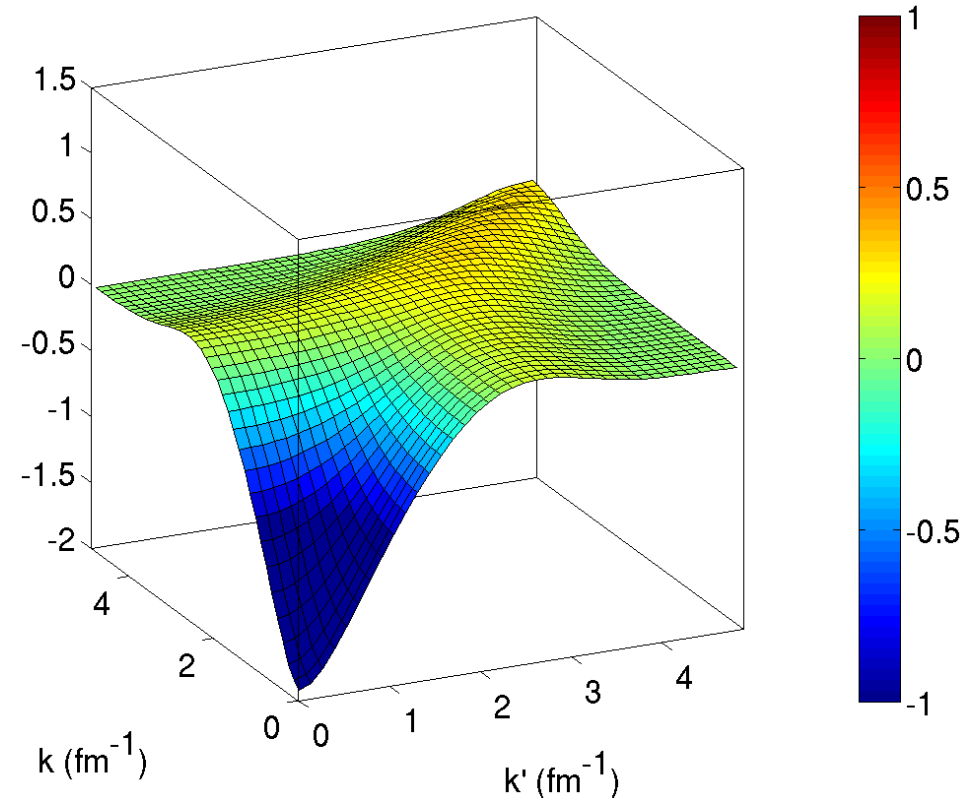


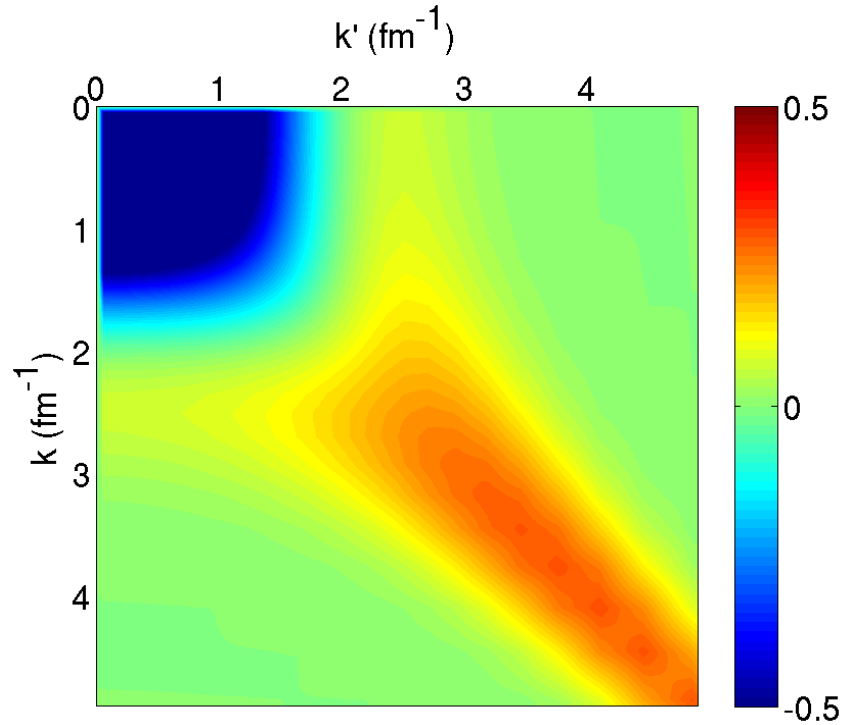
Illustration of SRG Flow

Drive H to band-diagonal form with standard choice:

$$\eta(s) = [T, H(s)]$$

With alternate definition of flow parameter: $\lambda^2 = \frac{1}{\sqrt{s}}$

Argonne V_{18} 1S_0



$$\lambda = 2.5 \text{ fm}^{-1}$$

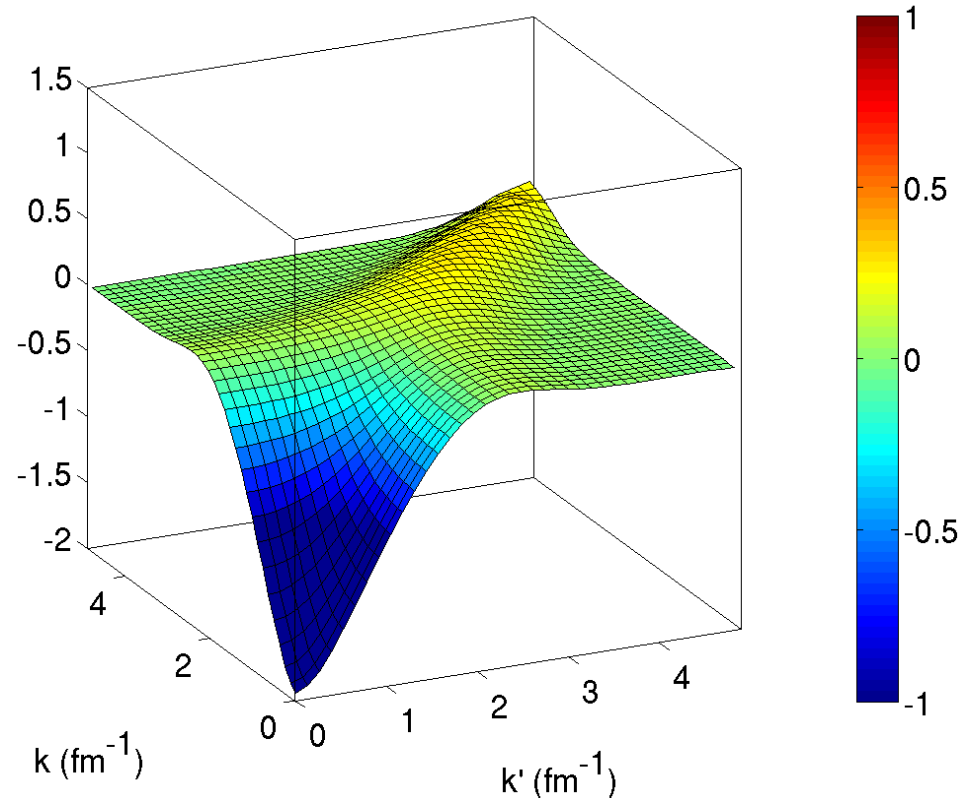


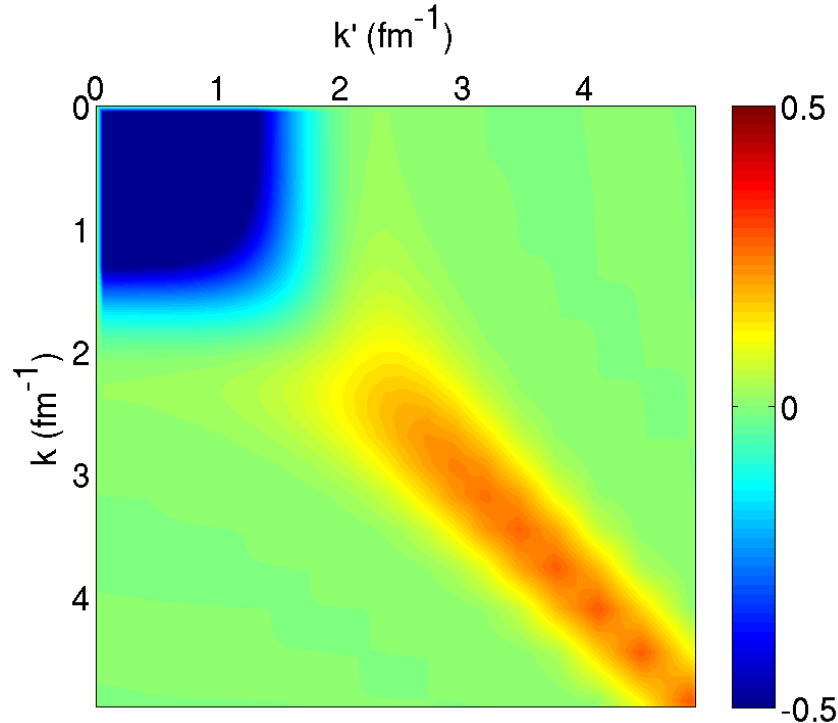
Illustration of SRG Flow

Drive H to band-diagonal form with standard choice:

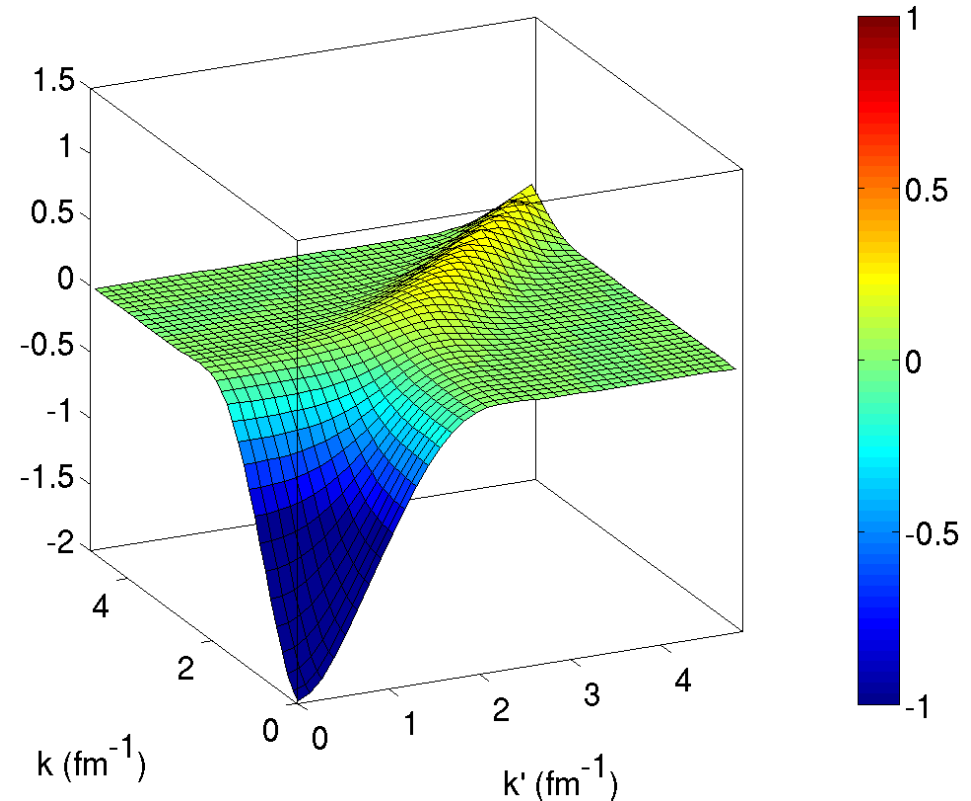
$$\eta(s) = [T, H(s)]$$

With alternate definition of flow parameter: $\lambda^2 = \frac{1}{\sqrt{s}}$

Argonne V_{18} 1S_0



$$\lambda = 2.0 \text{ fm}^{-1}$$

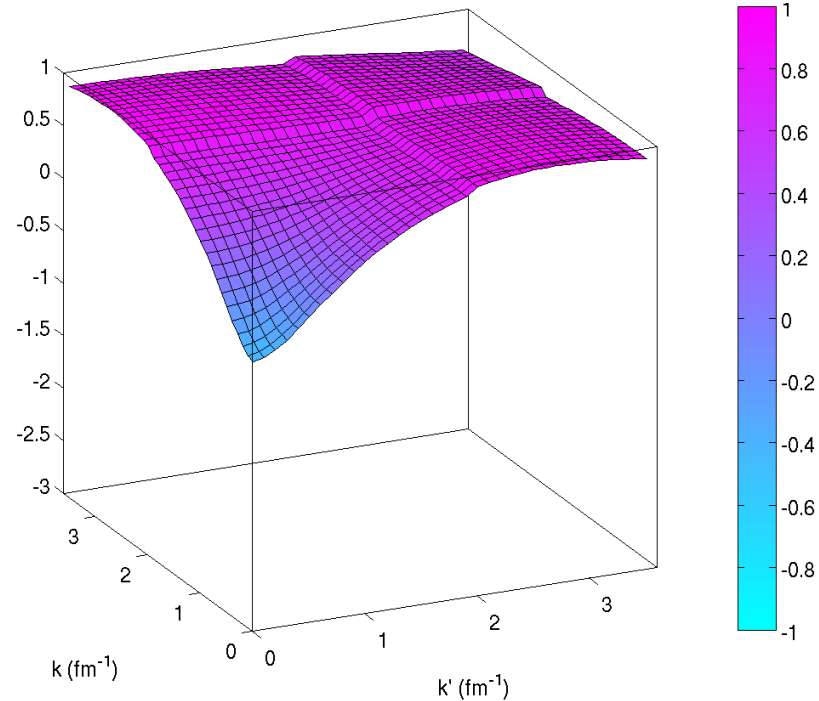
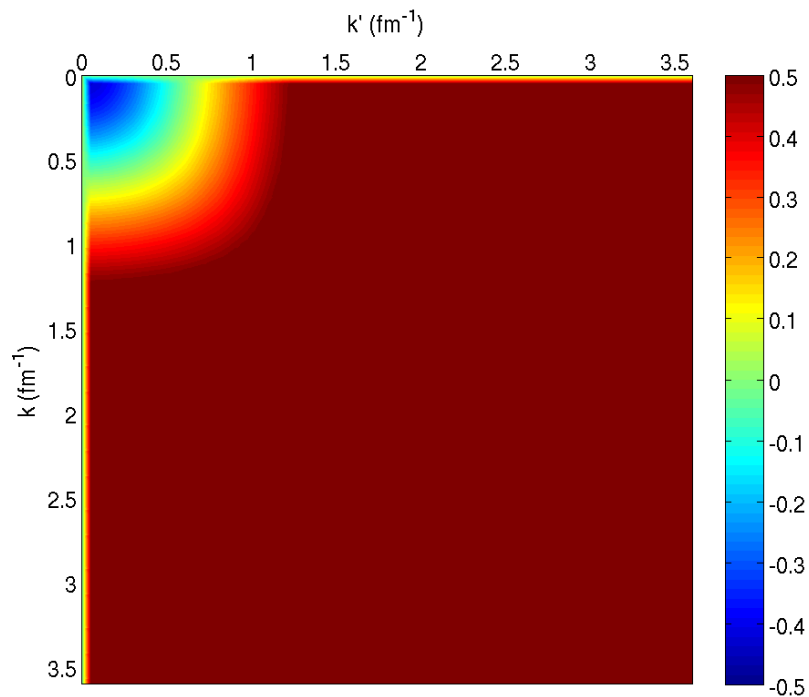


Other Generator Choices: Block Diagonal

Create block diagonal form like $V_{\text{low}k}$?

$$G(s) = H_{\text{BD}} = \begin{pmatrix} PH(s)P & 0 \\ 0 & QH(s)Q \end{pmatrix}$$

With alternate definition of flow parameter: $\lambda^2 = \frac{1}{\sqrt{s}}$



Argonne V_{18} 3S_1

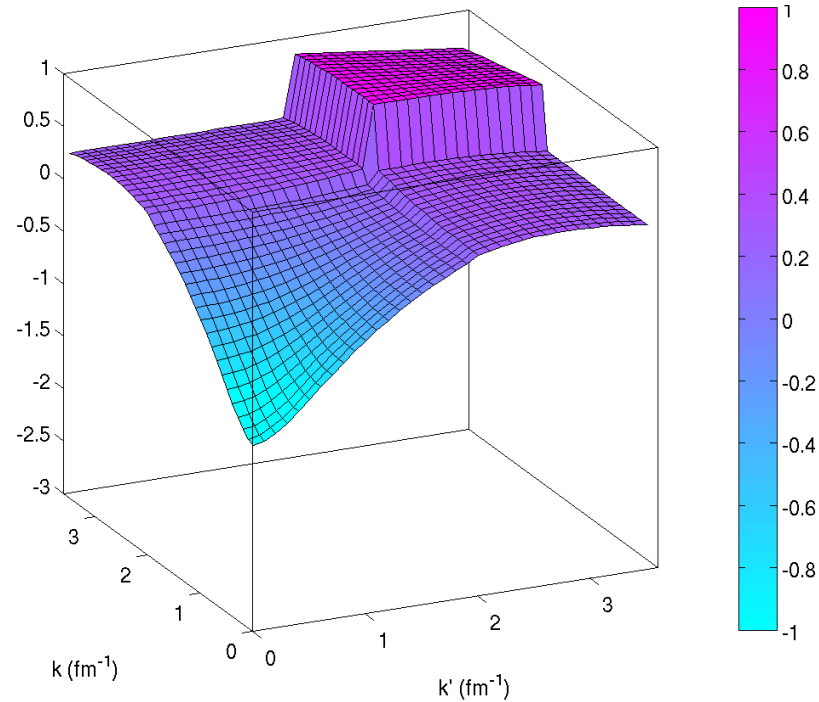
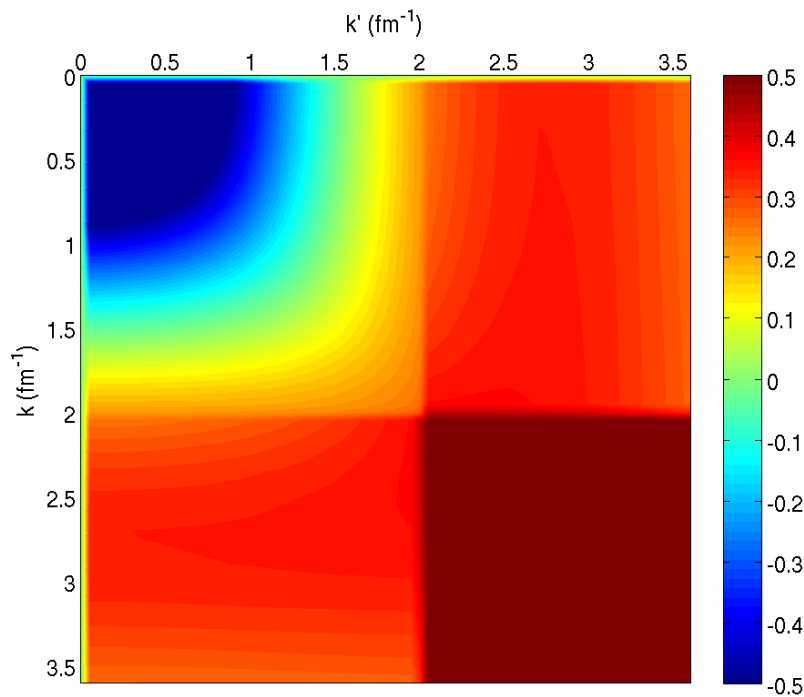
$\lambda = 10.0 \text{ fm}^{-1}$

Other Generator Choices: Block Diagonal

Create block diagonal form like $V_{\text{low}k}$?

$$G(s) = H_{\text{BD}} = \begin{pmatrix} PH(s)P & 0 \\ 0 & QH(s)Q \end{pmatrix}$$

With alternate definition of flow parameter: $\lambda^2 = \frac{1}{\sqrt{s}}$



Argonne V_{18} 3S_1

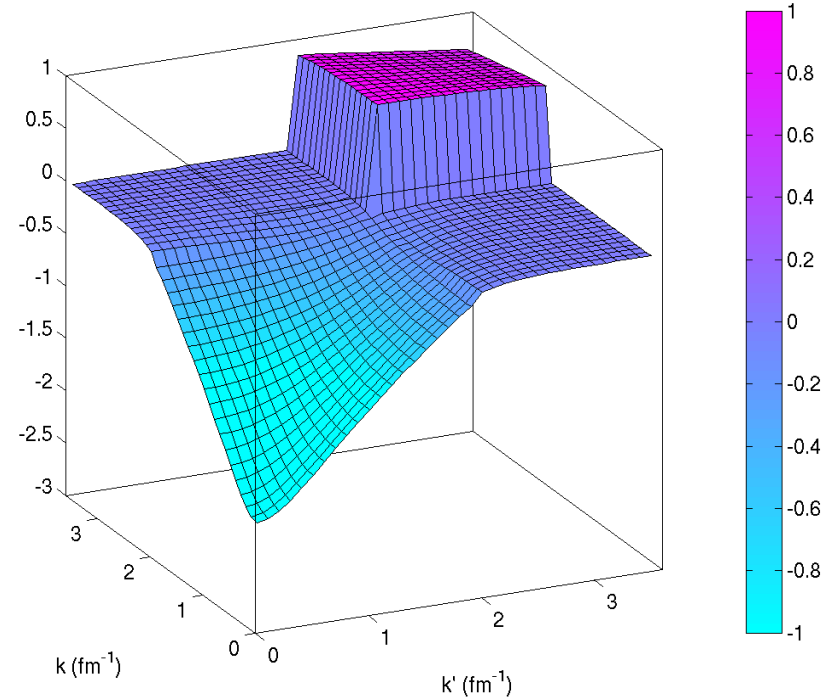
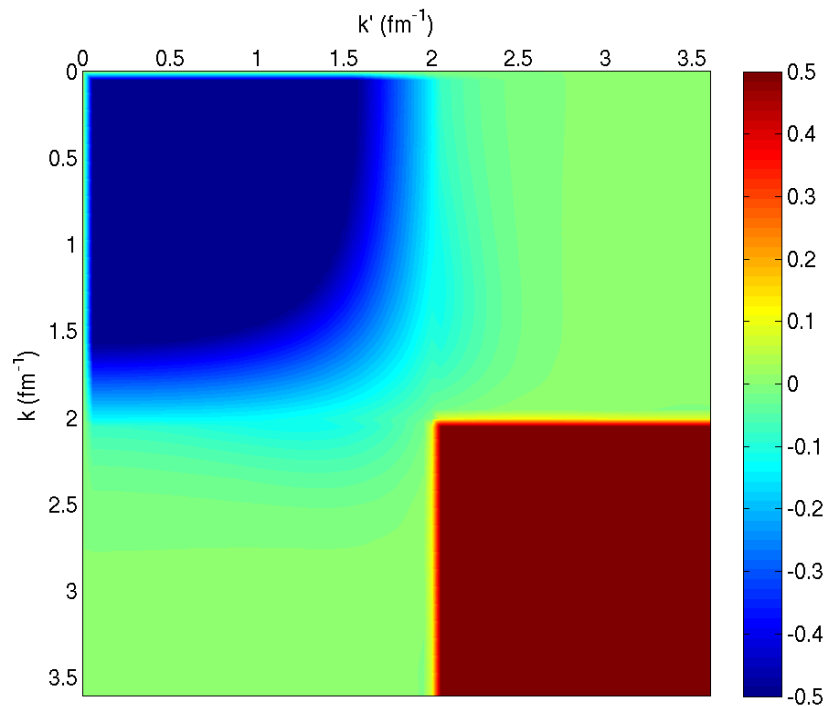
$\lambda = 5.0 \text{ fm}^{-1}$

Other Generator Choices: Block Diagonal

Create block diagonal form like $V_{\text{low}k}$?

$$G(s) = H_{\text{BD}} = \begin{pmatrix} PH(s)P & 0 \\ 0 & QH(s)Q \end{pmatrix}$$

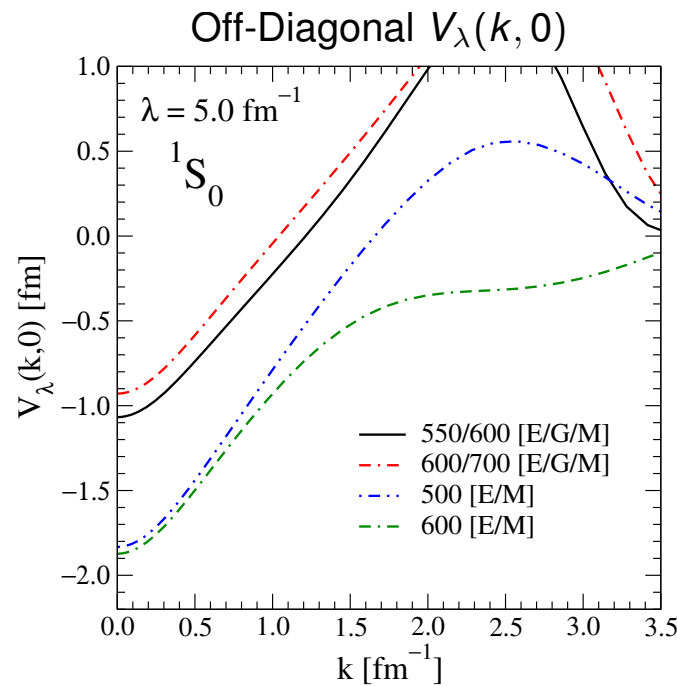
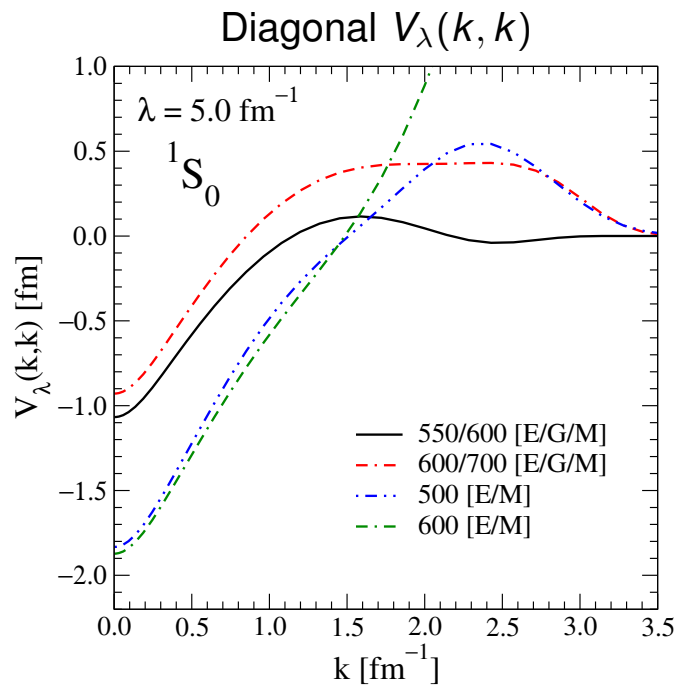
With alternate definition of flow parameter: $\lambda^2 = \frac{1}{\sqrt{s}}$



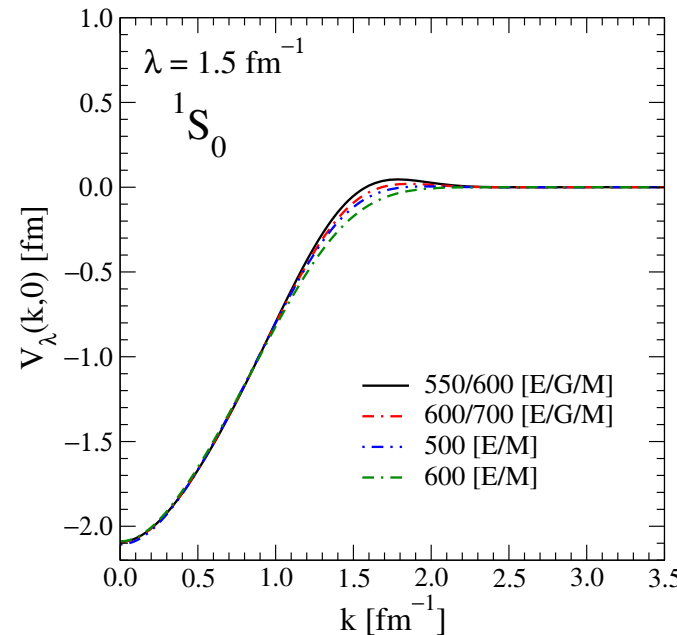
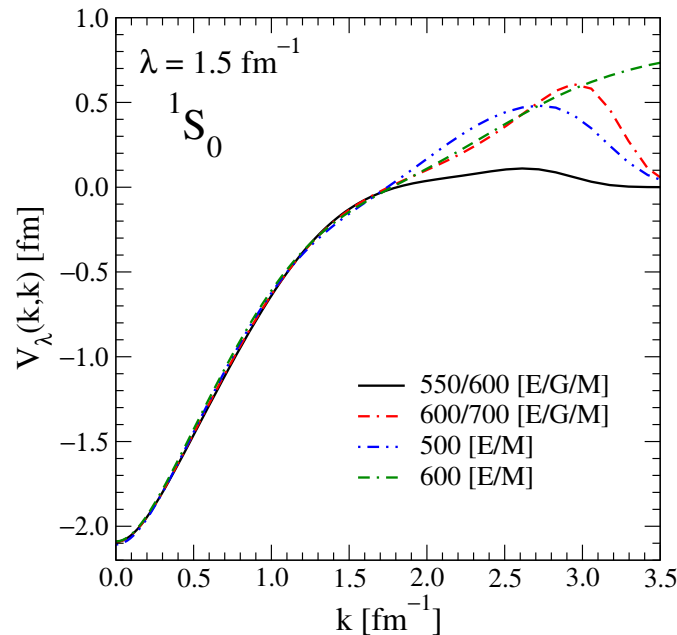
Argonne V_{18} 3S_1

$\lambda = 2.0 \text{ fm}^{-1}$

SRG Renormalization of Chiral EFT Potentials



These are all our favorite Chiral EFT NN potentials...



These are all our favorite Chiral EFT NN potentials...

SRG evolved

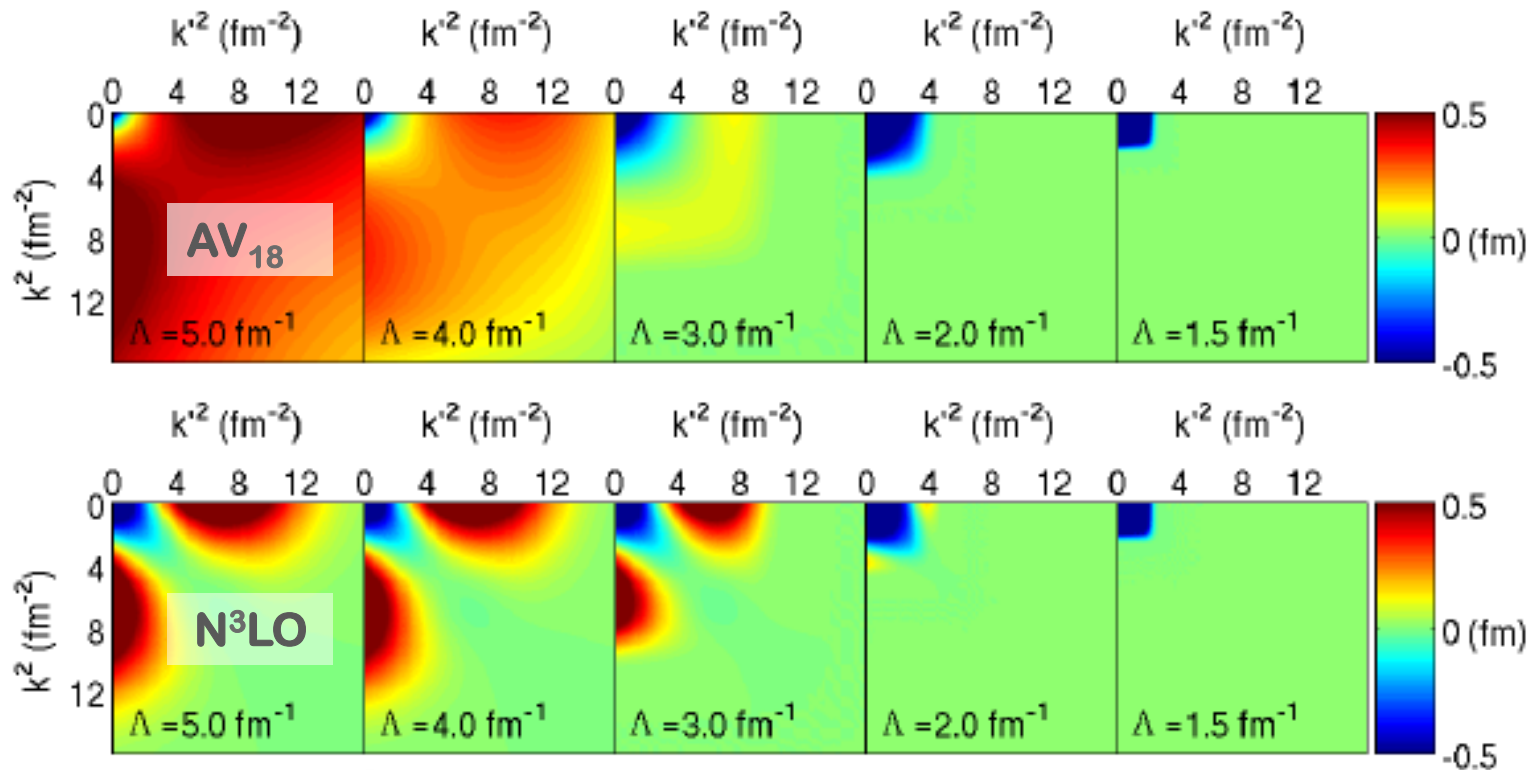
Exhibit similar “universal” behavior as low-momentum interactions!

Renormalization of Nuclear Interactions

$$H(\Lambda) = T + V_{\text{NN}}(\Lambda) + V_{\text{3N}}(\Lambda) + V_{\text{4N}}(\Lambda) + \dots$$

Evolve momentum resolution scale of chiral interactions from initial Λ_χ
 Remove coupling to high momenta, low-energy physics unchanged

Bogner, Kuo, Schwenk, Furnstahl



Universal at
low-momentum

$V_{\text{low } k}(\Lambda)$: lower cutoffs advantageous for nuclear structure calculations

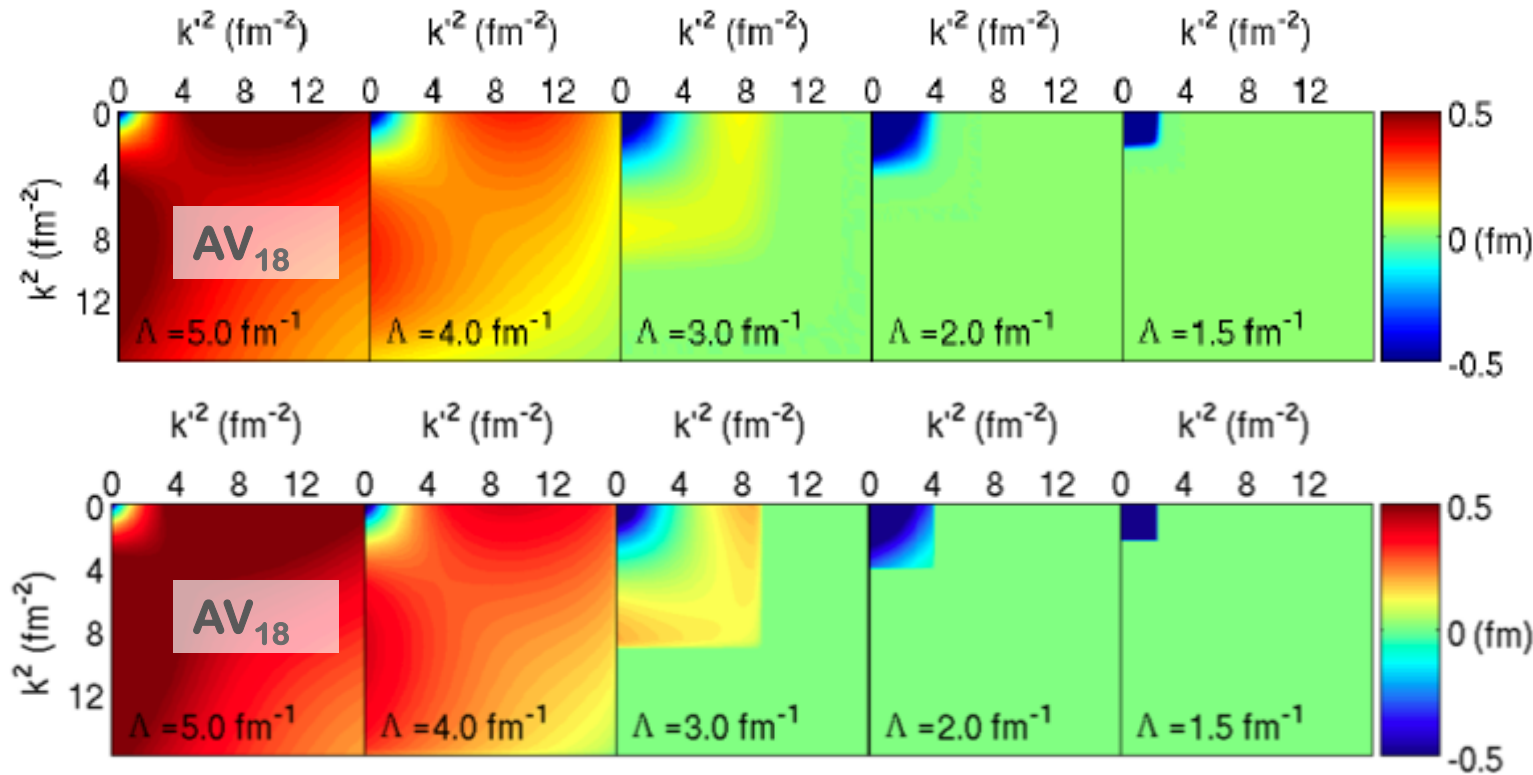
Smooth vs. Sharp Cutoffs

$$H(\Lambda) = T + V_{\text{NN}}(\Lambda) + V_{\text{3N}}(\Lambda) + V_{\text{4N}}(\Lambda) + \dots$$

Can have sharp as well as smooth cutoffs

Remove coupling to high momenta, low-energy physics unchanged

Bogner, Kuo, Schwenk, Furnstahl

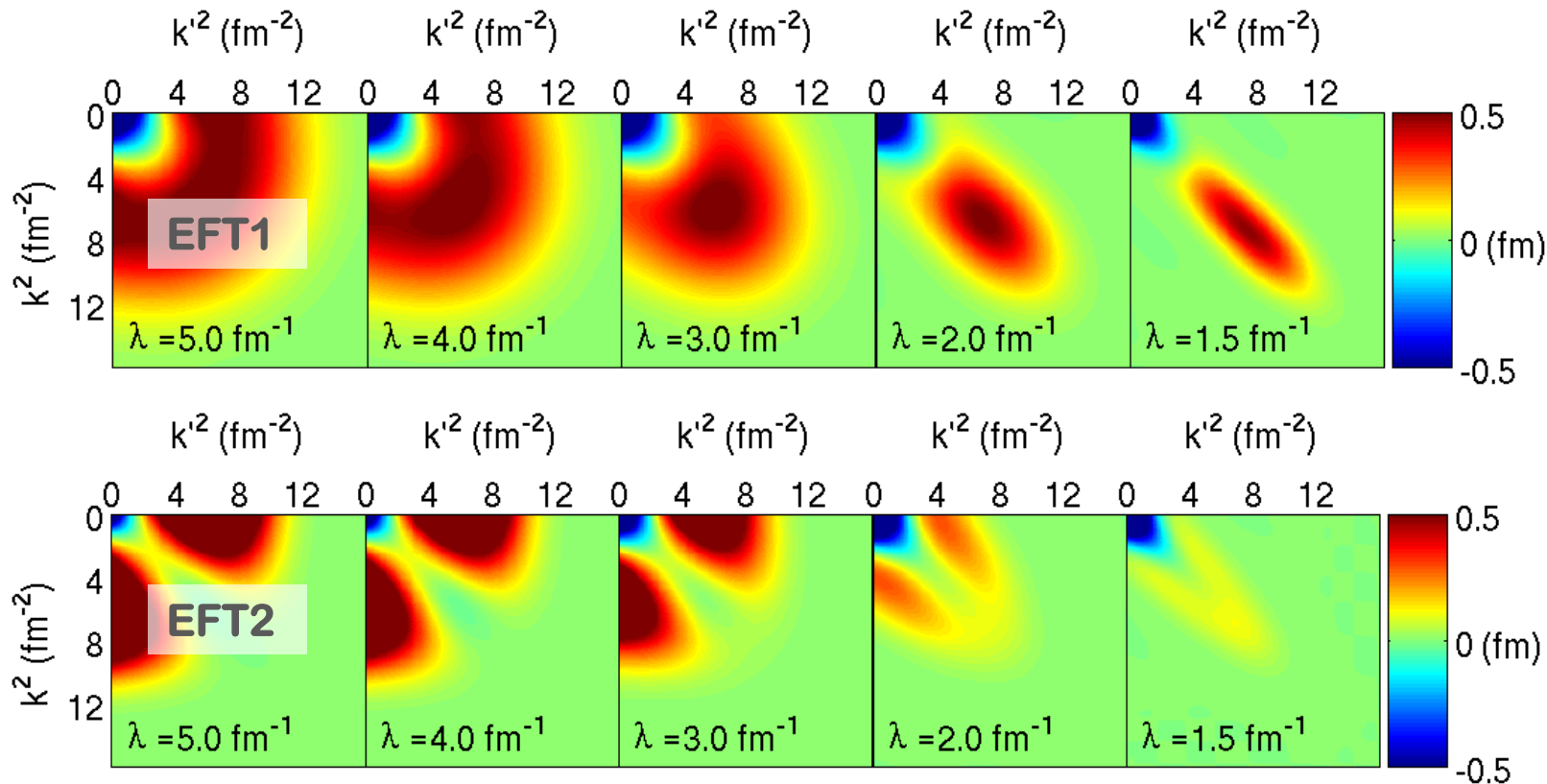


Similar but not exact same results – will be differences in calculations

SRG-Evolution of Different Initial Potentials

$$H(\Lambda) = T + V_{\text{NN}}(\Lambda) + V_{\text{3N}}(\Lambda) + V_{\text{4N}}(\Lambda) + \dots$$

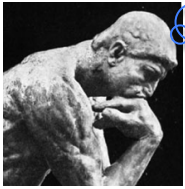
SRG evolution of two different chiral EFT potentials



Lots of pretty pictures, but how does it actually help?

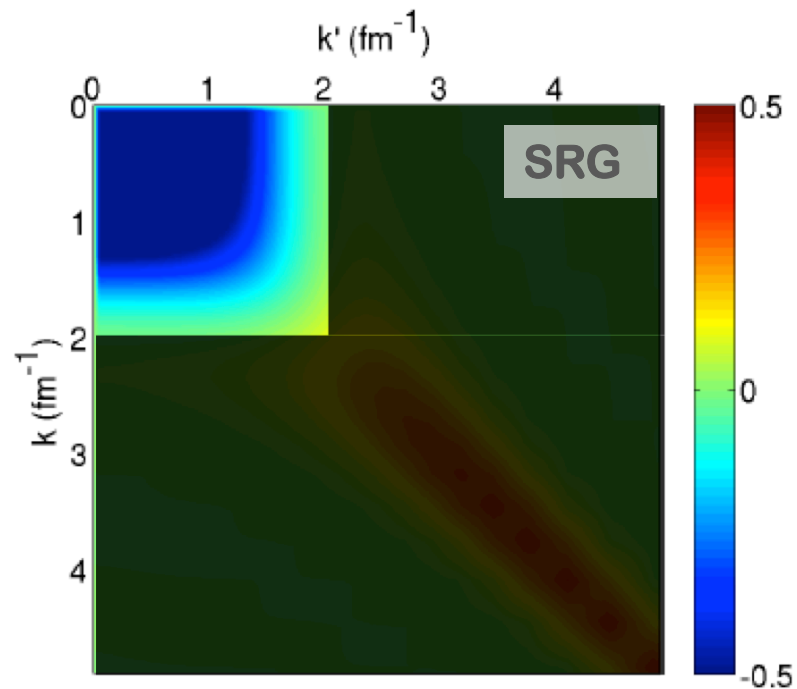
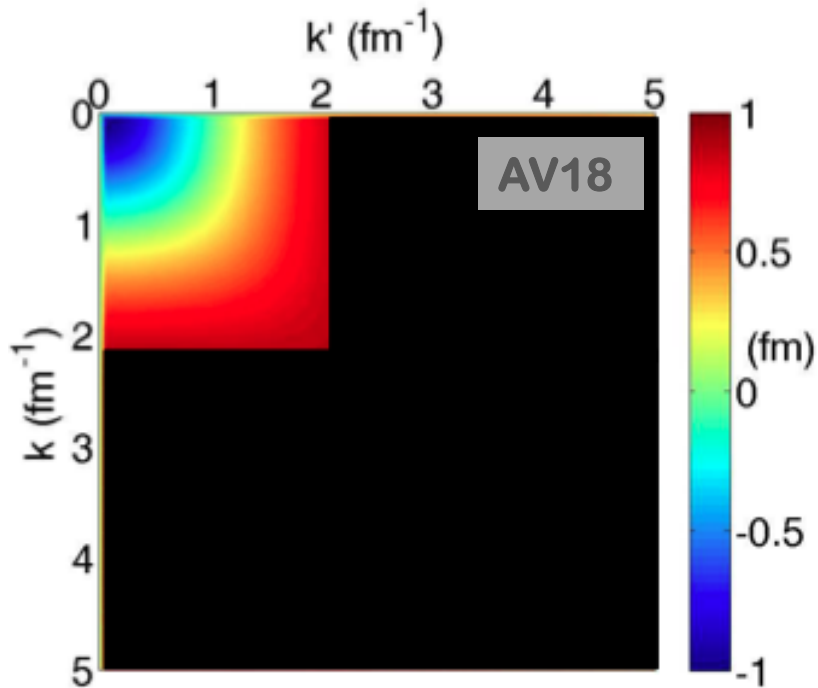
Revisit Low-Pass Filter Idea

Ok, high momentum is a pain. I wonder what would happen to low-energy observables...



Low-to-high momentum makes life difficult for low-energy nuclear theorists

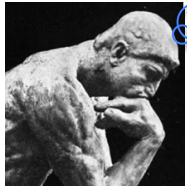
What's the difference now?



$$V_{\text{filter}}(k', k) \equiv 0; \quad k, k' > 2.2 \text{ MeV}$$

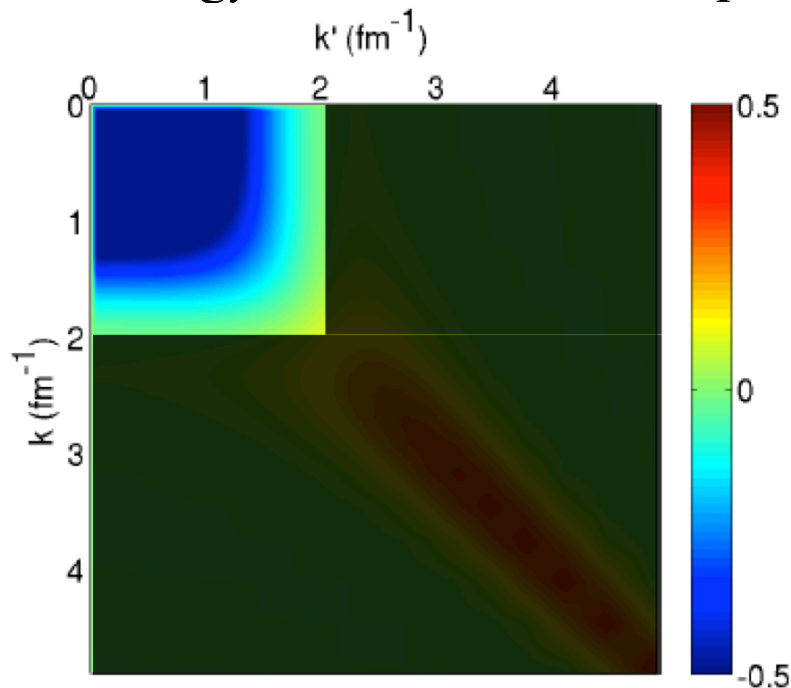
Revisit Low-Pass Filter Idea

Ok, high momentum is a pain. I wonder what would happen to low-energy observables...

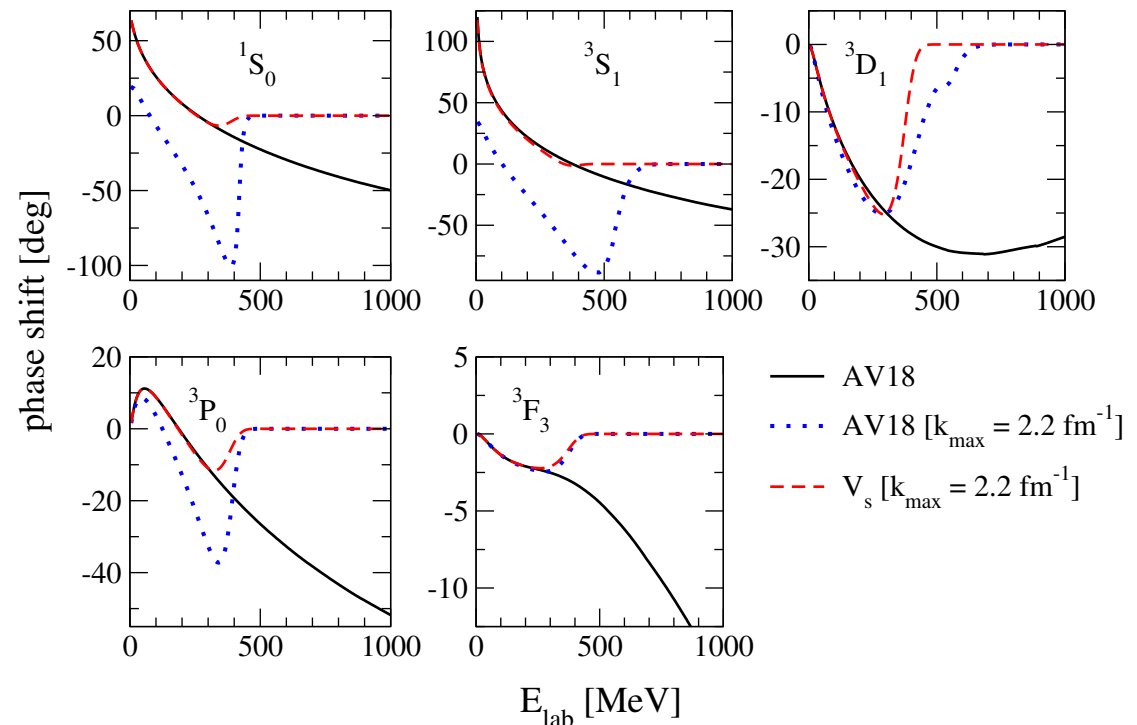


Low-to-high momentum makes life difficult for low-energy nuclear theorists

Low-energy observables were preserved – now sharp cut makes sense!



$$V_{\text{filter}}(k', k) \equiv 0; \quad k, k' > 2.2 \text{ MeV}$$

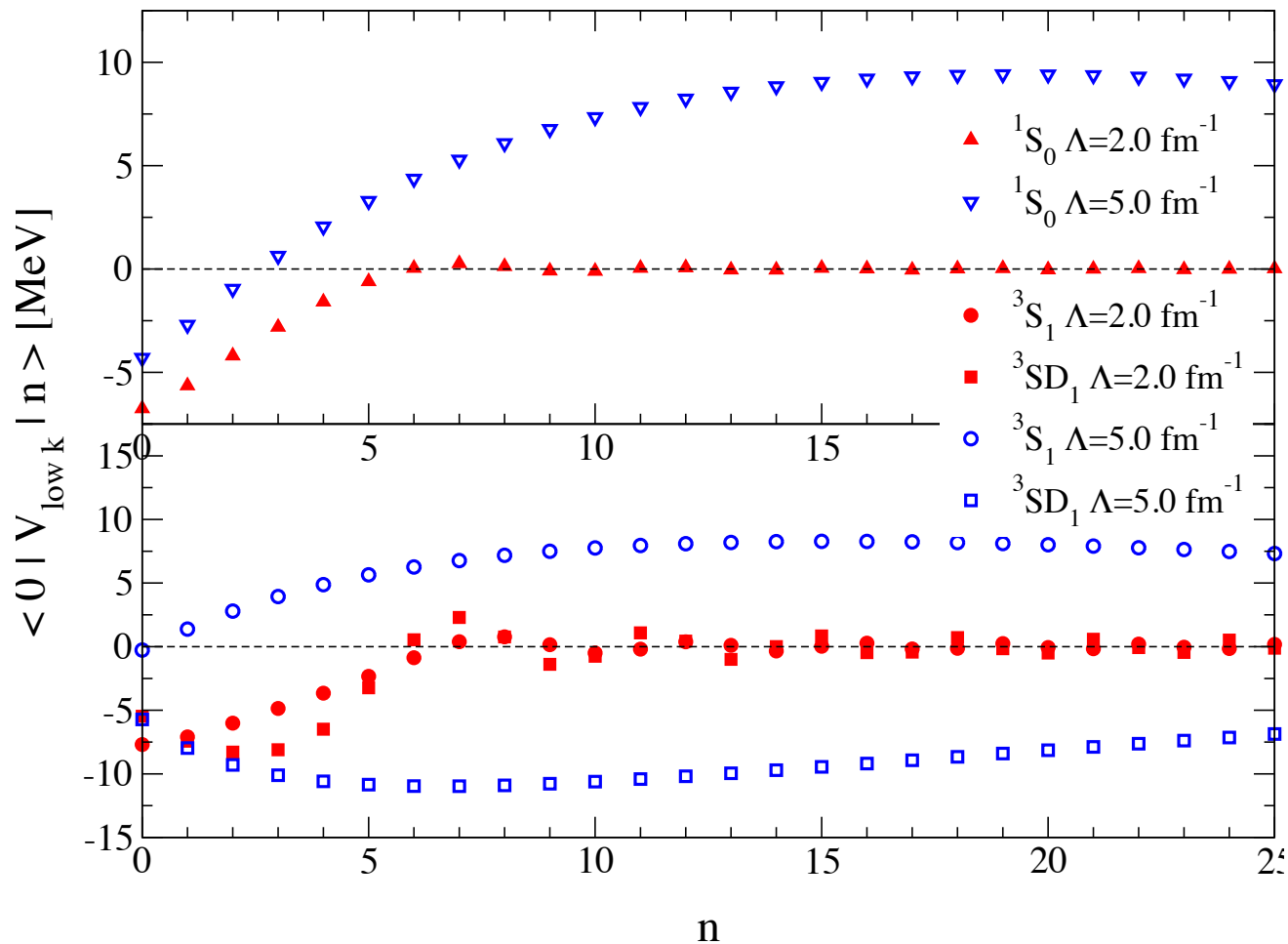


Benefits of Lower Cutoffs

Often work in HO basis – does this make a difference there?

Removes coupling from low-to-high harmonic oscillator states

Expect to speed convergence in HO basis



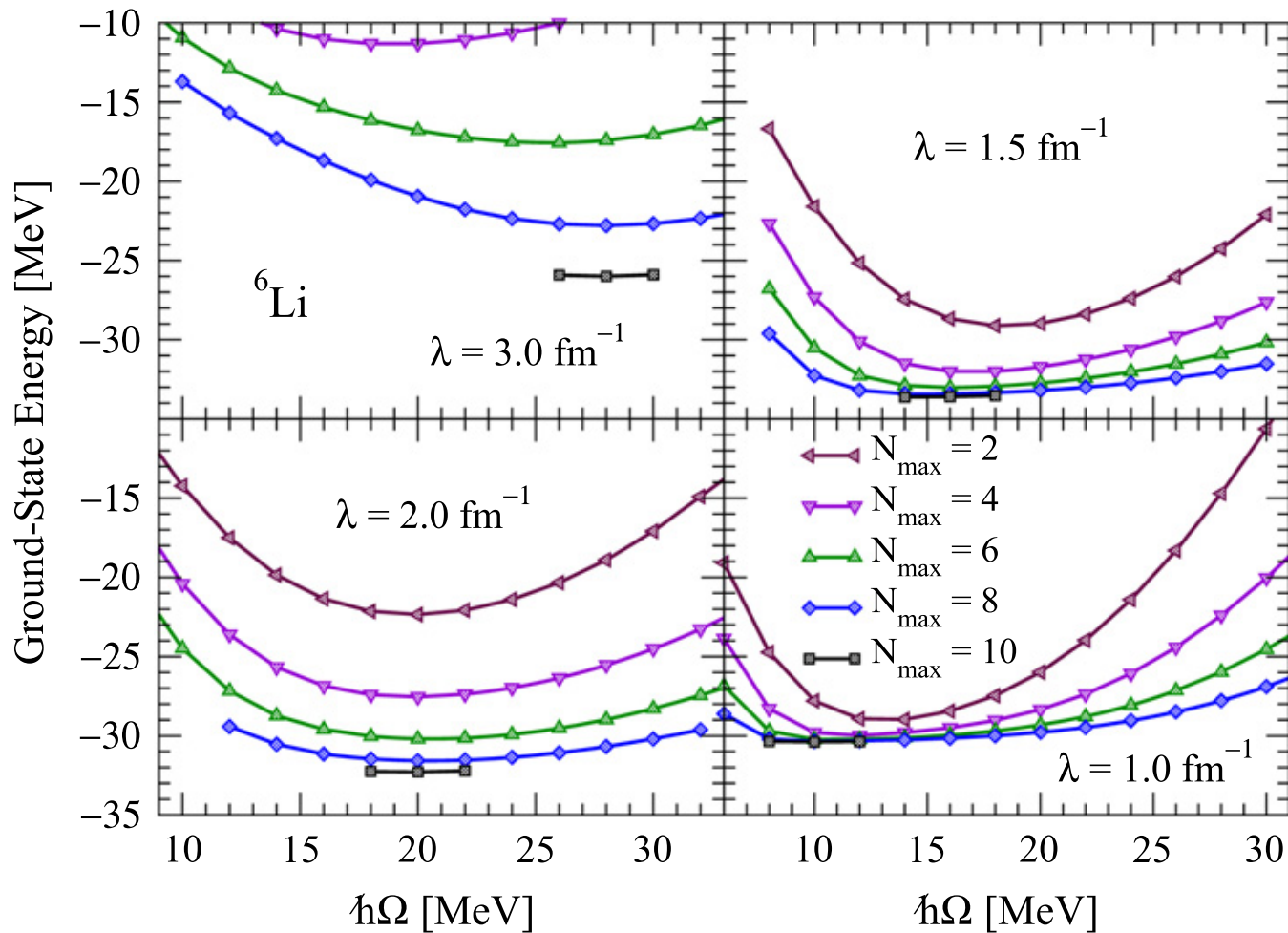
Explicitly see why this causes problems later!

Benefits of Lower Cutoffs

Exactly what happens in **no-core shell model calculations**

Probably equally helpful in normal shell-model calculations?

Come back to this later...



Benefits of Lower Cutoffs

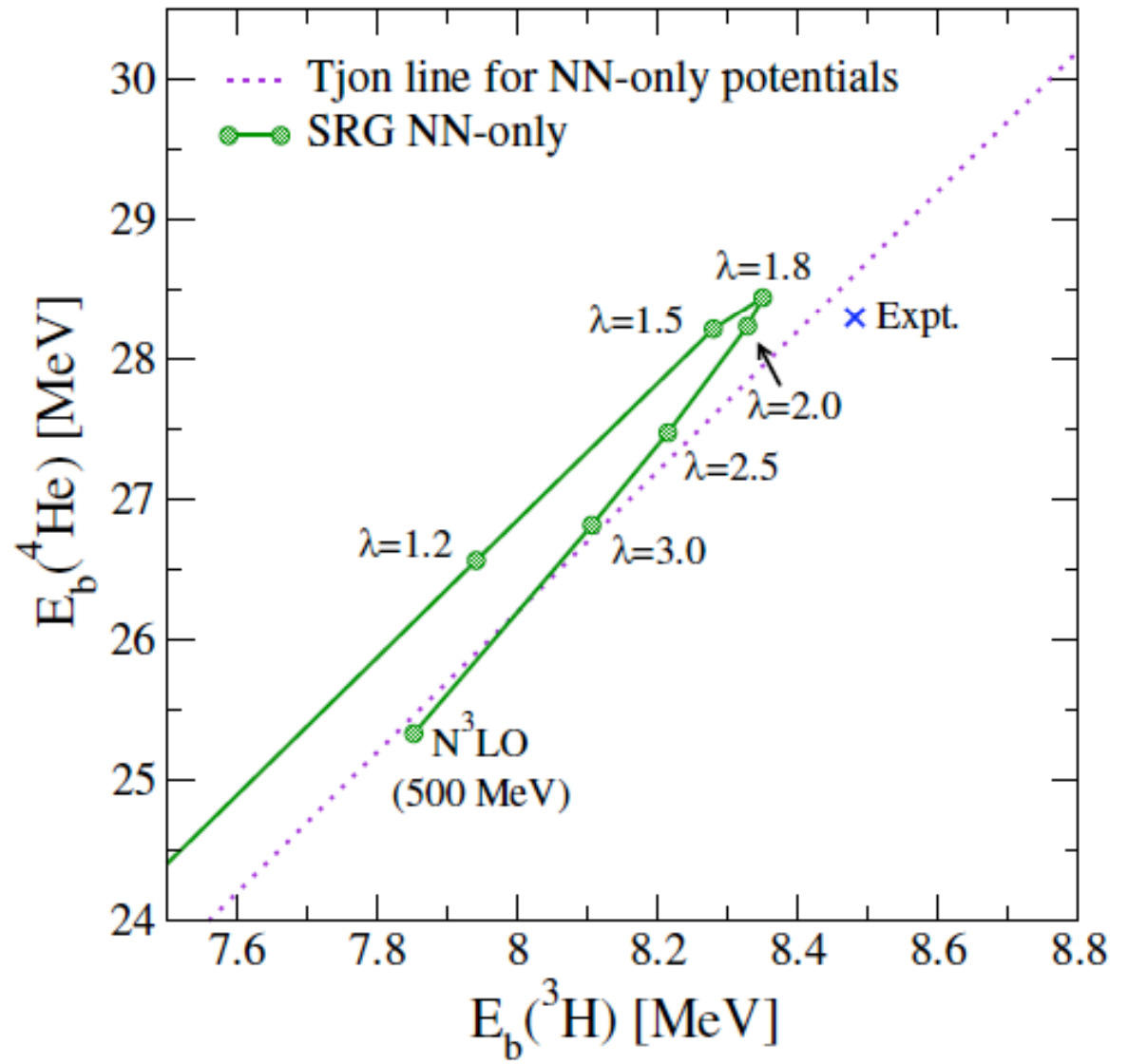
Use cutoff dependence to assess missing physics: return to Tjon line

Varying cutoff moves along line

Still never reaches experiment

Lesson: Variation in physical observables with cutoff indicates missing physics

Tool, not a parameter!

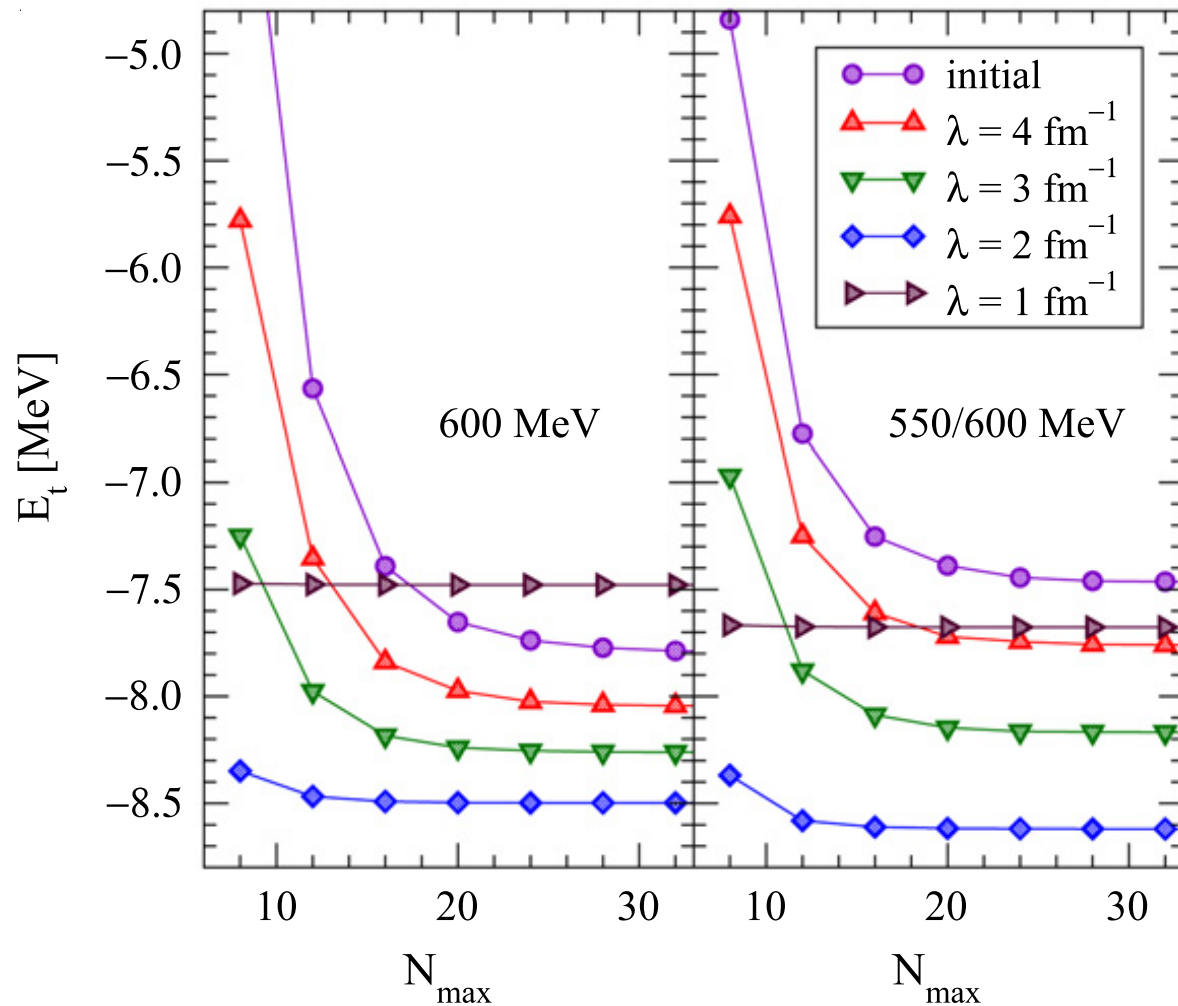


Benefits of Lower Cutoffs

Triton binding energy - again clearly improved convergence behavior

Clear dependence on cutoff – more than one, look closely...

What is the source(s)?

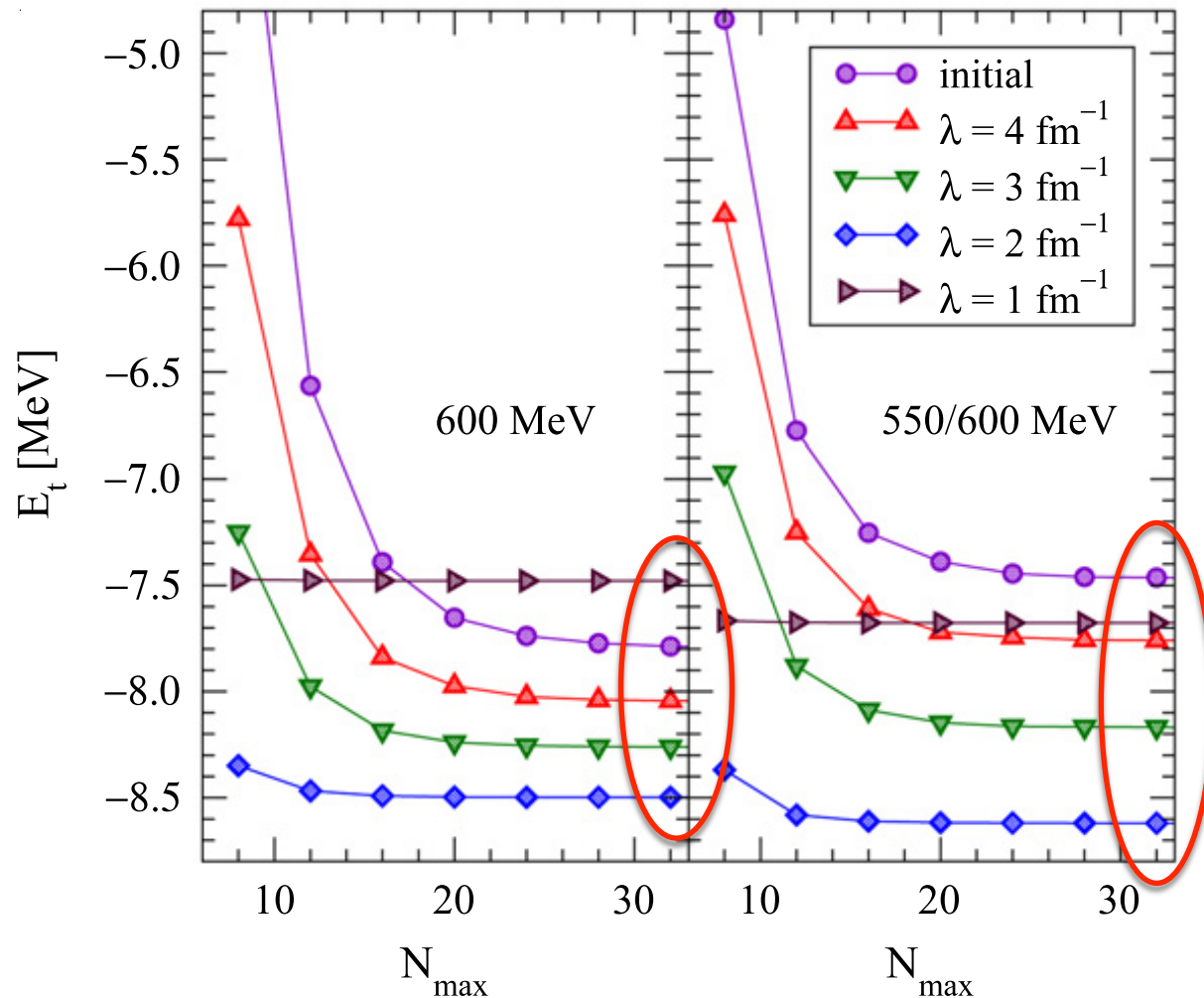


Benefits of Lower Cutoffs

Triton binding energy - again clearly improved convergence behavior

Clear dependence on cutoff – more than one, look closely...

What is the source(s)?



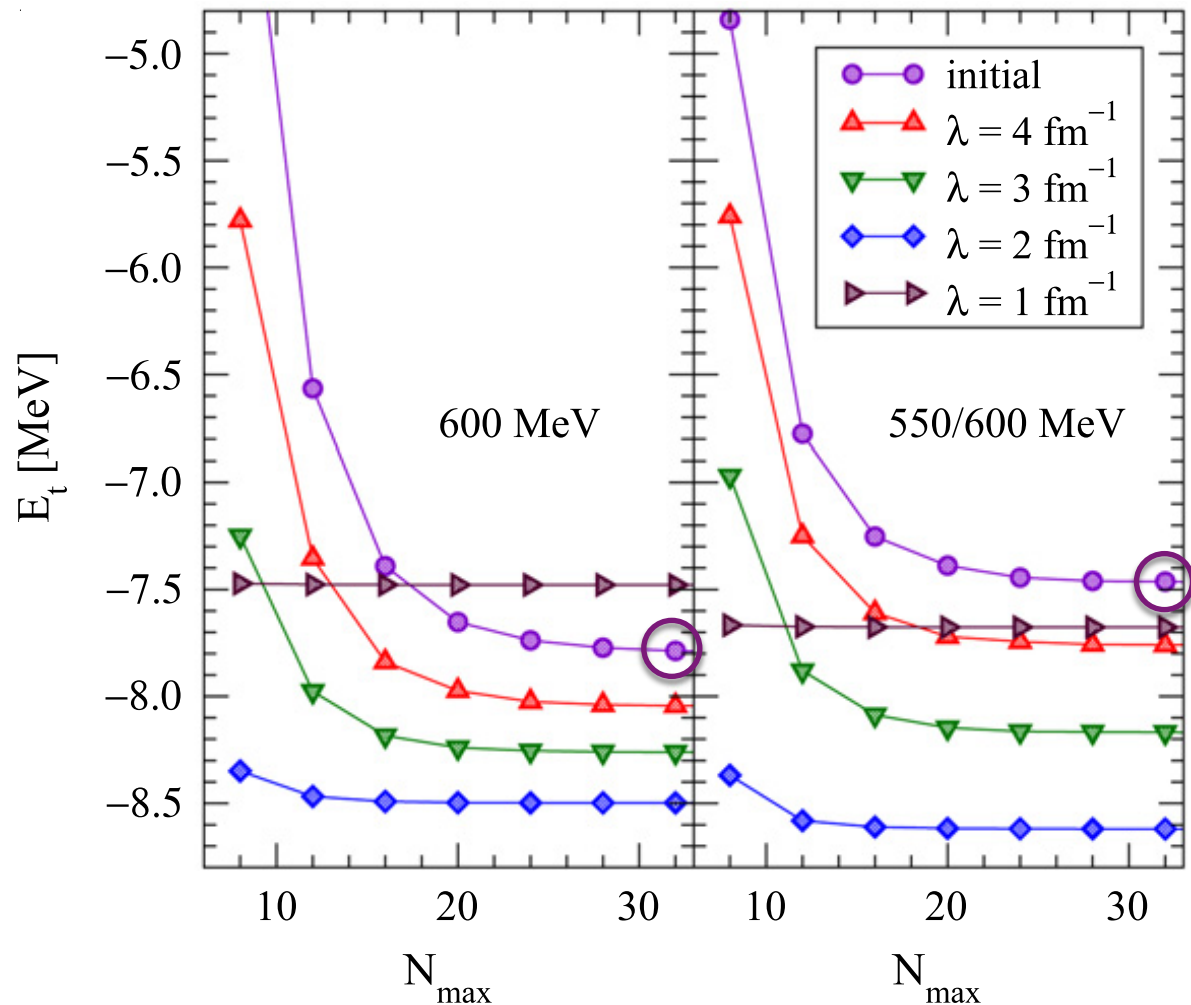
1) SRG cutoff dependence

Benefits of Lower Cutoffs

Triton binding energy - again clearly improved convergence behavior

Clear dependence on cutoff – more than one, look closely...

What is the source(s)?



- 1) SRG cutoff dependence
- 2) Initial cutoff dependence

Something missing in each case!

Case 1: Price of Low Cutoffs = Induced Forces

Life Lesson: no free lunch – not even at Summer Schools, apparently ☹️

Consider Hamiltonian with only two-body forces:

$$H = T + V_{\text{NN}}$$

And $\eta(s) = [T, H(s)]$

$$\frac{dH(s)}{ds} = [\eta(s), H(s)] = [[T, T + V(s)], T + V(s)]$$

Simply expand with creation/annihilation operators:

Case 1: Price of Low Cutoffs = Induced Forces

Life Lesson: no free lunch – not even at Summer Schools, apparently ☹️

Consider Hamiltonian with only two-body forces:

$$H = T + V_{\text{NN}}$$

And $\eta(s) = [T, H(s)]$

$$\frac{dH(s)}{ds} = [\eta(s), H(s)] = [[T, T + V(s)], T + V(s)]$$

Simply expand with creation/annihilation operators:

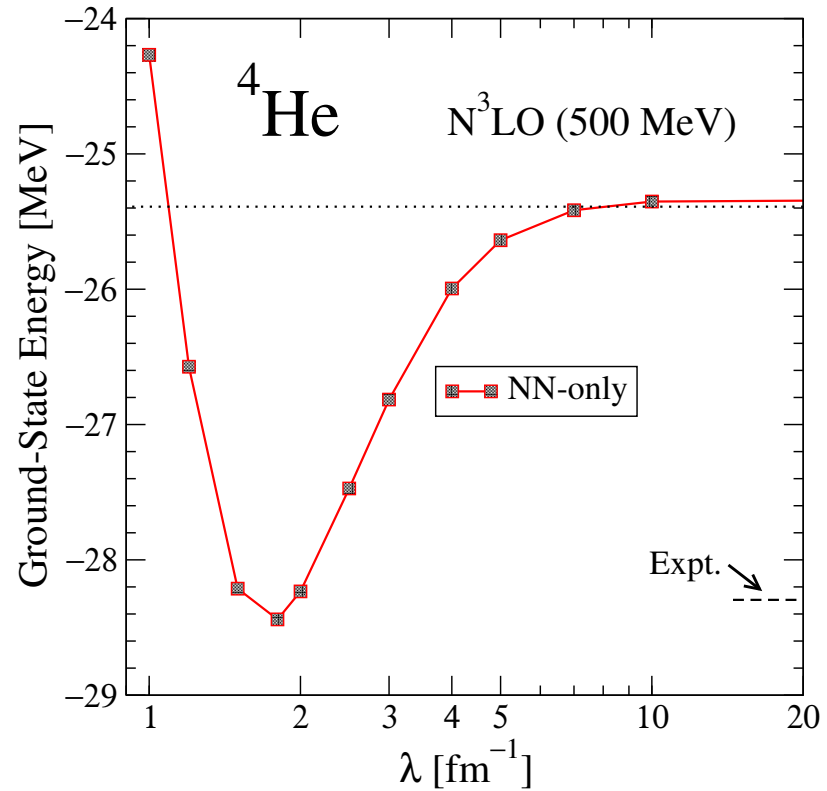
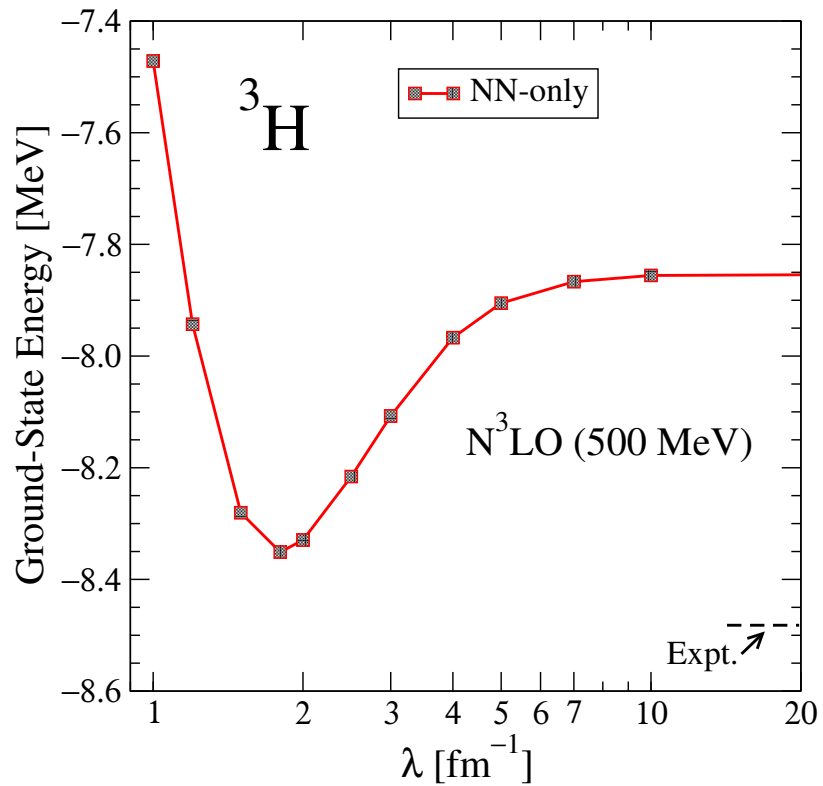
$$\frac{dV(s)}{ds} = \left[\left[\sum a^\dagger a, \sum a^\dagger a^\dagger aa \right], \sum a^\dagger a^\dagger aa \right] = \dots + \sum a^\dagger a^\dagger a^\dagger aaa + \dots$$

Three-body terms will appear even when initial 3-body forces absent

Call these **induced 3N forces (3N-ind)**

Induced 3N Forces

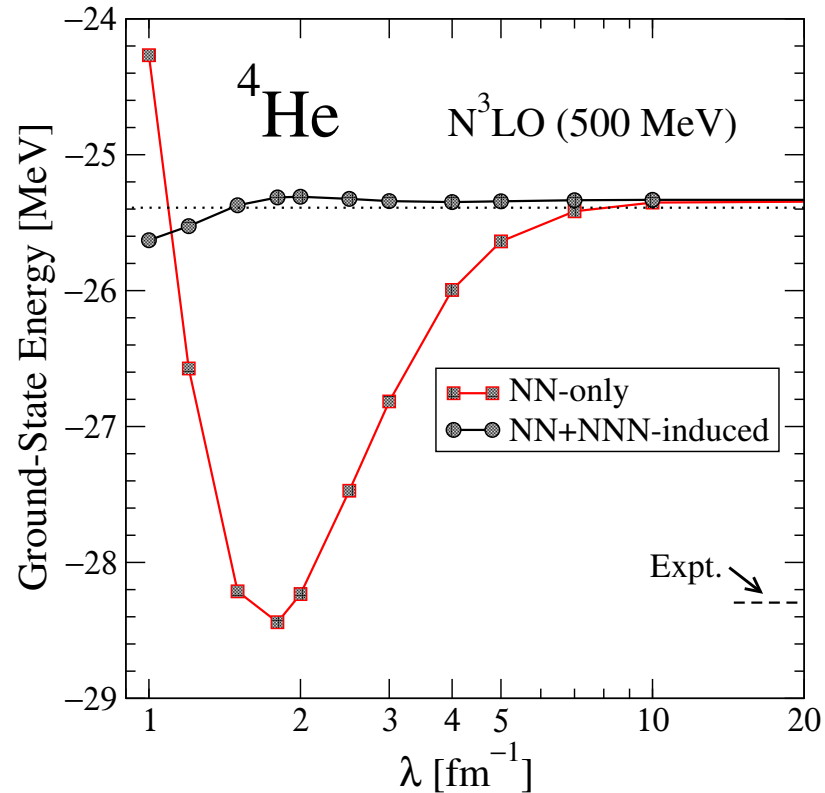
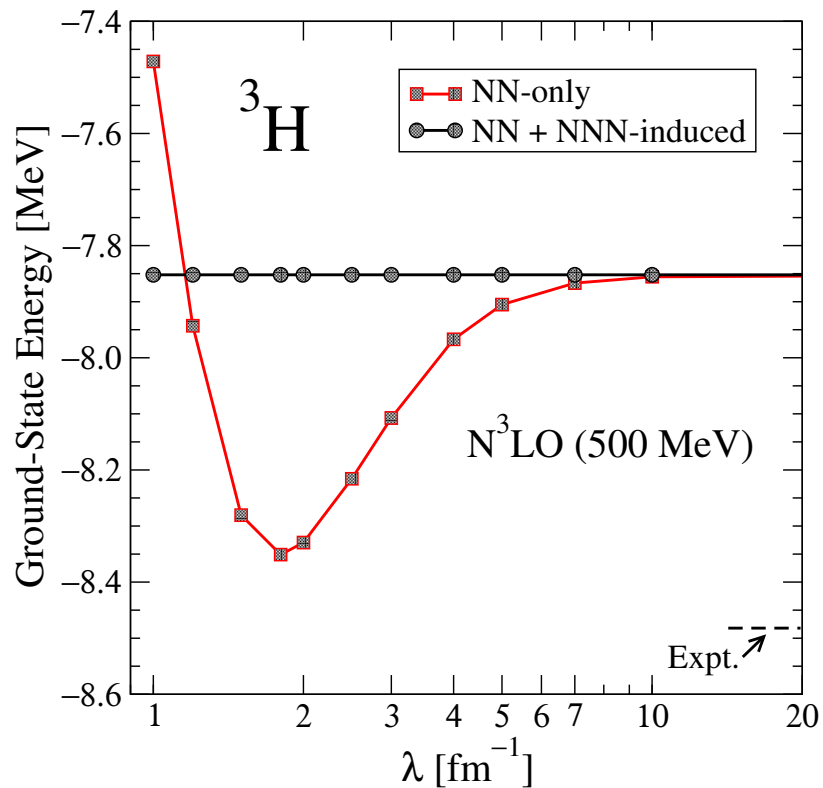
Effect of including 3N-ind? Exactly initial V_{NN} up to neglected 4N-ind



NN-only clear cutoff dependences

Induced 3N Forces

Effect of including 3N-ind? Exactly initial V_{NN} up to neglected 4N-ind



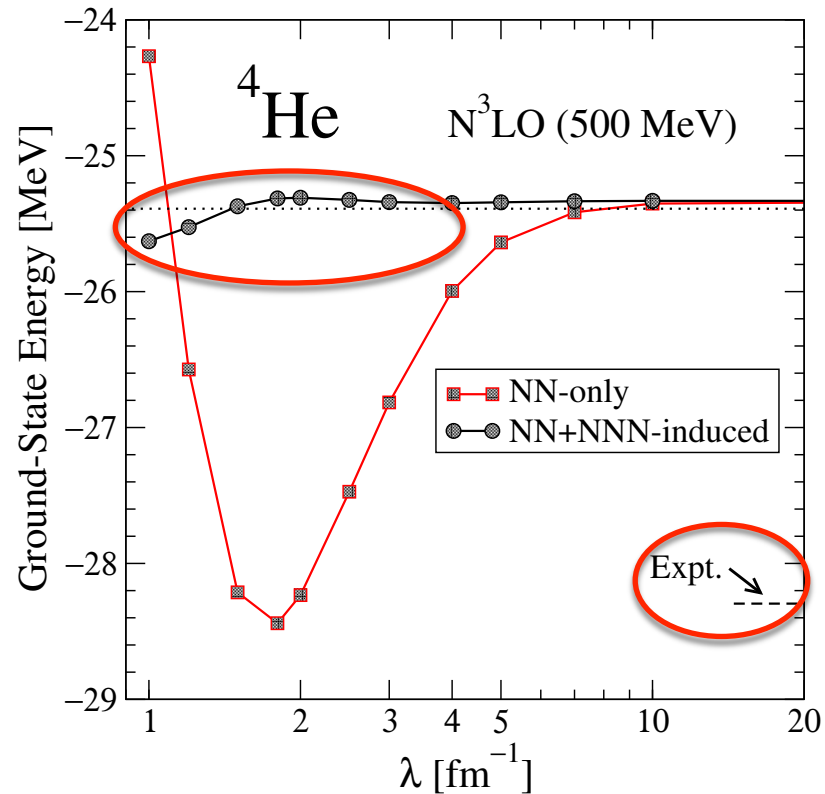
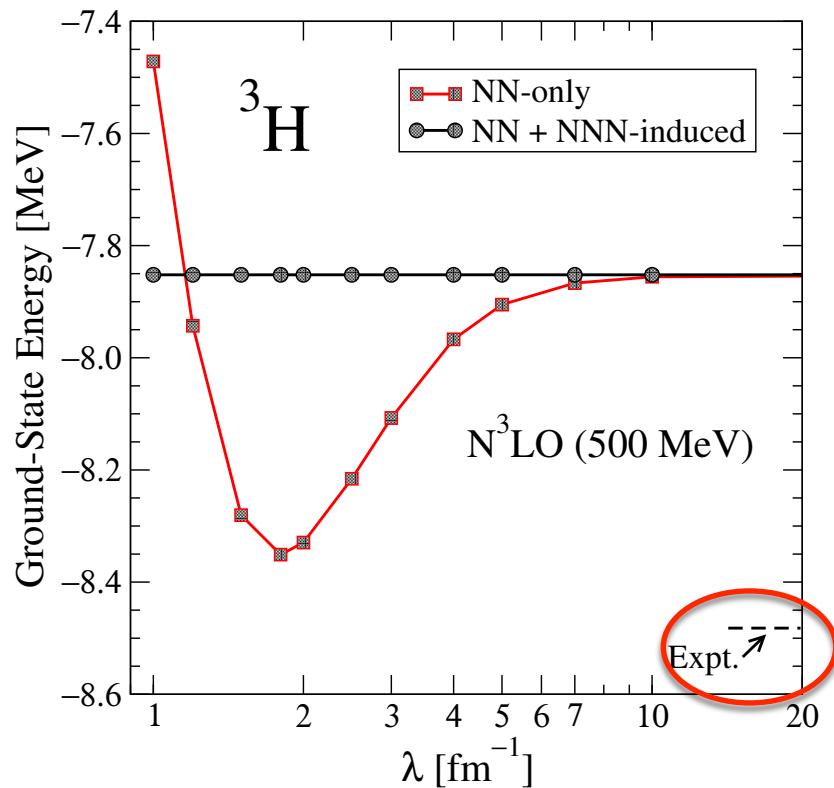
NN-only clear cutoff dependences

3N-induced – dramatic reduction in cutoff dependence!

Lesson: SRG cutoff variation a sign of neglected induced forces

Induced 3N Forces

Effect of including 3N-ind? Exactly initial V_{NN} up to neglected 4N-ind



NN-only clear cutoff dependences

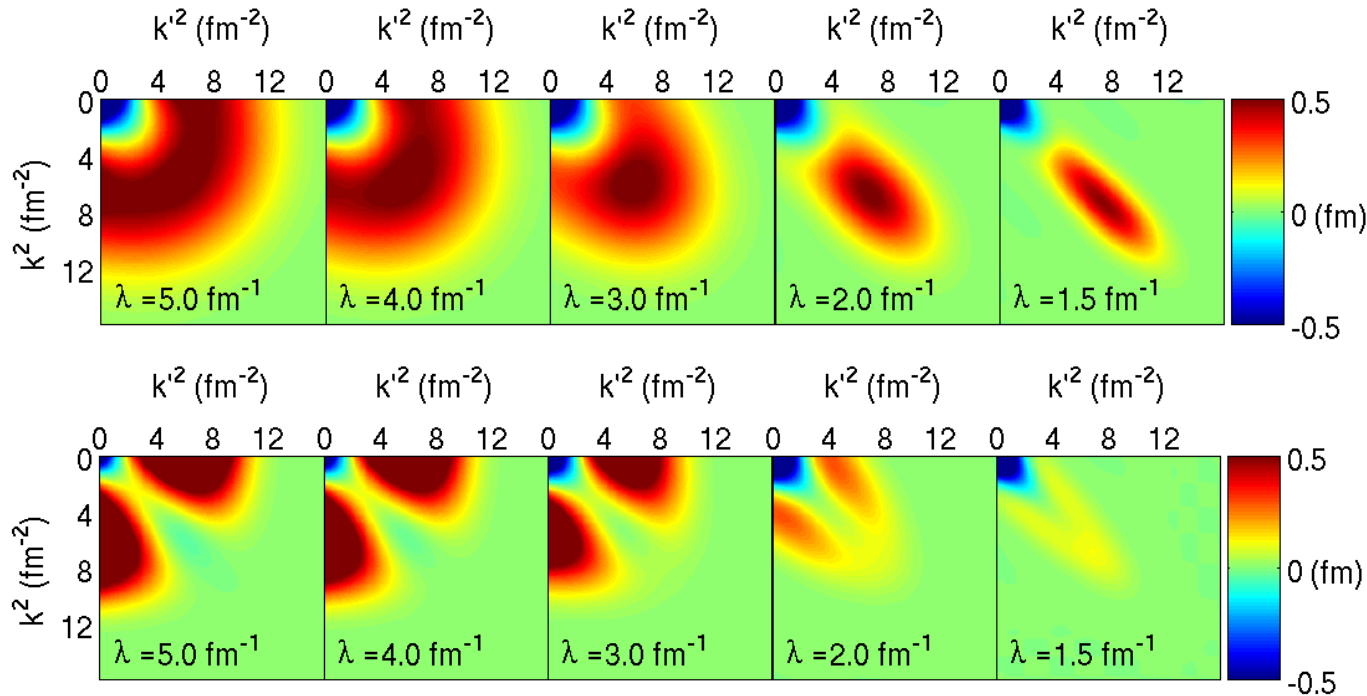
3N-induced – dramatic reduction in cutoff dependence!

Lesson: SRG cutoff variation a sign of neglected induced forces

Still far from experiment and remaining (minor) cutoff dependence!

Summary

Low-momentum interactions can be constructed from any V_{NN} via RG



Low-to-high momentum coupling not desirable in low-energy nuclear physics

Evolve to low-momentum while preserving low-energy physics

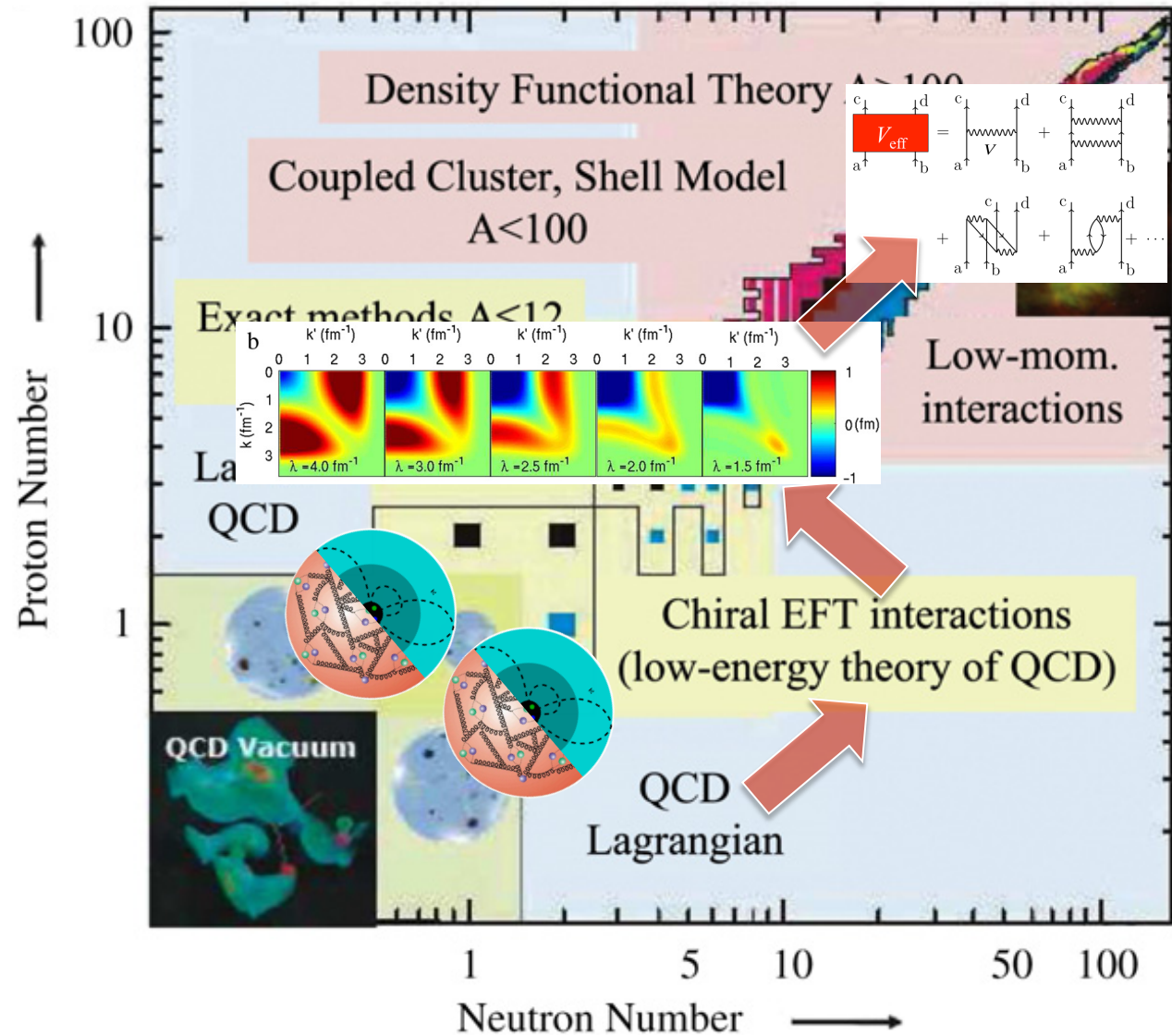
Universality attained near cutoff of data

Low-momentum cutoffs remove low-to-high harmonic oscillator couplings

Cutoff variation assesses missing physics interaction level: tool not a parameter

Part III: The Nuclear Many-Body Problem

To understand the properties of complex nuclei from first principles



Microscopic Valence-Space Interactions

Model spaces

Many-body perturbation theory (MBPT)

Calculating effective interaction

In-medium Similarity RG

Monopole part of interaction

Deficiencies of this approach

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow "Solve" many-body problem \rightarrow Predictions

The Nuclear Many-Body Problem

Nucleus strongly interacting many-body system – how to solve A -body problem?

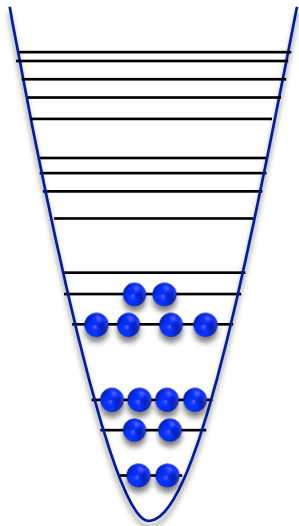
$$H\psi_n = E_n\psi_n$$

Quasi-exact solutions only in light nuclei (GFMC, NCSM...)

Large scale: controlled approximations to full Schrödinger Equation

Valence space: diagonalize exactly with reduced number of degrees of freedom

Medium-mass
Large scale

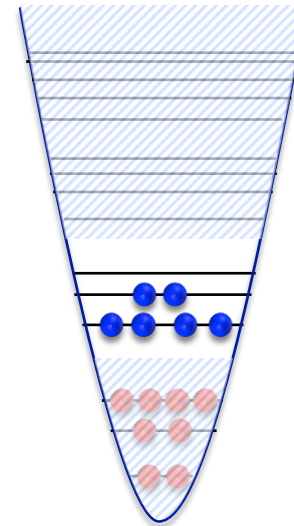


Limited range:
Closed shell ± 1
Even-even

Limited properties:
Ground states only
Some excited state

Coupled Cluster
In-Medium SRG
Green's Function

Medium-mass
Valence space



All nuclei near
closed-shell cores

All properties:
Ground states
Excited states
EW transitions

Coupled Cluster
In-Medium SRG
Perturbation Theory

From Momentum Space to HO Basis

To this point interaction matrix elements in momentum space, **partial waves**

$$\langle kK, lL | V | k'K, l'L \rangle_{\alpha}$$

To go to finite nuclei begin from Hamiltonian

$$H\psi_n = (T + V)\psi_n = E_n\psi_n$$

Assume many particles in the nucleus generate a **mean field** U :

U a one-body potential simple to solve (typically **Harmonic Oscillator**)

$$H = H_0 + H_1; \quad H_0 = T + U; \quad H_1 = V - U$$

So transform from momentum space to **Harmonic Oscillator Basis**

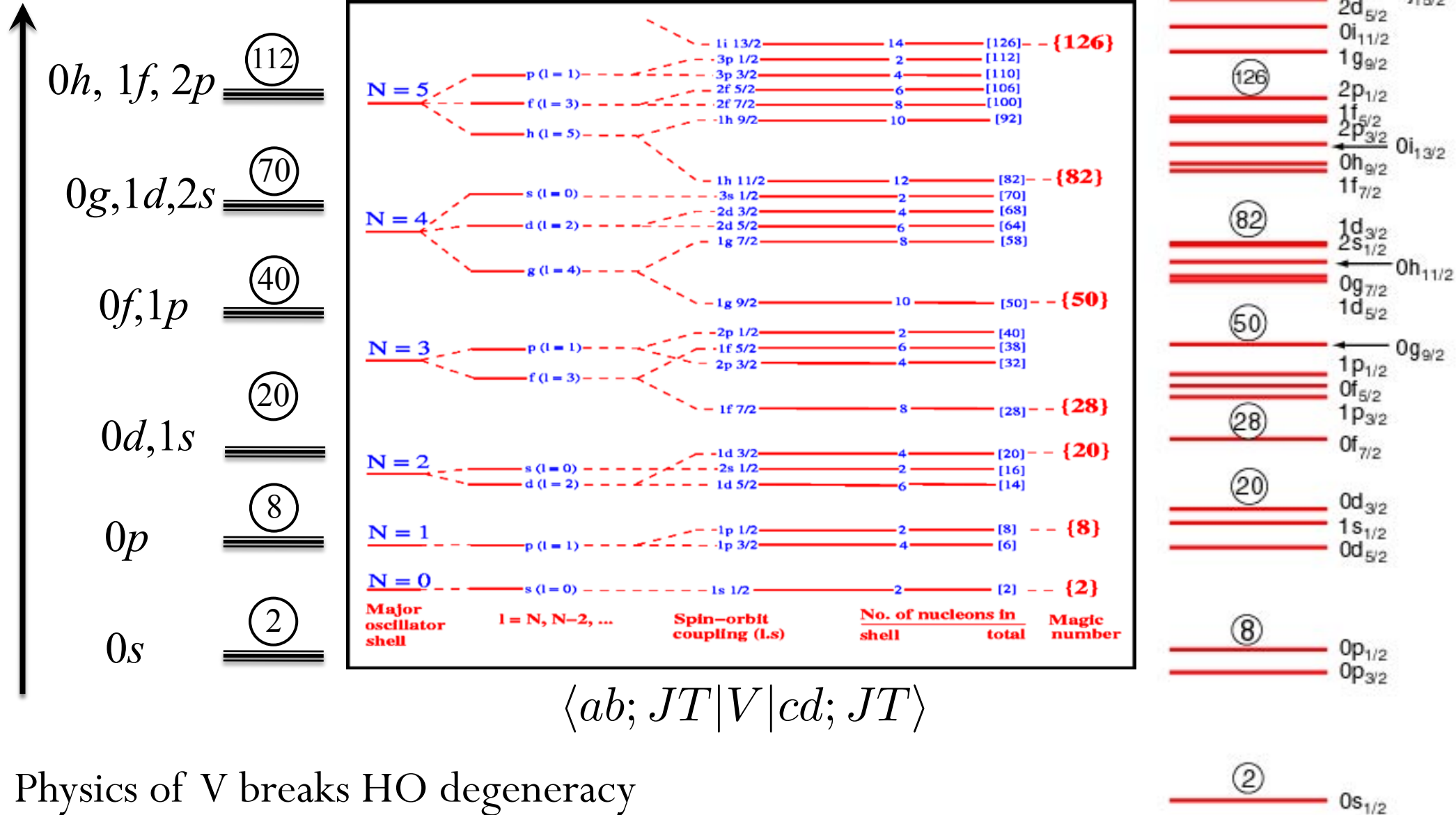
$$|nl, NL; \alpha\rangle = \int k^2 dk K^2 dK R_{nl}(\sqrt{2}\alpha k) R_{NL}(\sqrt{1/2}\alpha K) |kl, KL; \alpha\rangle$$

One more (ugly) transformation from center-of-mass to lab frame:

$$\rightarrow \langle ab; JT | V | cd; JT \rangle$$

Valence-Space Ideas

Begin with degenerate HO levels



$$\langle ab; JT | V | cd; JT \rangle$$

Physics of V breaks HO degeneracy

Problem: Can't solve Schrodinger equation in full Hilbert space

Valence-Space Ideas

Nuclei understood as many-body system starting from closed shell, add nucleons

Unperturbed

HO spectrum

$0h, 1f, 2p$ (112)

$0g, 1d, 2s$ (70)

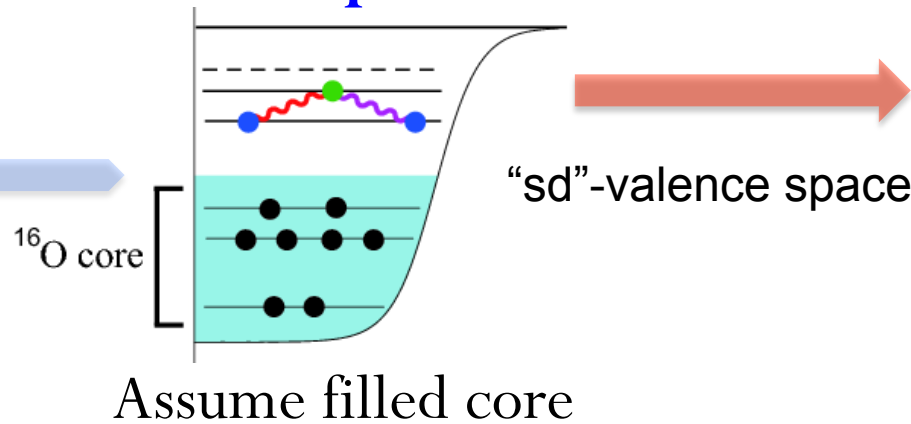
$0f, 1p$ (40)

$0d, 1s$ (20)

$0p$ (8)

$0s$ (2)

Active nucleons occupy
valence space



Removes degeneracy in
valence space only

$0h, 1f, 2p$ (112)

$0g, 1d, 2s$ (70)

$0f, 1p$ (40)

(20)
 $0d_{3/2}$
 $1s_{1/2}$
 $0d_{5/2}$

$0p$ (8)

$0s$ (2)

Valence-Space Ideas

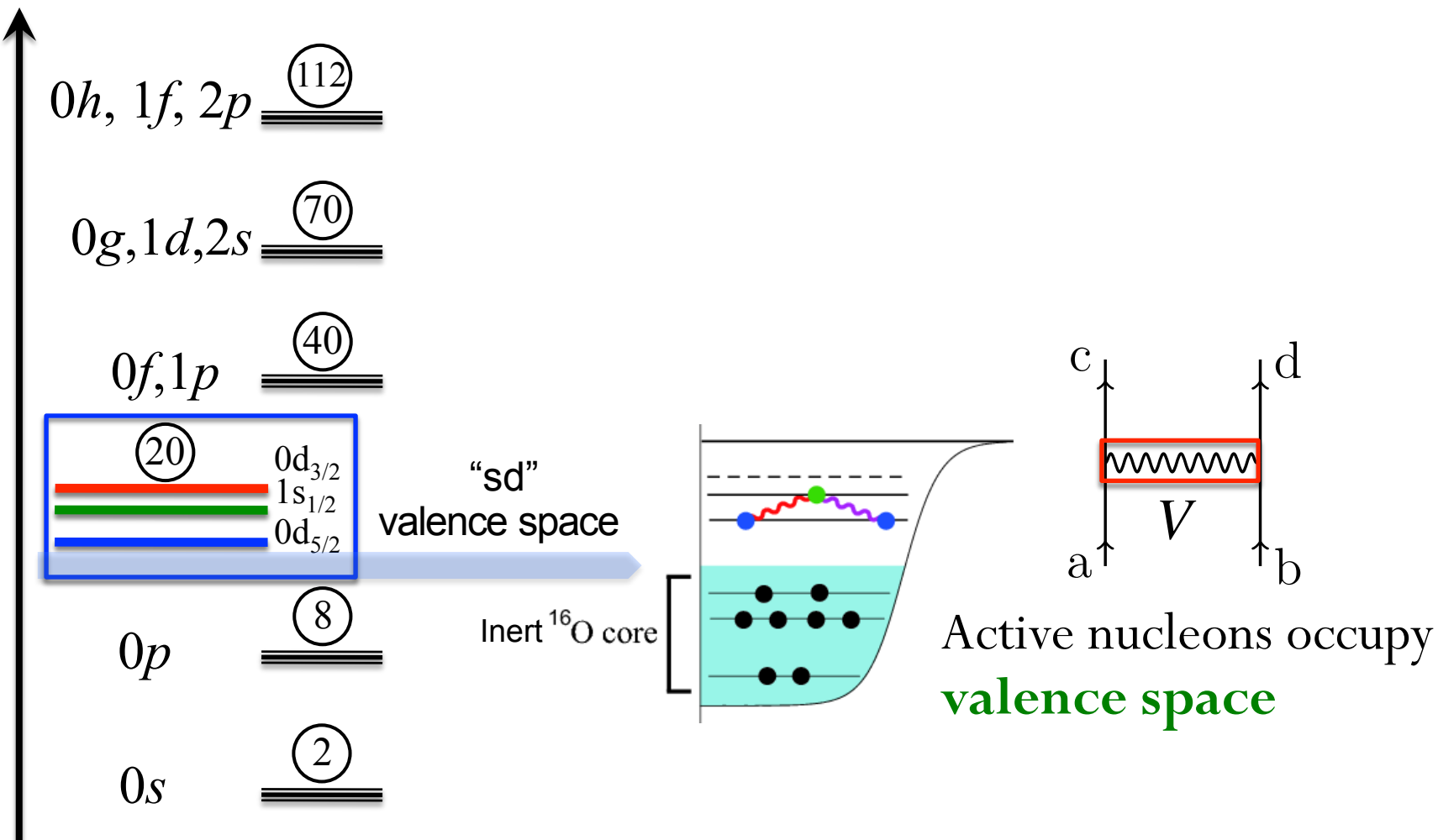
Nuclei understood as many-body system starting from closed shell, add nucleons

Valence-space Hamiltonian derived from nuclear forces:

Single-particle energies

$$H_{\text{v.s.}} = \sum_i \epsilon_i a_i^\dagger a_i + V_{\text{v.s.}}$$

Interaction matrix elements



Valence-Space Philosophy

Nuclei understood as many-body system starting from closed shell, add nucleons

Valence-space Hamiltonian derived from nuclear forces:

Single-particle energies

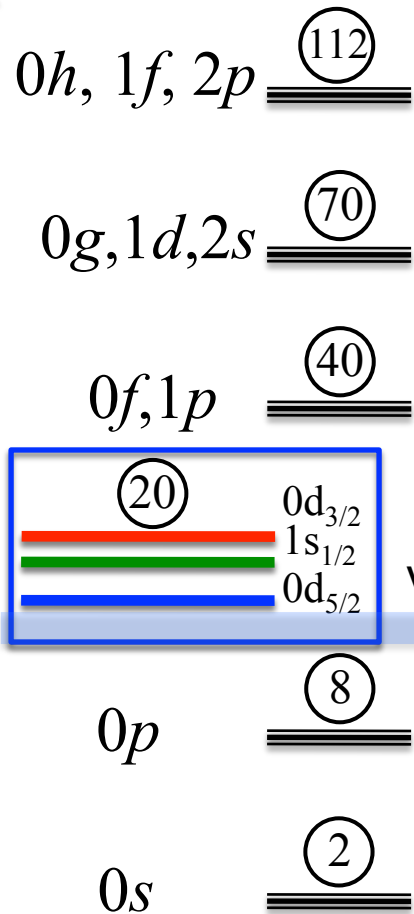
$$H_{\text{eff}} = \sum_i \epsilon_{i_{\text{eff}}} a_i^\dagger a_i + V_{\text{eff}}$$

Interaction matrix elements

$$H\psi_n = E_n\psi_n \rightarrow PH_{\text{eff}}P\psi_i = E_iP\psi_i$$

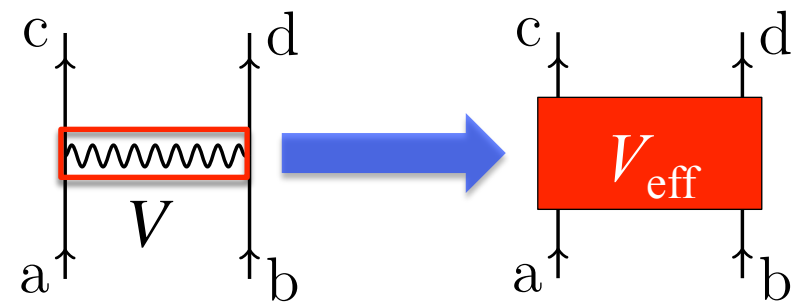
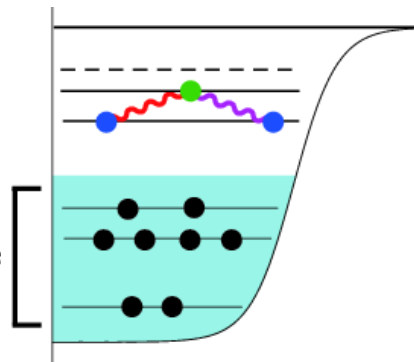
Effective valence space Hamiltonian:

Sum all excitations outside valence space



“sd”
valence space

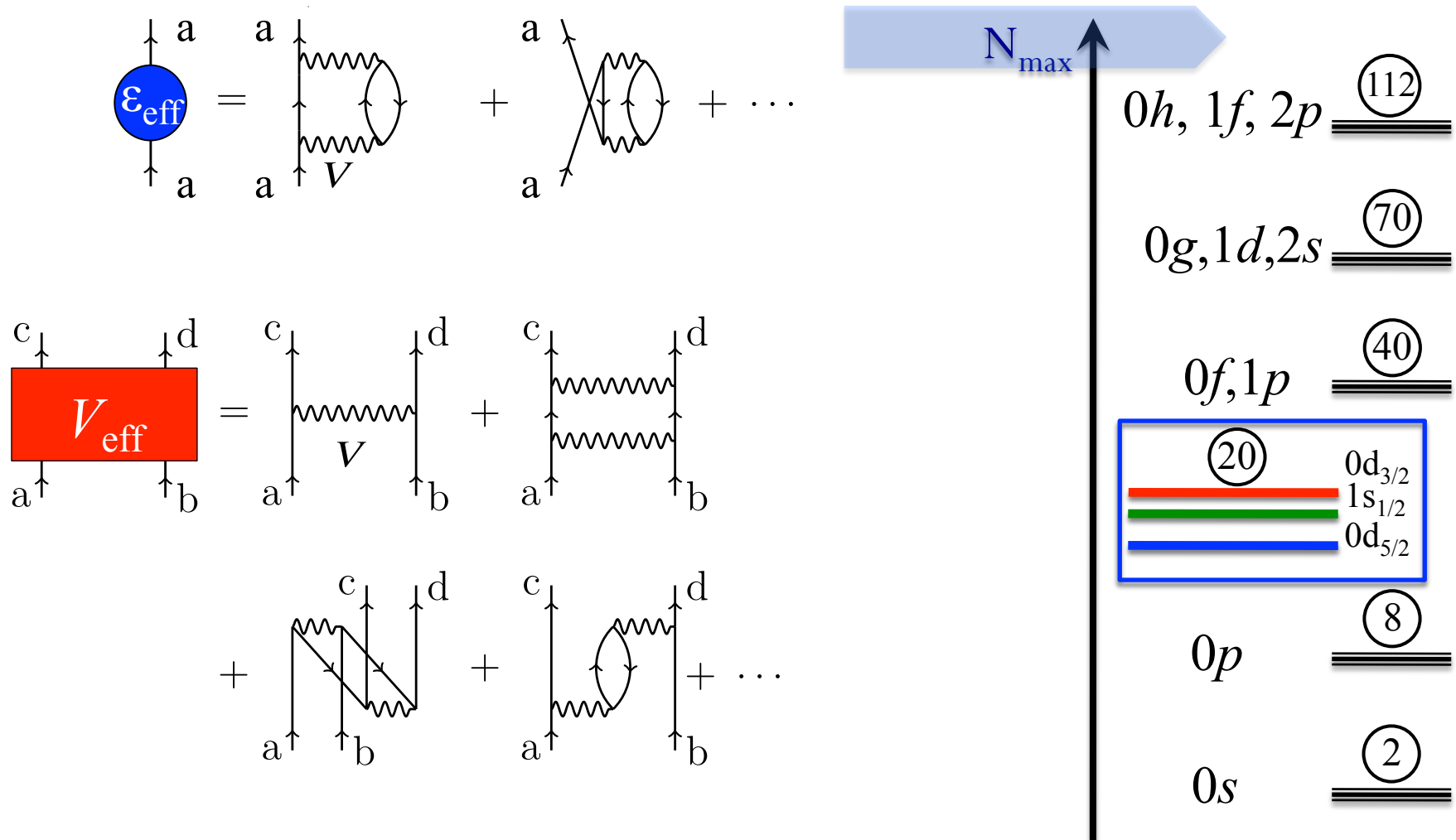
Inert ^{16}O core



Decouple valence space
from excitations

Perturbative Approach

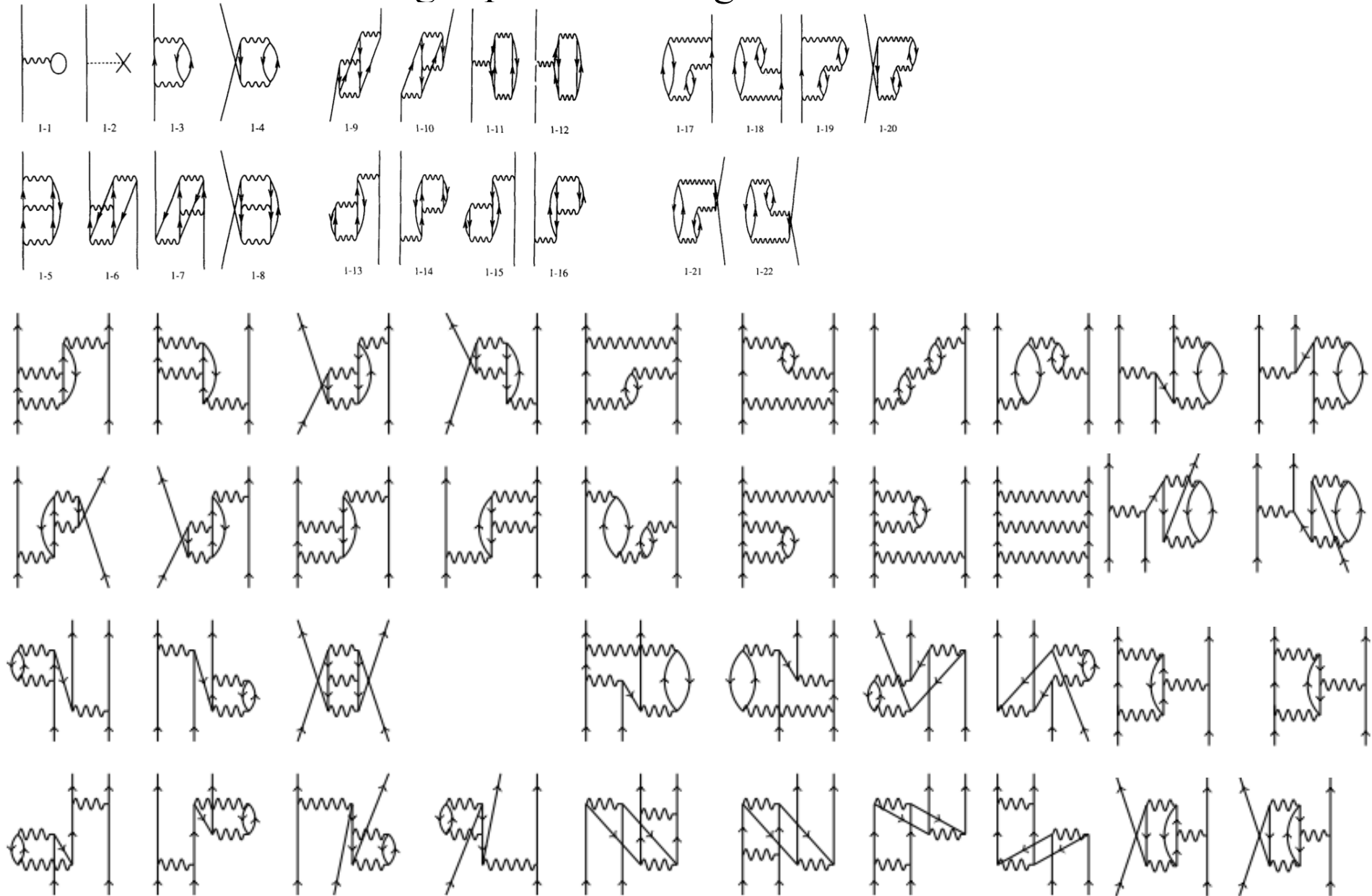
- 1) Effective Hamiltonian: sum excitations outside valence space
- 2) Self-consistent single-particle energies



Perturbative Approach

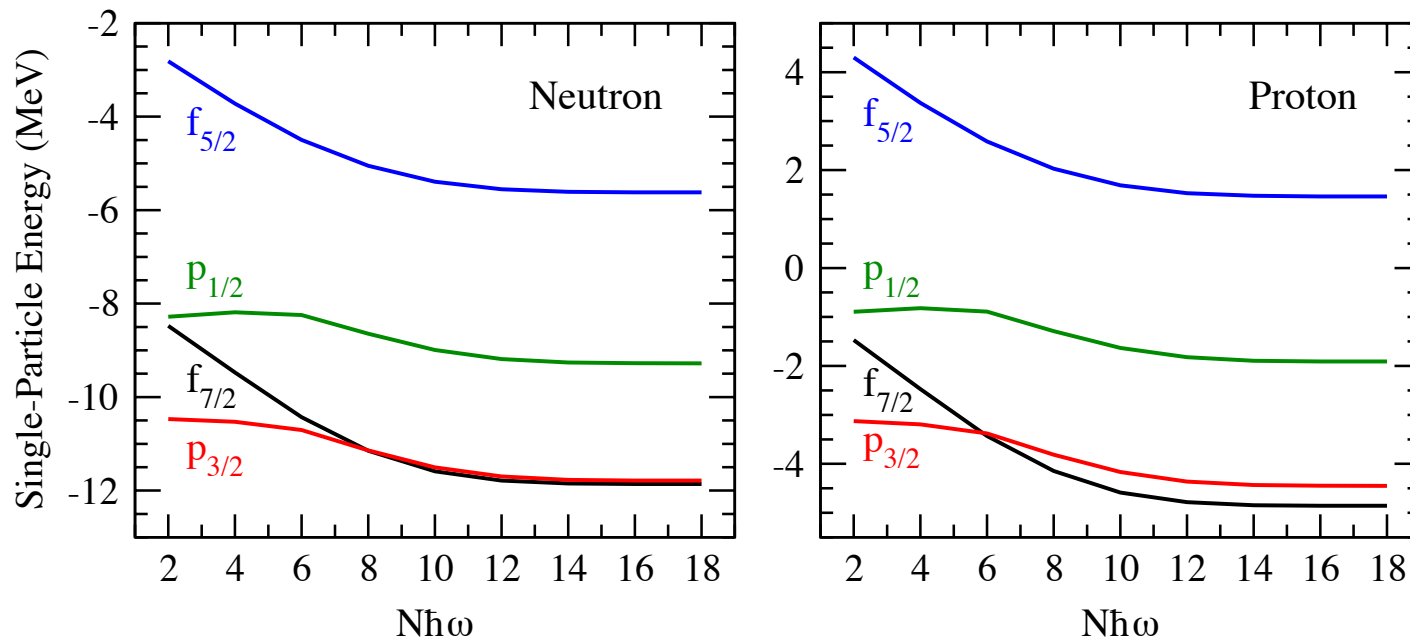
1) Effective Hamiltonian: sum excitations outside valence space to **MBPT(3)**

2) Self-consistent single-particle energies



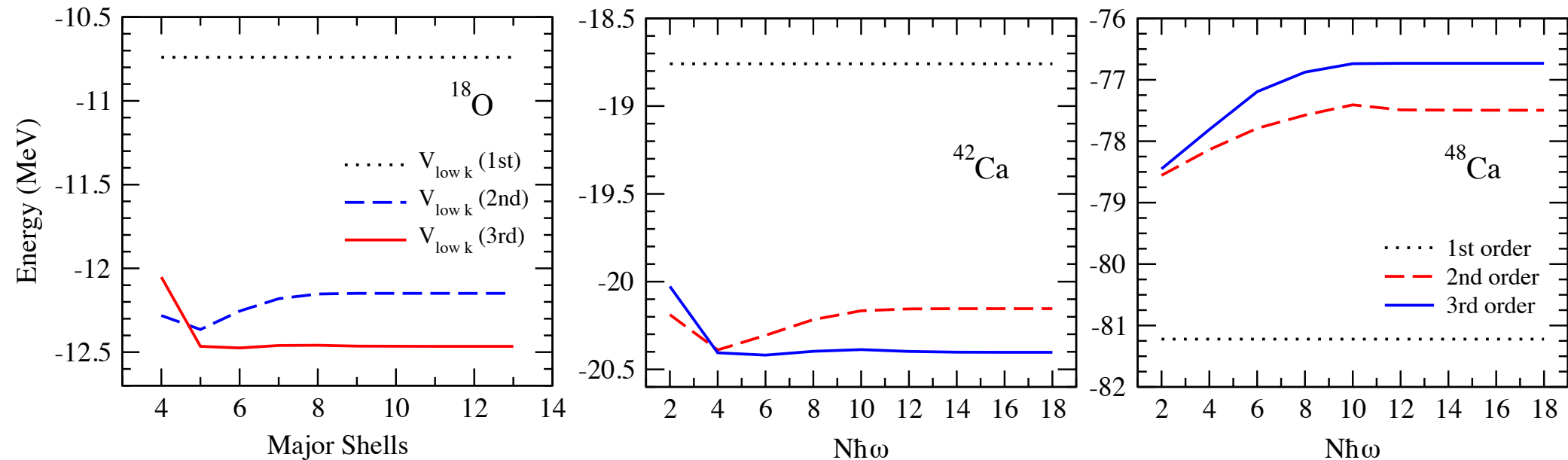
Perturbative Approach

- 1) Effective Hamiltonian: sum excitations outside valence space to MBPT(3)
- 2) Self-consistent single-particle energies
- 3) **Harmonic-oscillator basis** of 13-15 major shells: **converged!**



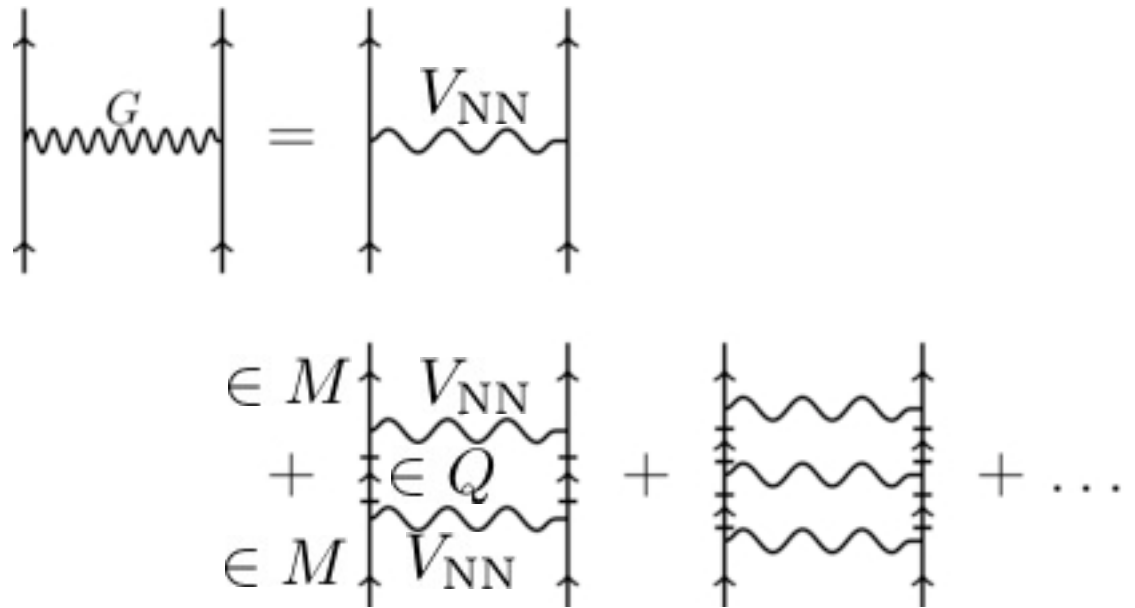
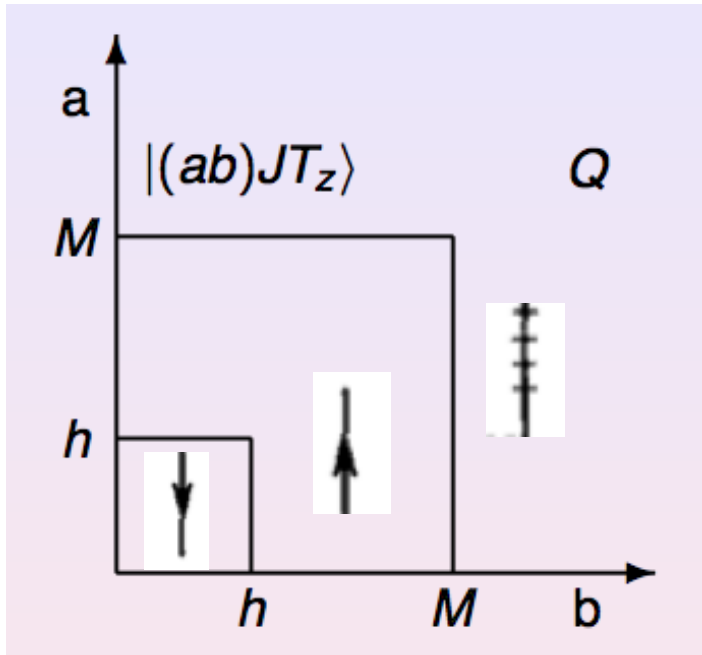
Perturbative Approach

- 1) Effective Hamiltonian: sum excitations outside valence space to MBPT(3)
- 2) Self-consistent single-particle energies
- 3) **Harmonic-oscillator basis** of 13-15 major shells: **converged!**



Aside: G-matrix Renormalization

Standard method for softening interaction in nuclear structure for decades:



Infinite summation of ladder diagrams

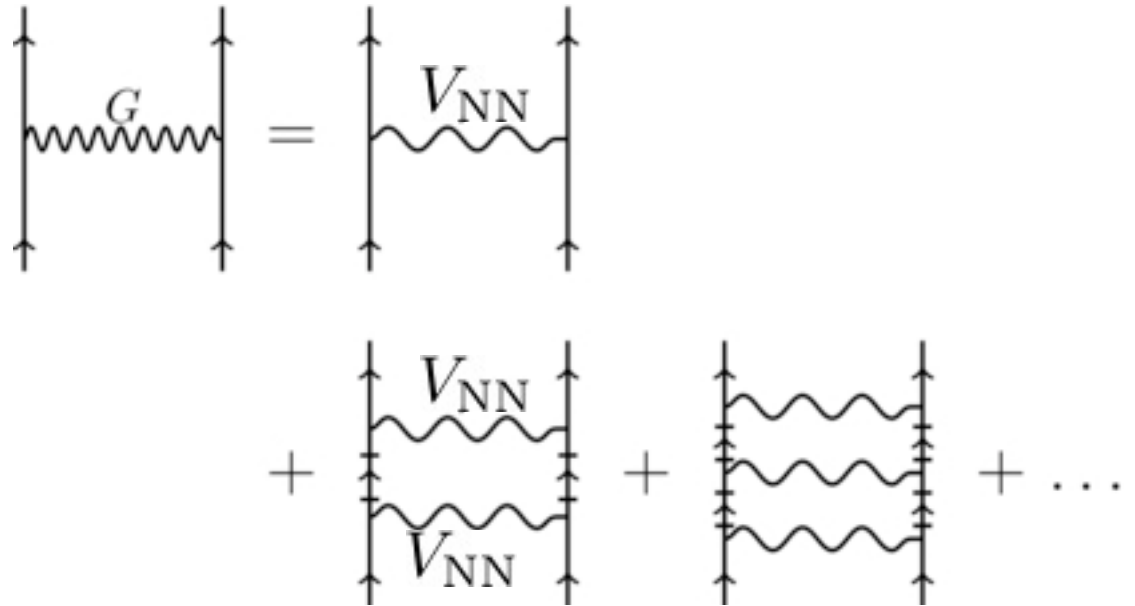
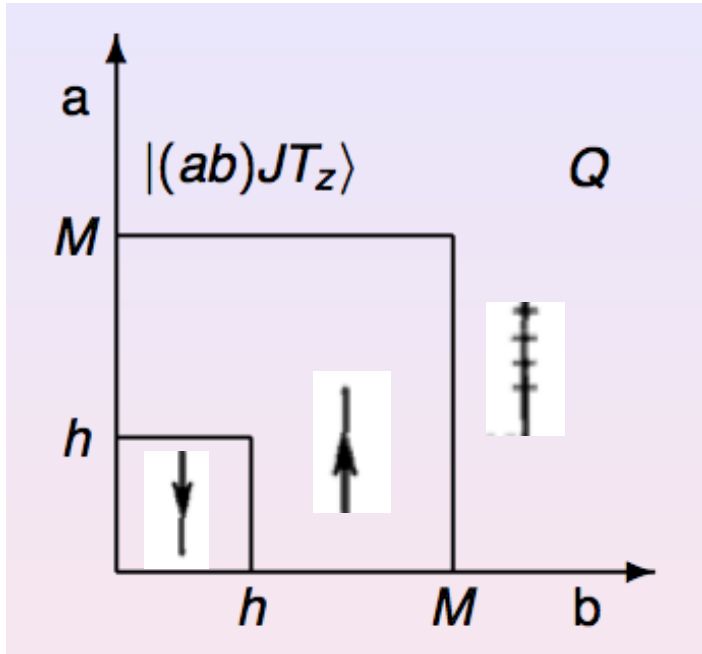
Need two model spaces:

- 1) \mathbf{M} space in which we will want to calculate (excitations allowed in \mathbf{M})
- 2) Large space \mathbf{Q} in which particle excitations are allowed

To avoid double counting, can't overlap – **matrix elements depend on \mathbf{M}**

Aside: G-matrix Renormalization

Standard method for softening interaction in nuclear structure for decades:



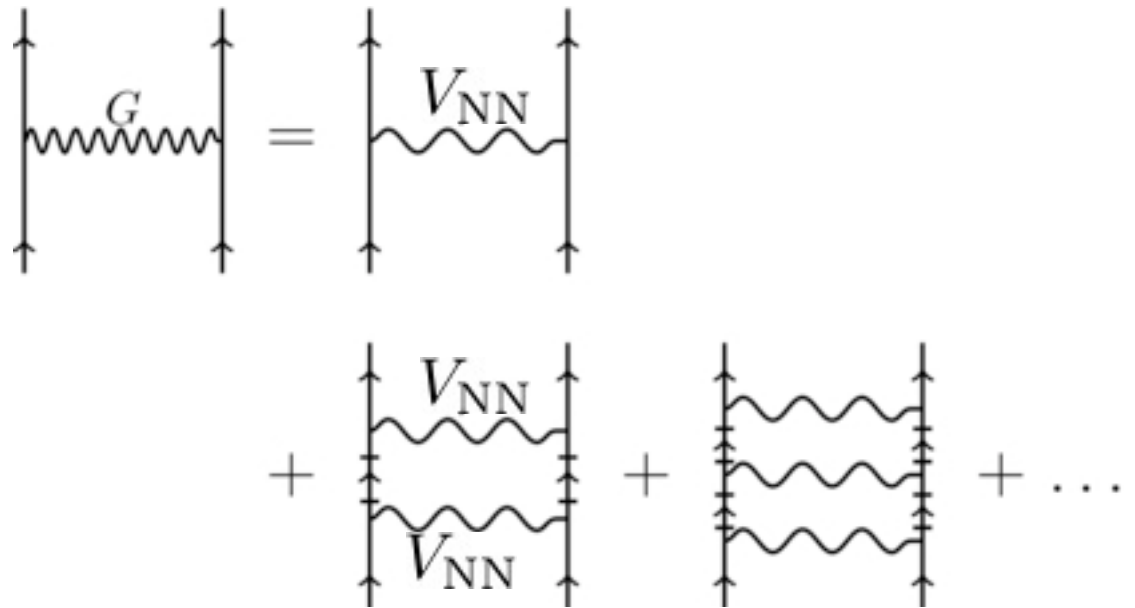
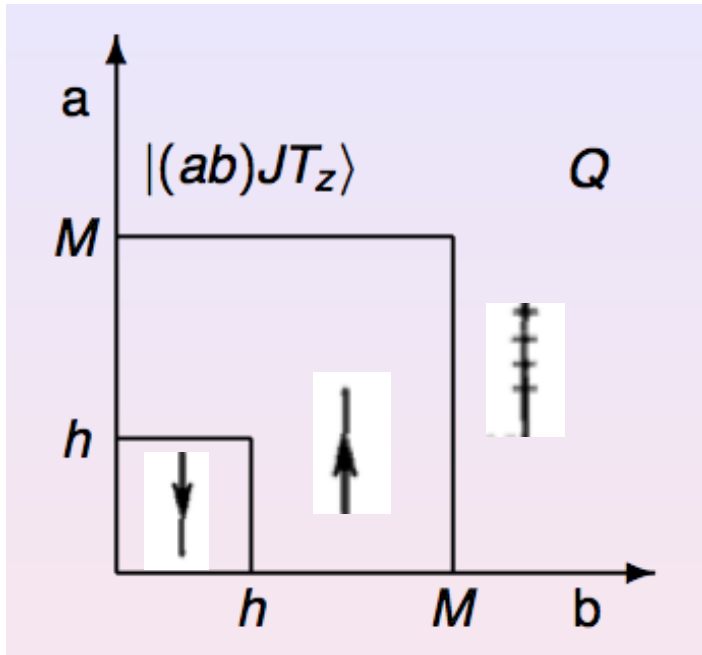
$$G_{ijkl}(\omega) = V_{ijkl} + \sum_{mn \in Q} V_{ijmn} \frac{Q}{\omega - \epsilon_m - \epsilon_n} G_{mnkl}(\omega)$$

Iterative procedure

Dependence on arbitrary starting energy!

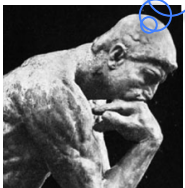
G-matrix Renormalization

Standard method for softening interaction in nuclear structure for decades:



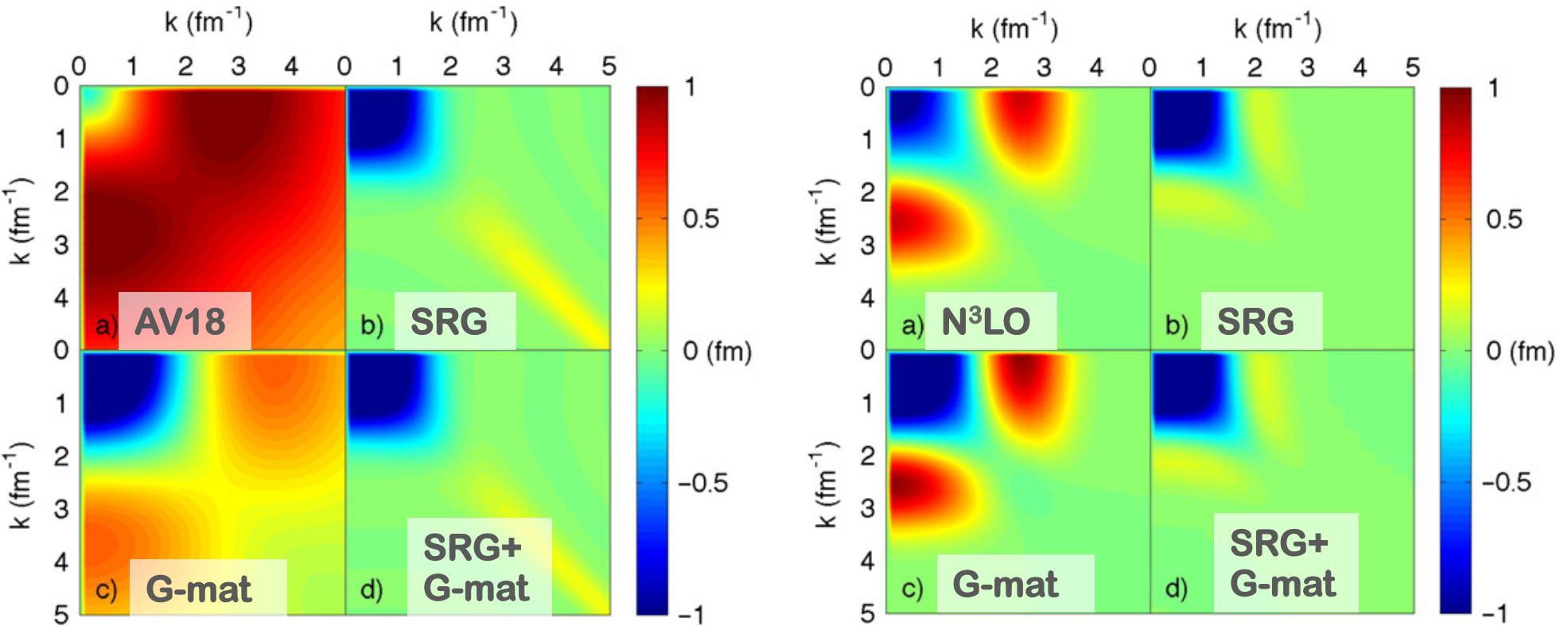
$$G_{ijkl}(\omega) = V_{ijkl} + \sum_{mn \in Q} V_{ijmn} \frac{Q}{\omega - \varepsilon_m - \varepsilon_n} G_{mnkl}(\omega)$$

What happens
as we keep
increasing M?



G-matrix Renormalization

Results of **G-matrix** renormalization vs. SRG



Removes some diagonal high-momentum components

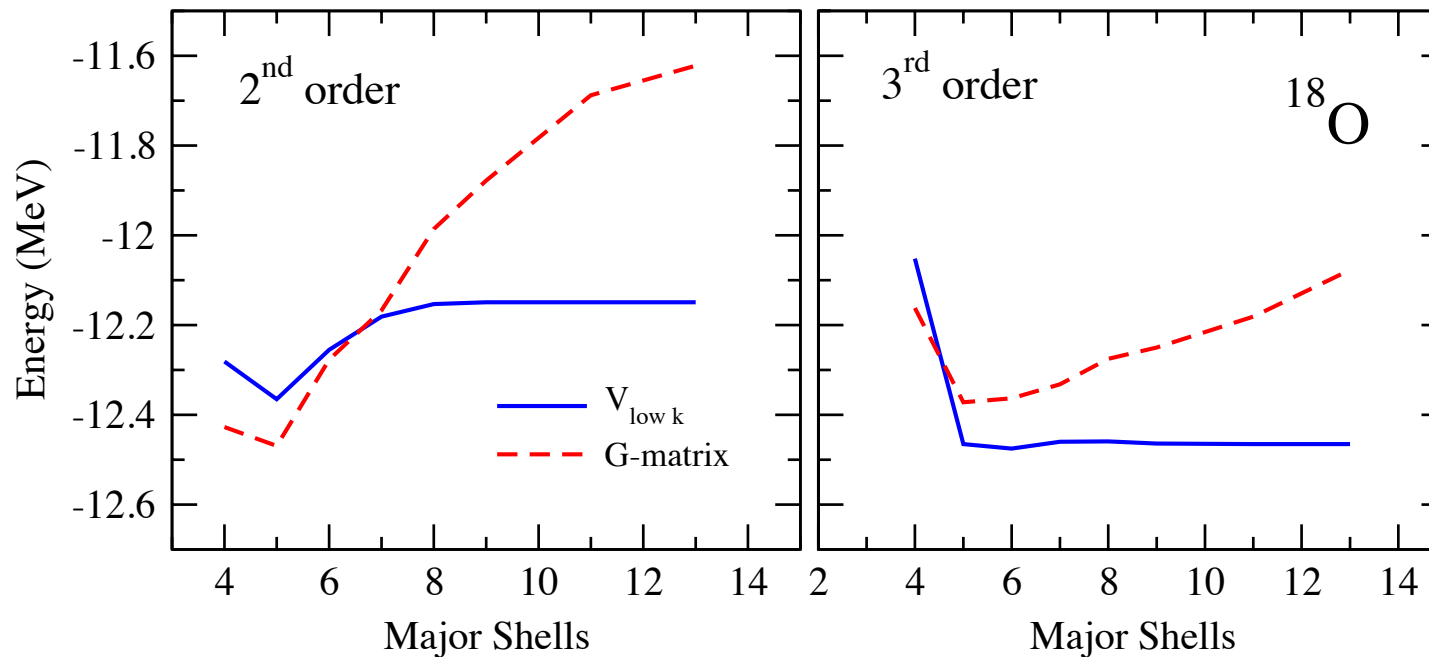
Still large low-to-high coupling in both interactions

No indication of universality

Clear difference compared with SRG-evolved interactions!

Perturbative Approach

- 1) Effective Hamiltonian: sum excitations outside valence space to MBPT(3)
- 2) Self-consistent single-particle energies
- 3) **Harmonic-oscillator basis** of 13-15 major shells: **converged!**

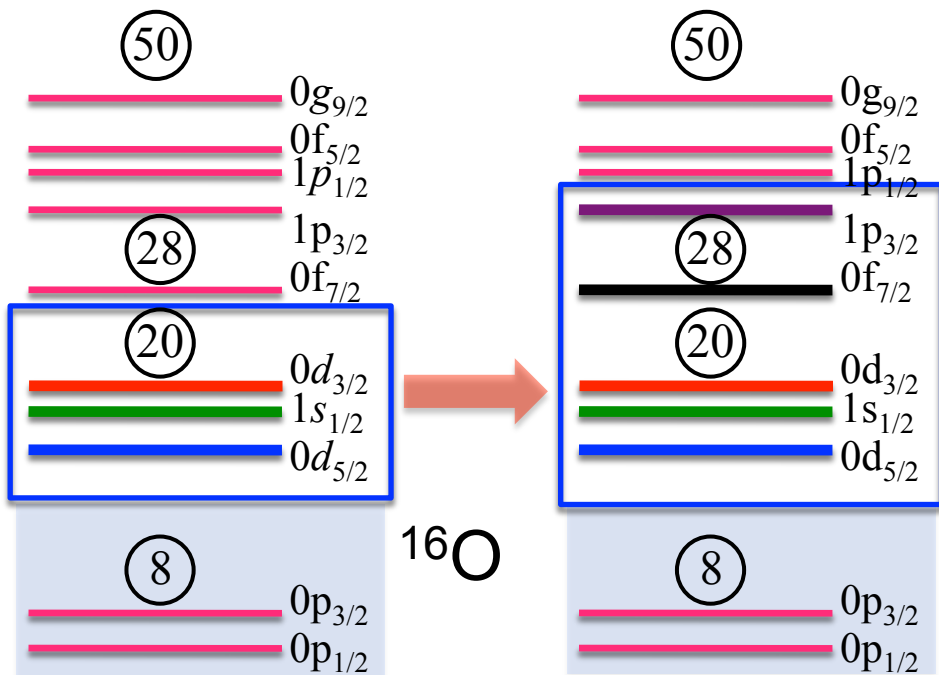


Compare vs G-matrix (no sign of convergence)

Clear benefit of low-momentum interactions!

Perturbative Approach

- 1) Effective Hamiltonian: sum excitations outside valence space to MBPT(3)
- 2) Self-consistent single-particle energies
- 3) Harmonic-oscillator basis of 13-15 major shells
- 4) Nuclear forces from chiral EFT
- 5) Requires **extended valence spaces**



Treat higher orbits nonperturbatively

Limits of Nuclear Existence: Oxygen Anomaly

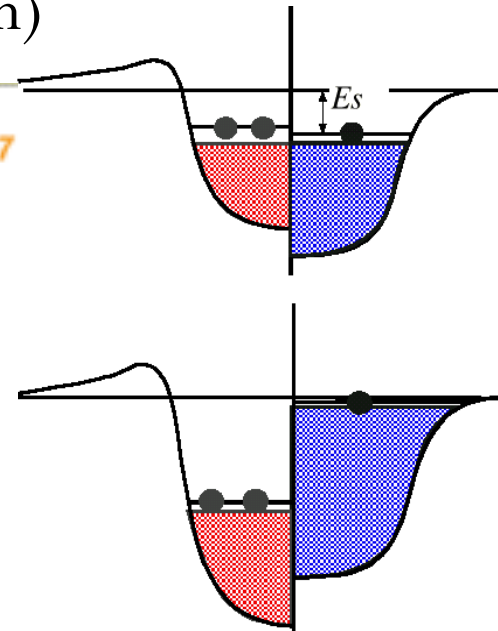
Where is the nuclear dripline?

Limits defined as last isotope with positive neutron separation energy

- Nucleons “drip” out of nucleus

Neutron dripline experimentally established to $Z=8$ (Oxygen)

²⁸ Si	²⁹ Si	³⁰ Si	³¹ Si	³² Si	³³ Si	³⁴ Si	³⁵ Si	³⁶ Si	³⁷ Si	³⁸ Si	³⁹ Si	⁴⁰ Si	⁴¹ Si	⁴² Si	⁴³ Si	⁴⁴ Si	2007
²⁷ Al	²⁸ Al	²⁹ Al	³⁰ Al	³¹ Al	³² Al	³³ Al	³⁴ Al	³⁵ Al	³⁶ Al	³⁷ Al	³⁸ Al	³⁹ Al	⁴⁰ Al	⁴¹ Al	⁴² Al	⁴³ Al	
²⁶ Mg	²⁷ Mg	²⁸ Mg	²⁹ Mg	³⁰ Mg	³¹ Mg	³² Mg	³³ Mg	³⁴ Mg	³⁵ Mg	³⁶ Mg	³⁷ Mg	³⁸ Mg	⁴⁰ Mg				
⁵⁷ Na	²⁶ Na	²⁷ Na	²⁸ Na	²⁹ Na	³⁰ Na	³¹ Na	³² Na	³³ Na	³⁴ Na	³⁵ Na	³⁷ Na	2002					
²⁴ Ne	²⁵ Ne	²⁶ Ne	²⁷ Ne	²⁸ Ne	²⁹ Ne	³⁰ Ne	³¹ Ne	³² Ne	³⁴ Ne	2002							
²³ F	²⁴ F	²⁵ F	²⁶ F	²⁷ F		²⁹ F	³¹ F	1999									
²² O	²³ O	²⁴ O	1970														
²¹ N	²² N	²³ N															
²⁰ C		²² C															



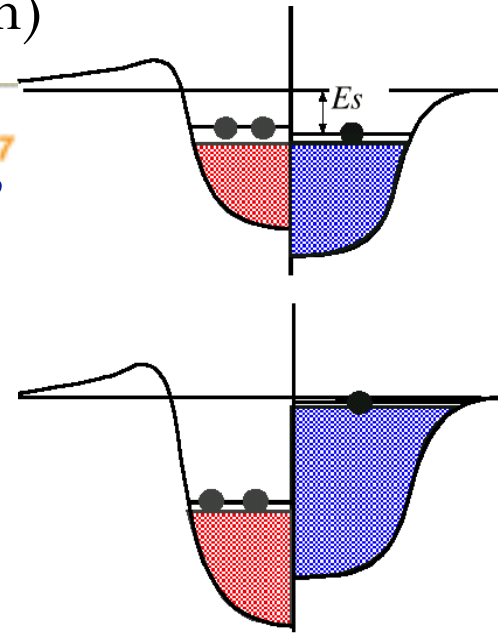
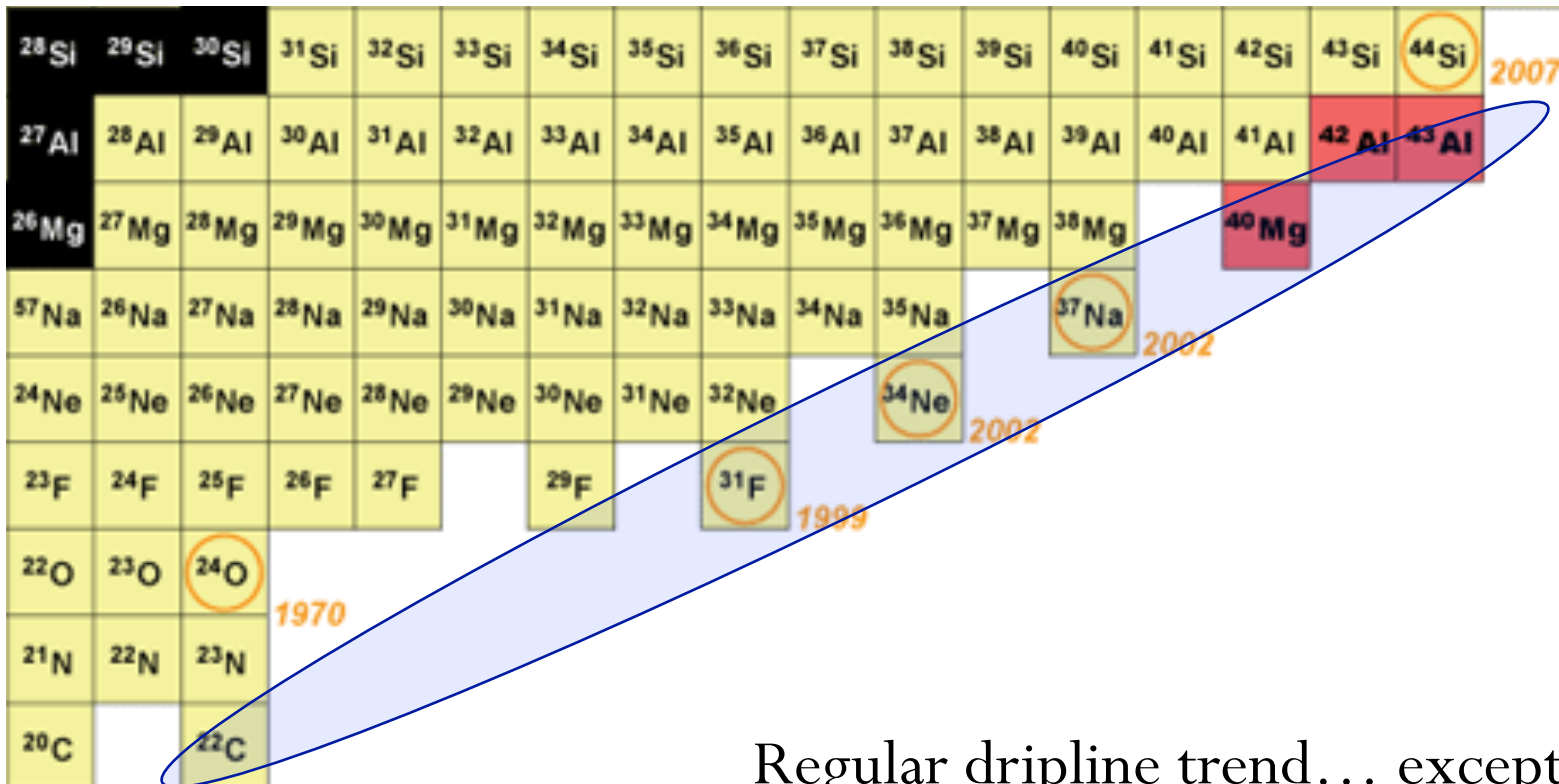
Limits of Nuclear Existence: Oxygen Anomaly

Where is the nuclear dripline?

Limits defined as last isotope with positive neutron separation energy

- Nucleons “drip” out of nucleus

Neutron dripline experimentally established to $Z=8$ (Oxygen)



Regular dripline trend... except oxygen

Adding one proton binds 6 additional neutrons

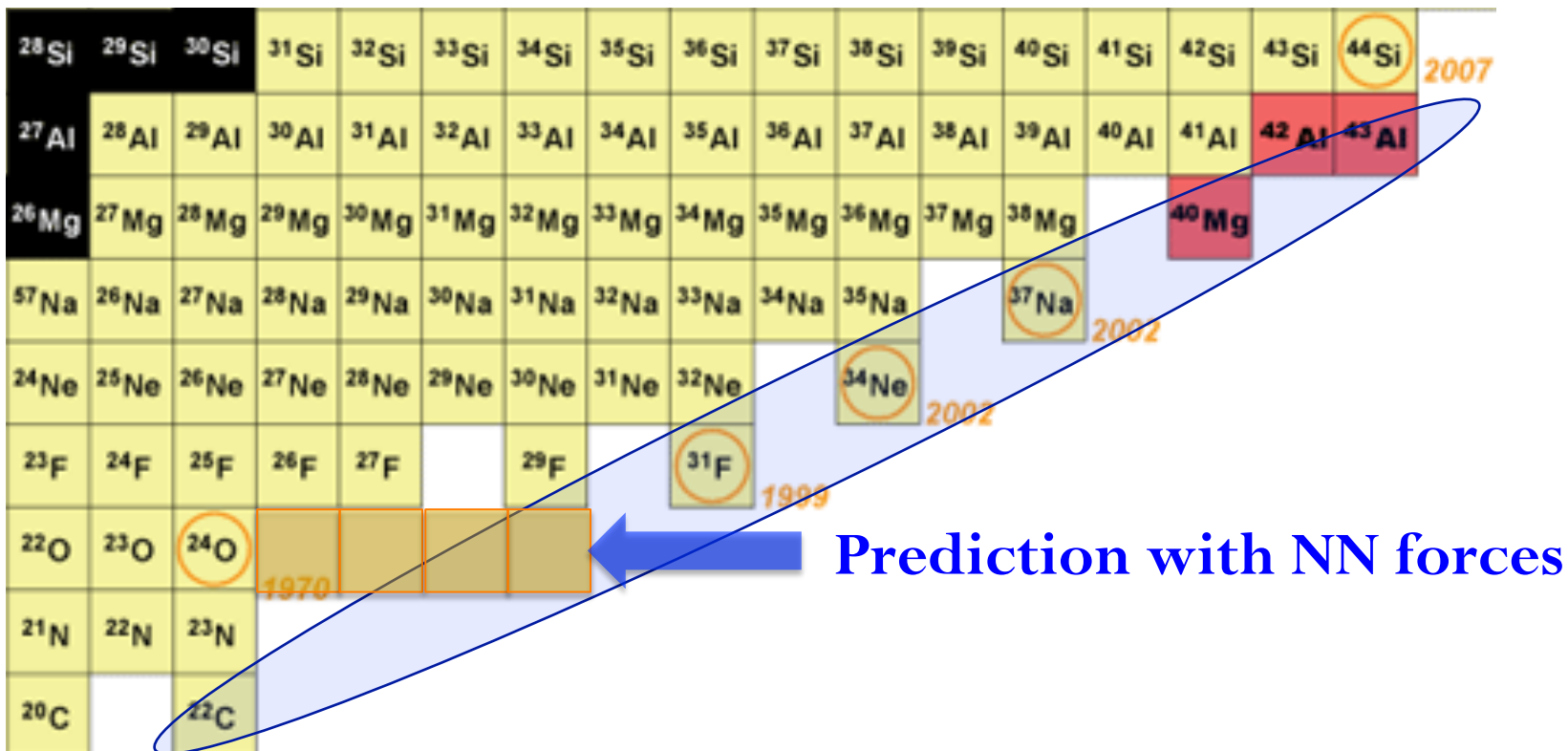
Limits of Nuclear Existence: Oxygen Anomaly

Where is the nuclear dripline?

Limits defined as last isotope with positive neutron separation energy

- Nucleons “drip” out of nucleus

Neutron dripline experimentally established to $Z=8$ (Oxygen)



Microscopic picture: **NN-forces too attractive**

Incorrect prediction of dripline

Monopole Part of Valence-Space Interactions

Microscopic MBPT – effective interaction in chosen model space

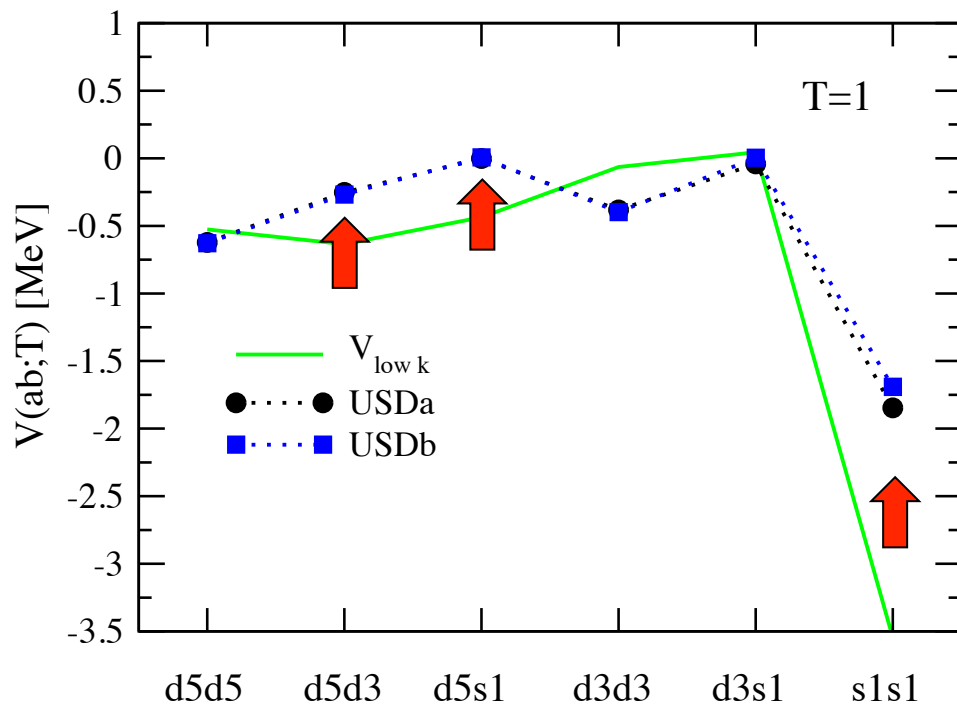
Works near closed shells: deteriorates beyond this

Deficiencies improved adjusting particular two-body matrix elements

Monopoles:
Angular average of interaction

$$V_{ab}^T = \frac{\sum_J (2J + 1) V_{abab}^{JT}}{\sum_J (2J + 1)}$$

Determines interaction of orbit a with b : evolution of orbital energies



$$\Delta \varepsilon_a = V_{ab} n_b$$

Microscopic **low-momentum** interactions

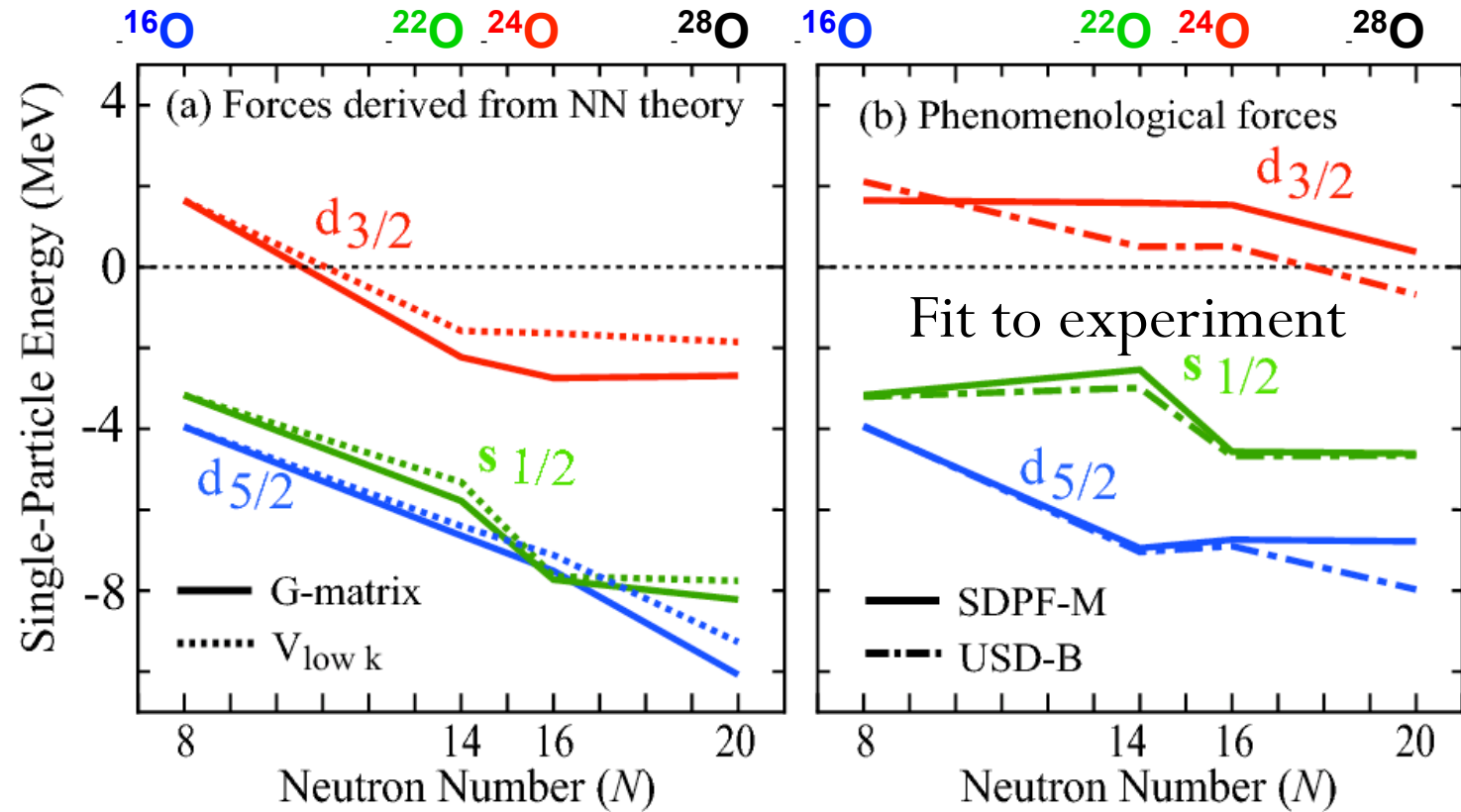
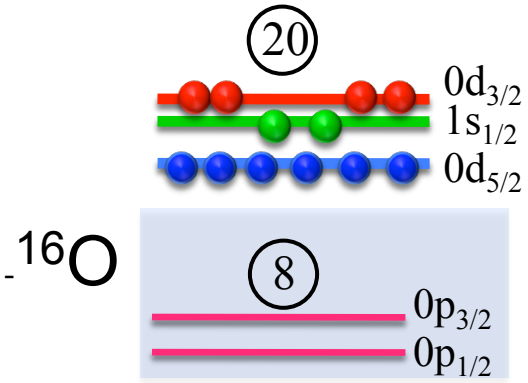
Phenomenological **USD** interactions

Clear shifts in **low-lying orbitals**:

- T=1 repulsive shift

Physics in Oxygen Isotopes

Calculate evolution of sd -orbital energies from interactions



Microscopic NN Theories

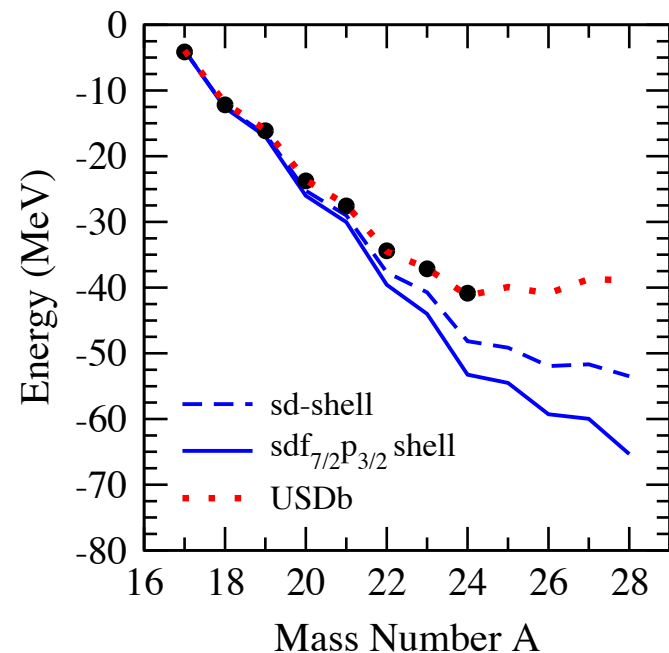
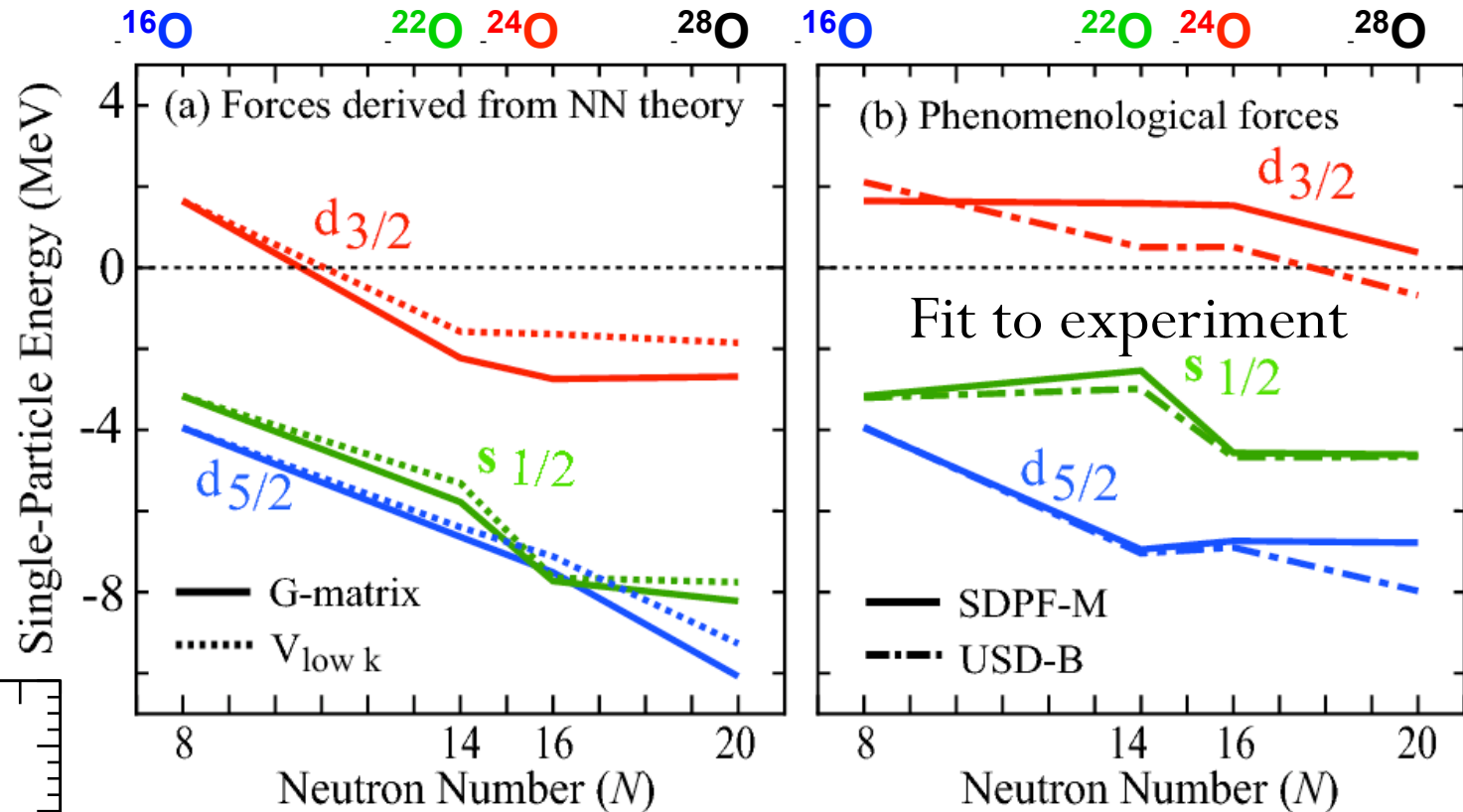
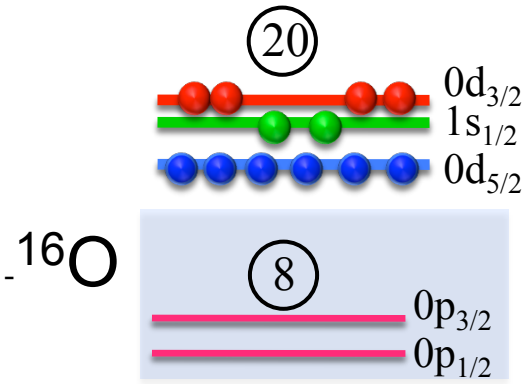
$d_{3/2}$ orbit bound to ^{28}O

Phenomenological Models

$d_{3/2}$ orbit unbound

Physics in Oxygen Isotopes

Calculate evolution of sd -orbital energies from interactions



Microscopic NN Theories

$d_{3/2}$ orbit bound to ^{28}O

Dripline at ^{28}O

Phenomenological Models

$d_{3/2}$ orbit unbound

Dripline at ^{24}O

Oxygen anomaly unexplained with NN forces

Origin of monopole shifts: Neglected 3N forces

-- See lecture of A. Poves

Perturbative Approach

- 1) Effective Hamiltonian: sum excitations outside valence space to MBPT(3)
- 2) Self-consistent single-particle energies
- 3) Harmonic-oscillator basis of 13-15 major shells
- 4) Nuclear forces from chiral EFT
- 5) Requires extended valence spaces

Limitations

- Uncertain perturbative convergence
- Core physics inconsistent or absent
- Degenerate valence space requires HO basis (HF requires nontrivial extension)
- Must treat additional orbitals nonperturbatively (extend valence space)

Particle/Hole Excitations

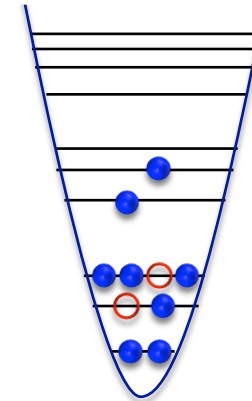
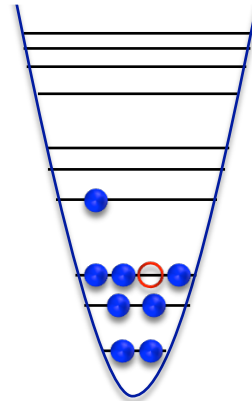
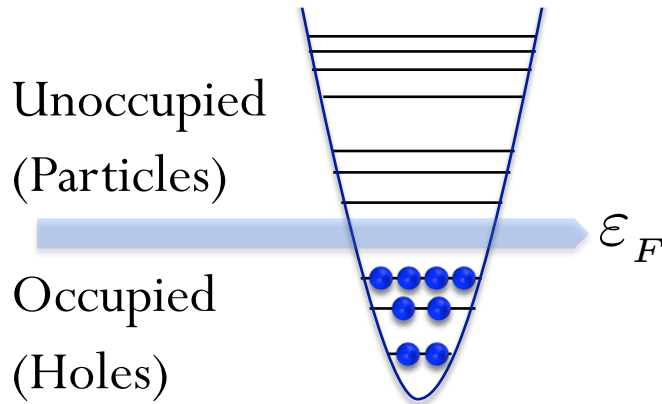
Consider basis states as excitations from some reference state:

Reference

Slater Determinant

1p-1h excitation

2p-2h excitation



$$|\Phi\rangle = \prod_{i=1}^N a_i^\dagger |0\rangle$$

$$|\Phi_i^a\rangle = a_a^\dagger a_i |\Phi\rangle$$

$$|\Phi_{ij}^{ab}\rangle = a_a^\dagger a_i a_b^\dagger a_j |\Phi\rangle$$

Hamiltonian schematically given in terms of ph excitations

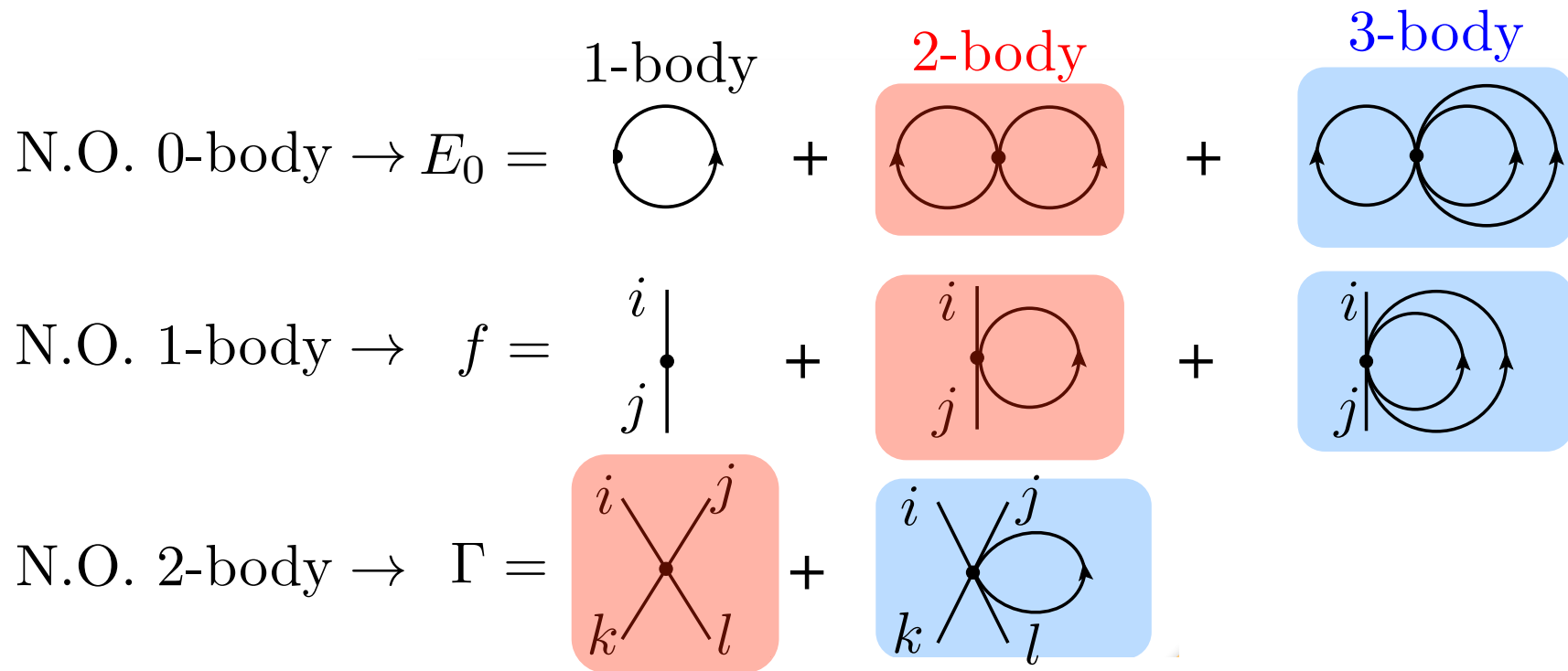
	0p-0h	1p-1h	2p-2h	3p-3h
0p-0h				
1p-1h				
2p-2h				
3p-3h				

$\langle i|H|j\rangle$

Normal-Ordered Hamiltonian

Now rewrite exactly the initial Hamiltonian in normal-ordered form

$$H_{\text{N.O.}} = E_0 + \sum_{ij} f_{ij} \{a_i^\dagger a_j\} + \frac{1}{4} \sum_{ijkl} \Gamma_{ijkl} \{a_i^\dagger a_j^\dagger a_l a_k\} + \frac{1}{36} \sum_{ijklmn} W_{ijklmn} \{a_i^\dagger a_j^\dagger a_k^\dagger a_l a_m a_n\}$$



Normal-ordered Hamiltonian w.r.t. reference state

Loop = **sum over occupied states**

Include dominant 1-, 2-, 3-body physics in NO

Nonperturbative In-Medium SRG

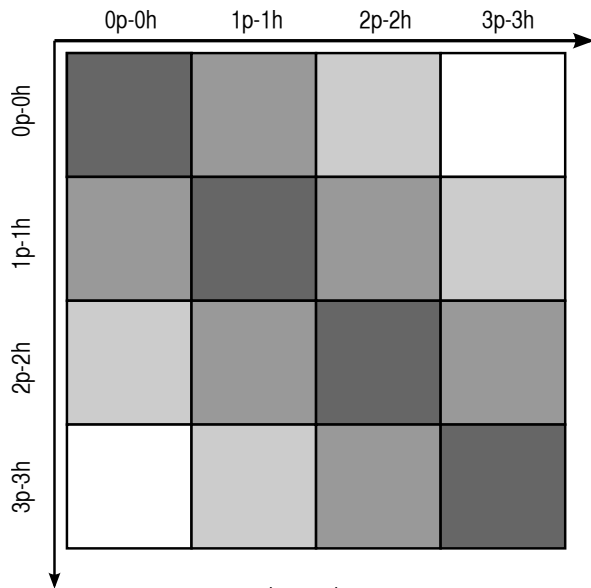
Tsukiyama, **Bogner**, Schwenk, PRL (2011)

In-Medium SRG continuous unitary trans. drives off-diagonal physics to zero

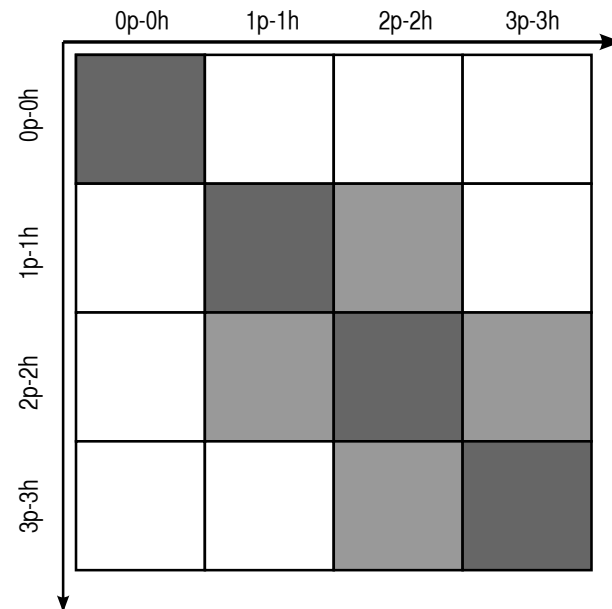
$$H(s) = U(s)HU^\dagger(s) \equiv H^d(s) + H^{\text{od}}(s) \rightarrow H^d(\infty)$$

From uncorrelated Hartree-Fock reference state (e.g., ^{16}O) define:

$$H^{\text{od}} = \langle p|H|h\rangle + \langle pp|H|hh\rangle + \dots + \text{h.c.}$$



$$\langle i|H|j\rangle$$



$$\langle npnh|H(\infty)|\Phi_{\text{core}}\rangle = 0$$

Drives all n-particle n-hole couplings to 0 – **decouples** core from excitations

IM-SRG: Flow Equation Formulation

Define $U(s)$ implicitly from particular choice of generator:

$$\eta(s) \equiv (dU(s)/ds) U^\dagger(s)$$

chosen for desired decoupling behavior – e.g.,

$$\eta_I(s) = [H^d(s), H^{\text{od}}(s)] \quad \text{Wegner (1994)}$$

Solve **flow equation** for Hamiltonian (coupled DEs for 0, 1, 2-body parts)

$$\frac{dH(s)}{ds} = [\eta(s), H(s)] \quad H(s) = E_0(s) + f(s) + \Gamma(s) + \dots$$

Hamiltonian and generator truncated at 2-body level: **IM-SRG(2)**

0-body flow drives uncorrelated ref. state to fully correlated ground state

$$E_0(\infty) \rightarrow \text{Core Energy}$$

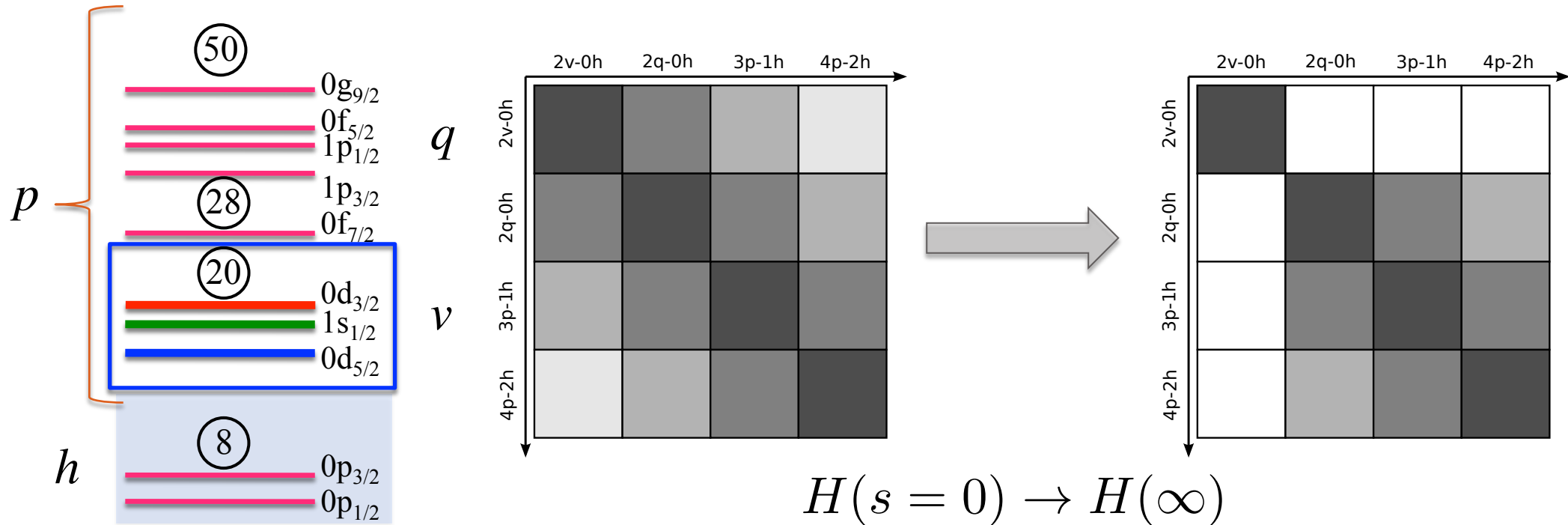
Ab initio method for energies of **closed-shell systems**

IM-SRG: Valence-Space Hamiltonians

Tsukiyama, **Bogner**, Schwenk, PRC (2012)

Open-shell systems

Separate p states into valence states (v) and those above valence space (q)



Redefine H^{od} to **decouple valence space from excitations** outside v

$$H^{\text{od}} = \langle p|H|h\rangle + \langle pp|H|hh\rangle + \langle v|H|q\rangle + \langle pq|H|vv\rangle + \langle pp|H|hv\rangle + \text{h.c.}$$

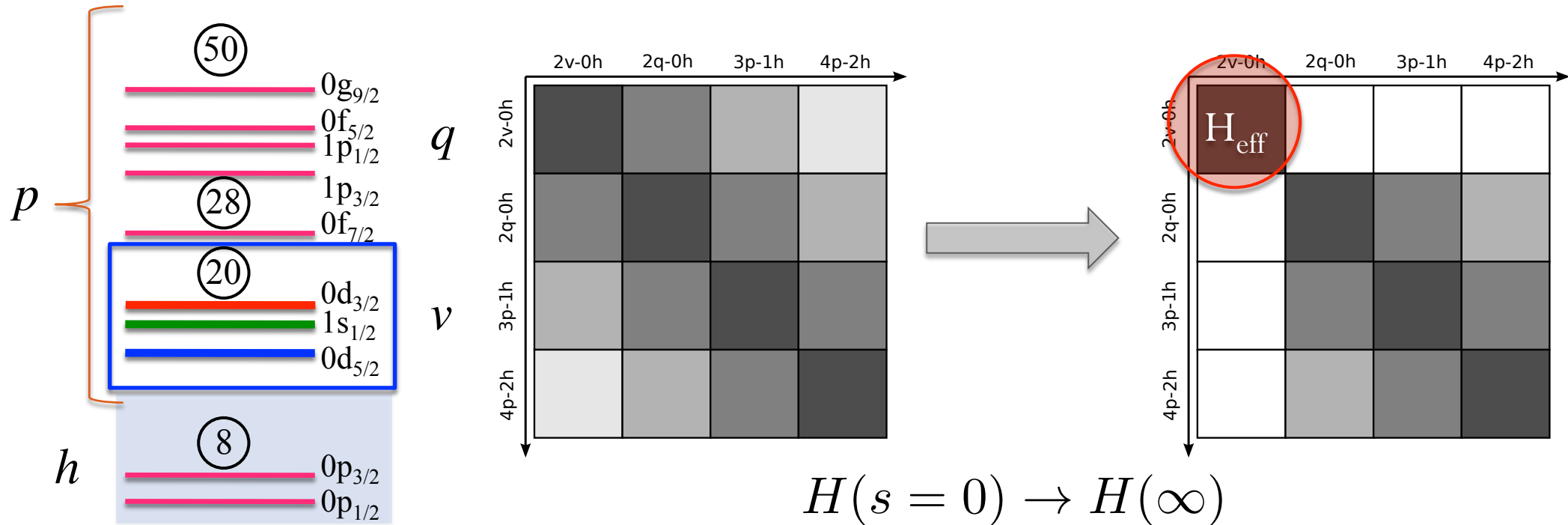
$$E_0(\infty) \rightarrow \text{Core Energy} \quad f(\infty) \rightarrow \text{SPEs} \quad \Gamma(\infty) \rightarrow V_{\text{eff}}$$

IM-SRG: Valence-Space Hamiltonians

Tsukiyama, **Bogner**, Schwenk, PRC (2012)

Open-shell systems

Separate p states into valence states (v) and those above valence space (q)



Core physics included consistently (**absolute energies, radii...**)

Inherently nonperturbative – no need for extended valence space

Non-degenerate valence-space orbitals

NN-only IM-SRG Monopoles

Testing ab initio IM-SRG shell model monopoles

Monopoles:
Angular average of interaction

$$V_{ab}^T = \frac{\sum_J (2J + 1) V_{abab}^{JT}}{\sum_J (2J + 1)}$$

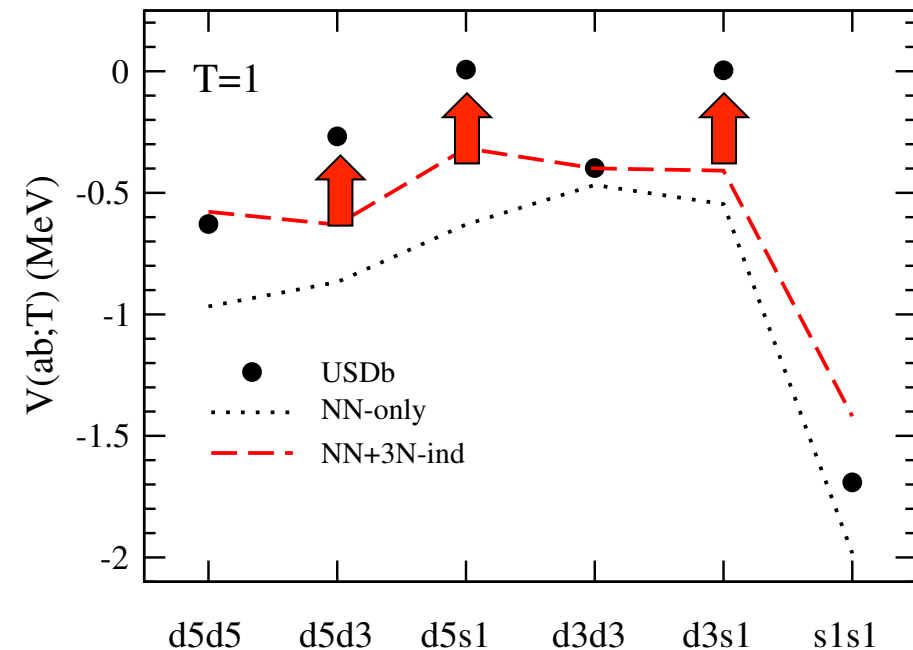
Determines interaction of orbit a with b : evolution of orbital energies

$$\Delta \epsilon_a = V_{ab} n_b$$

Improvements over MBPT?

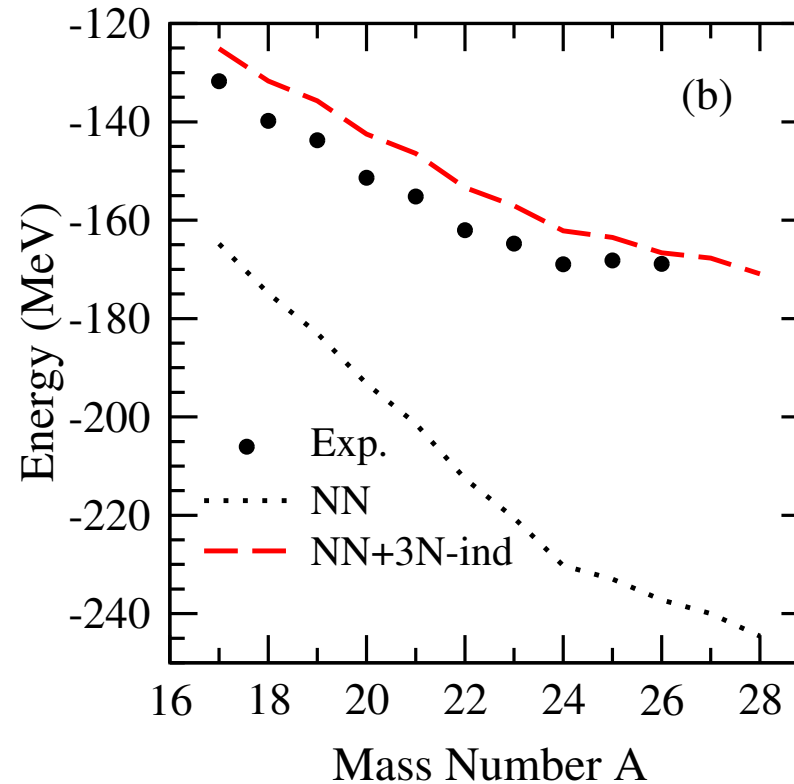
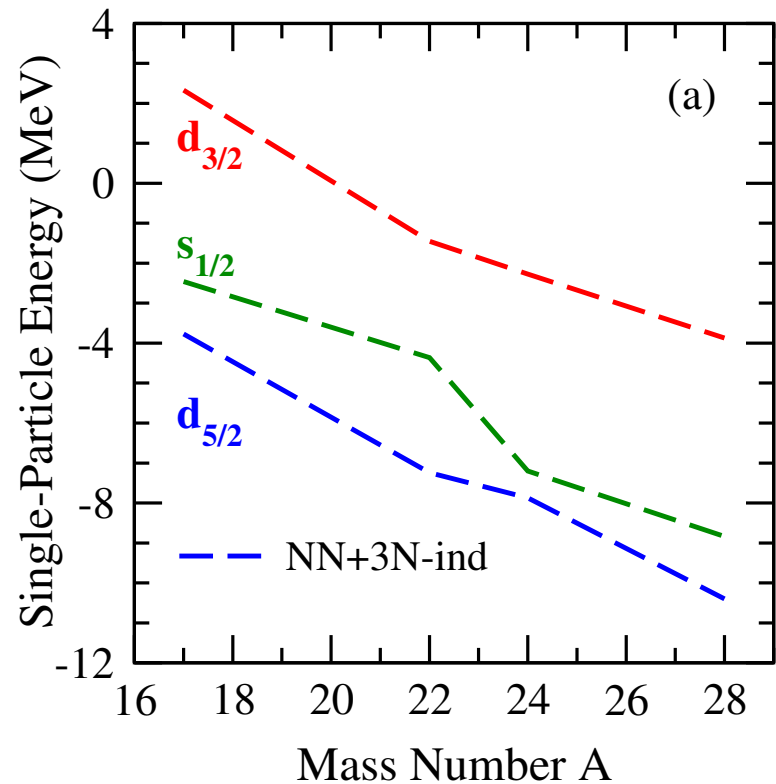
NN-only significantly too attractive

NN+3N-ind improved but $d_{3/2}$ monopoles too attractive



Comparison with Large-Space Methods

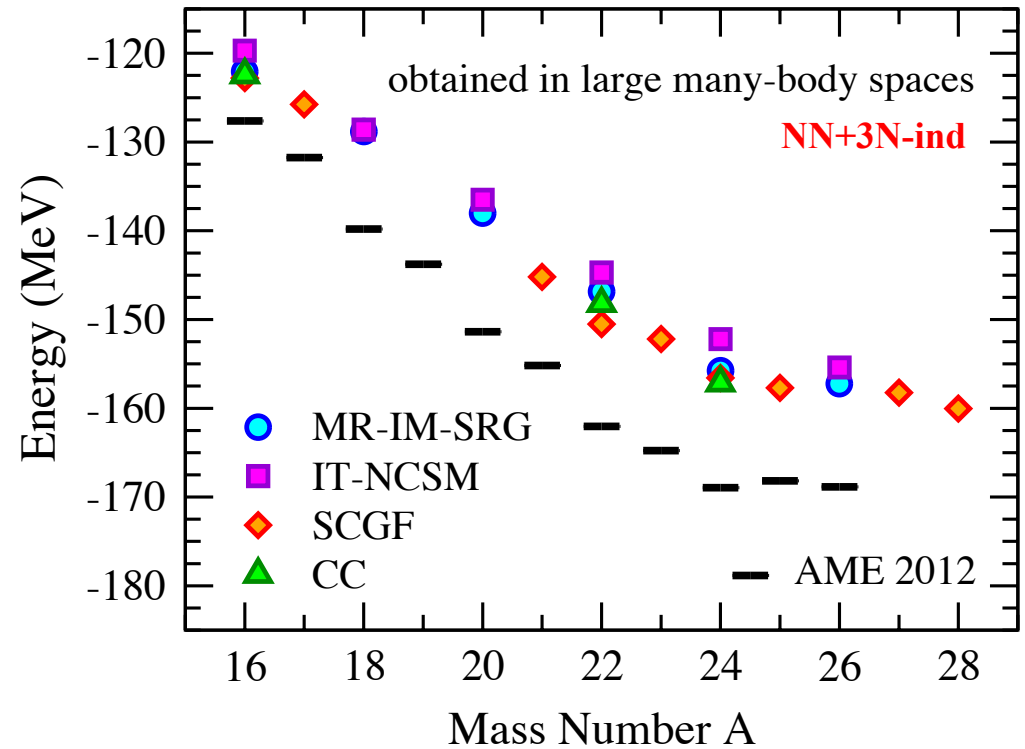
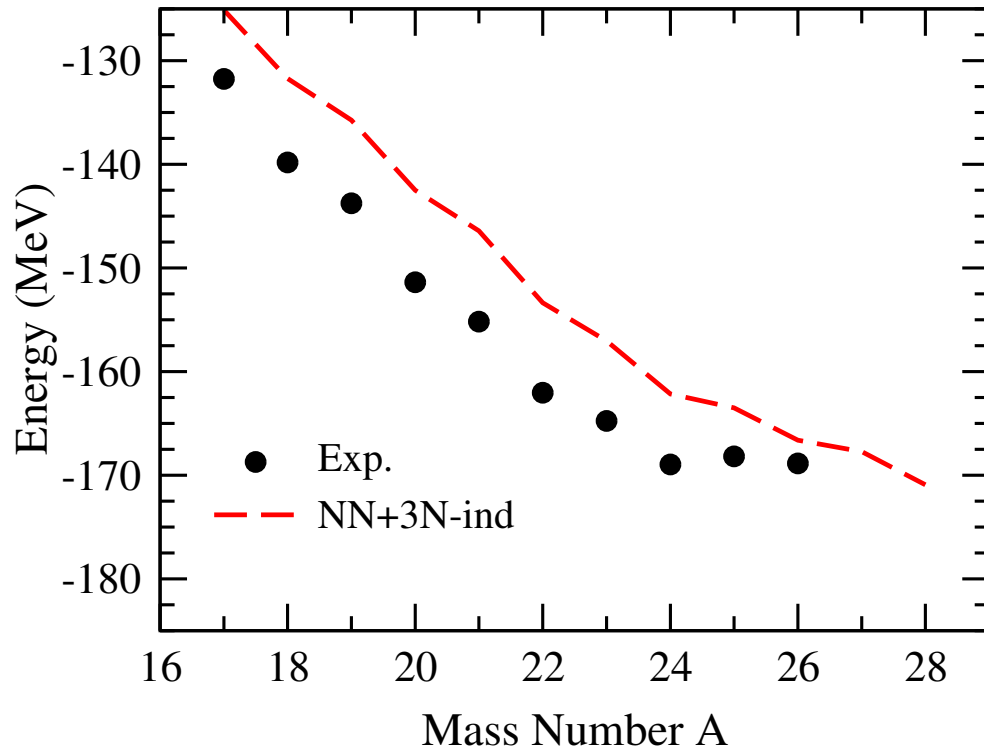
Results from SRG-evolved NN and NN+3N-ind forces



Dripline still not reproduced

Comparison with Large-Space Methods

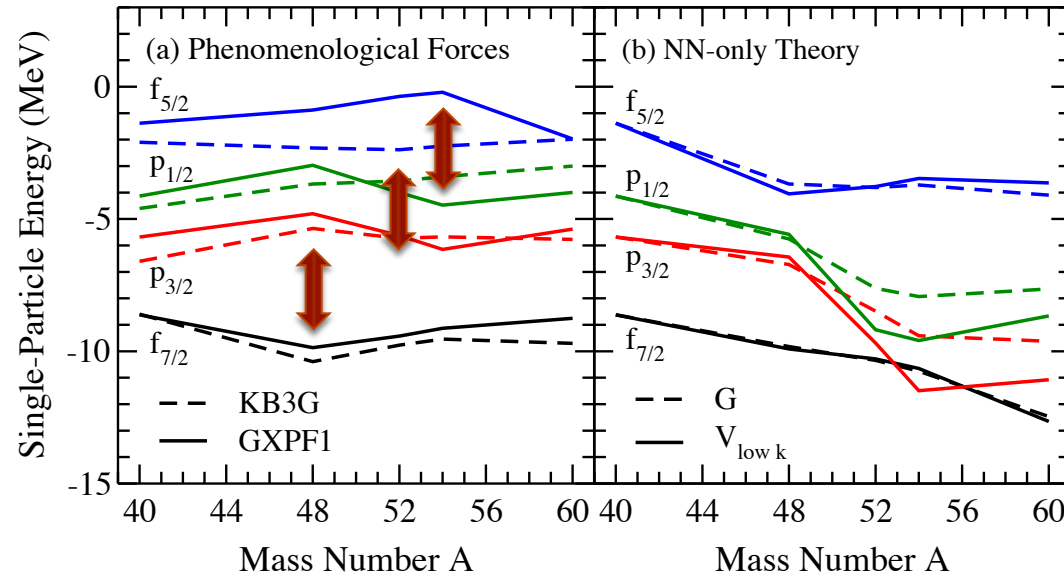
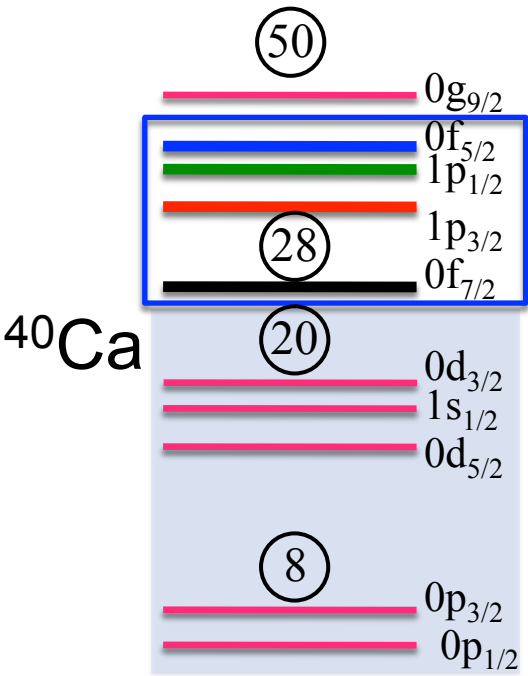
Large-space methods with **same SRG-evolved NN+3N-ind forces**



Agreement between all methods with same input forces

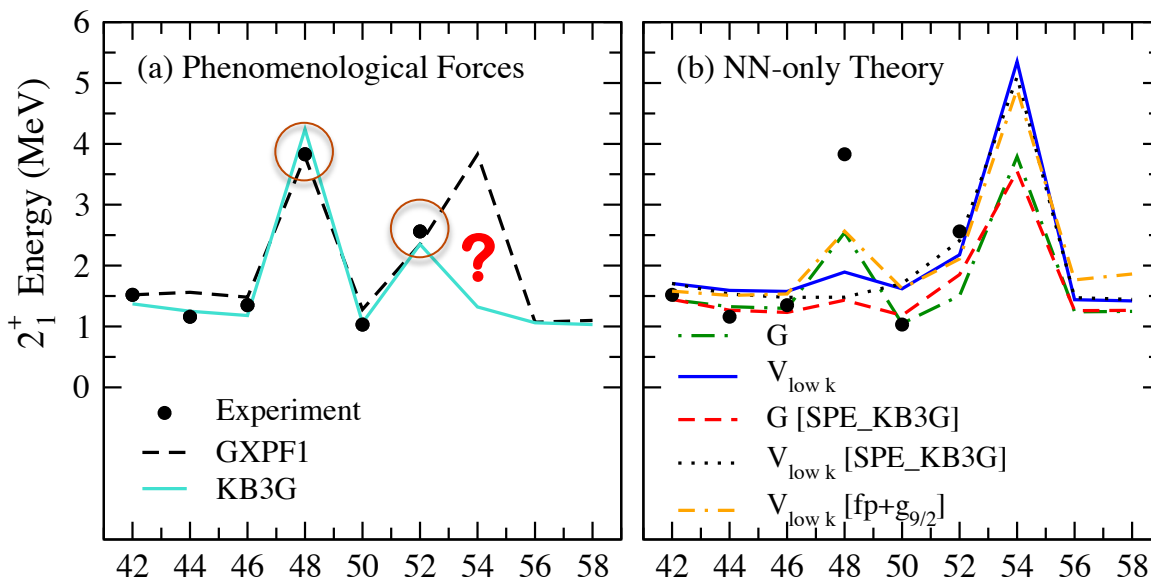
No reproduction of dripline in any case

Calcium Isotopes: Magic Numbers



GXPF1: Honma, Otsuka, Brown, Mizusaki (2004)

KB3G: Poves, Sanchez-Solano, Caurier, Nowacki (2001)



Phenomenological Forces

Large gap at ^{48}Ca

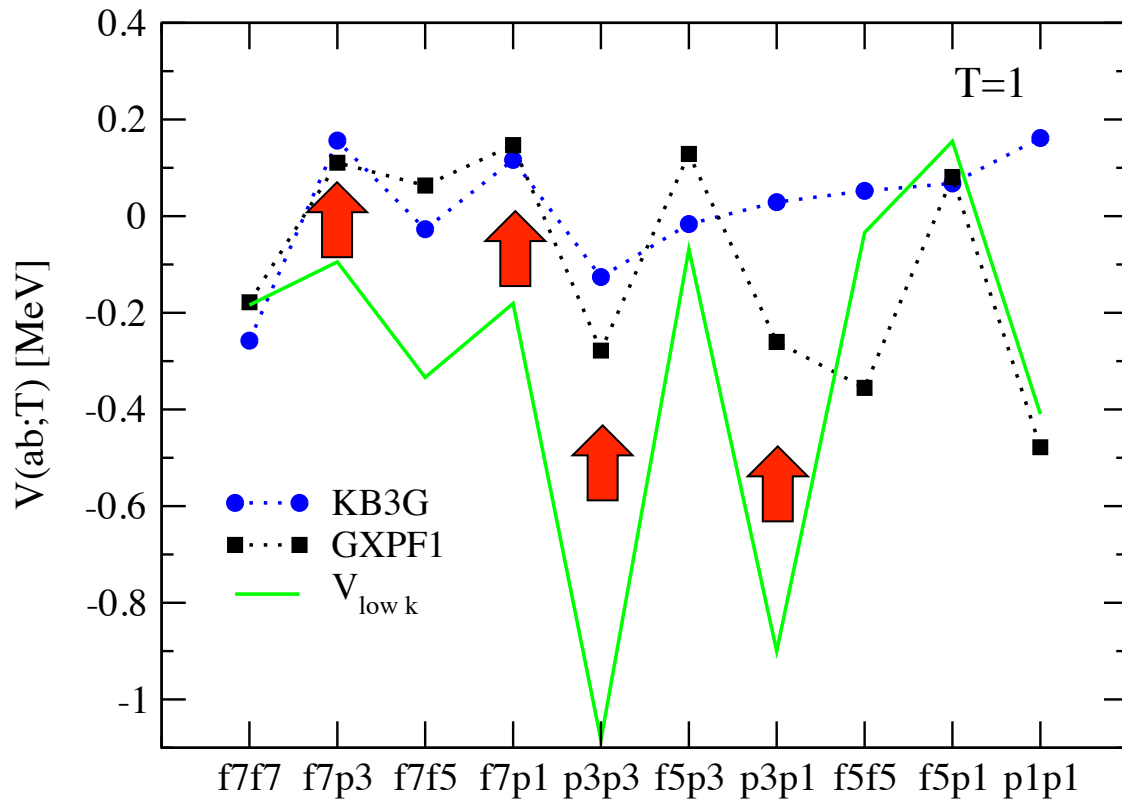
Discrepancy at $N=34$

Microscopic NN Theory

Small gap at ^{48}Ca

N=28: first standard magic number not reproduced in microscopic NN theories

Phenomenological vs. Microscopic



Compare monopoles from:

Microscopic **low-momentum** interactions

Phenomenological **KB3G, GXPF1** interactions

Shifts in **low-lying orbitals:**

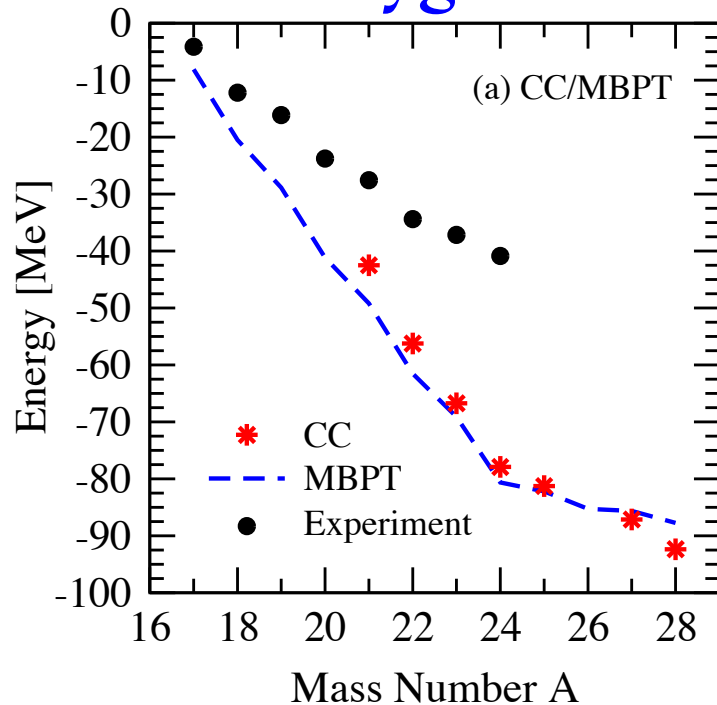
- T=1 repulsive shift

Comparison to Coupled Cluster

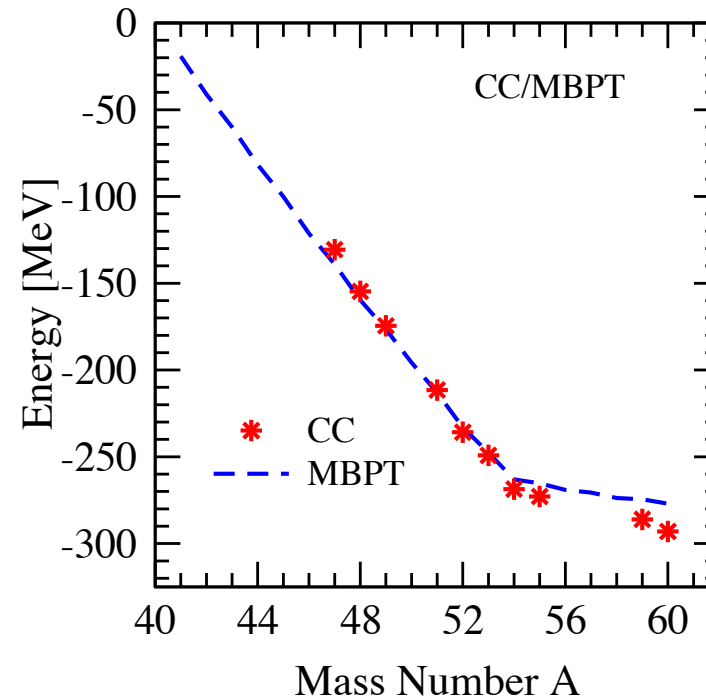
Many-body method insufficient?

Benchmark against *ab-initio* Coupled Cluster at NN-only level

Oxygen



Calcium



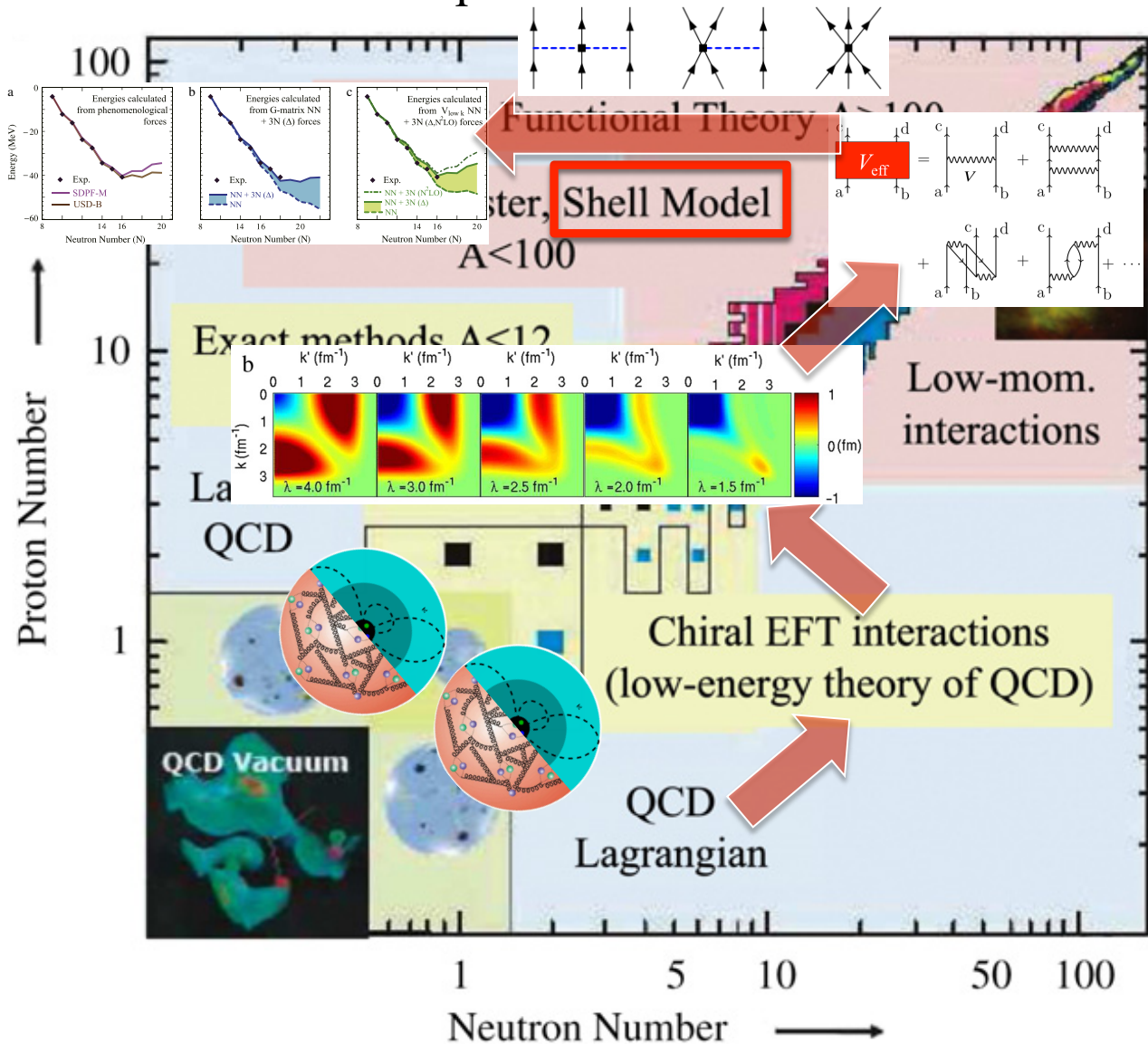
SPEs: one-particle attached CC energies in ^{17}O and ^{41}Ca

Small difference in many-body methods

Include **3N forces** to improve agreement with experiment

Part IV: Three-Nucleon Forces to Nuclei

To understand the properties of complex nuclei from first principles



Three-Nucleon Forces

Basic ideas – why needed?

3N from chiral EFT

Implementing in shell model

Relation to monopoles

Predictions/new discoveries

Connections beyond structure

How will we approach this problem:

QCD → NN (3N) forces → Renormalize → “Solve” many-body problem → Predictions

Chiral Effective Field Theory: Nuclear Forces

Nucleons interact via pion exchanges and contact interactions

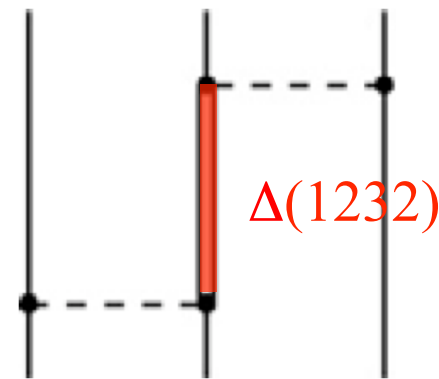
Consistent treatment of NN, 3N, ...

NN couplings fit to scattering data

	NN	3N	4N
LO $O\left(\frac{Q^0}{\Lambda^0}\right)$		—	—
NLO $O\left(\frac{Q^2}{\Lambda^2}\right)$		—	—
N ² LO $O\left(\frac{Q^3}{\Lambda^3}\right)$			—
N ³ LO $O\left(\frac{Q^4}{\Lambda^4}\right)$			

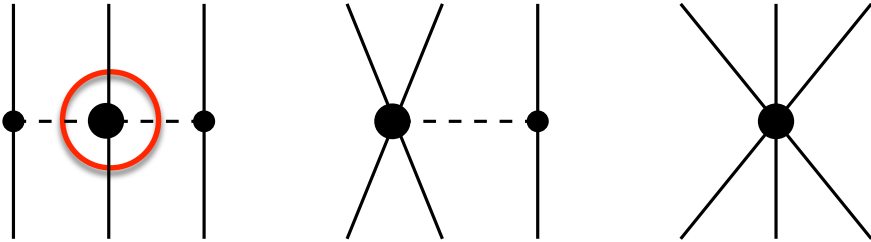
derived in (1994/2002)

(2011) (2006)



Chiral EFT: N²LO 3N

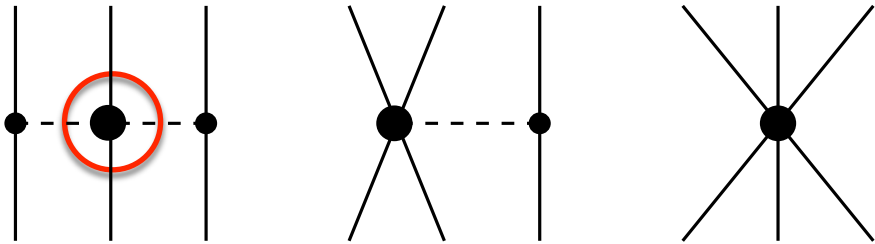
First non-vanishing 3N contributions: Next-to-next-to-leading order $\nu = 3$



$$\begin{aligned}
 V_{3N}^{(3)} = & \frac{g_A^2}{8F_\pi^4} \frac{\vec{\sigma}_1 \cdot \vec{q}_1 \vec{\sigma}_3 \cdot \vec{q}_3}{[q_1^2 + M_\pi^2][q_3^2 + M_\pi^2]} \left[\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 (-4c_1 M_\pi^2 \right. \\
 & + 2c_3 \vec{q}_1 \cdot \vec{q}_3) + c_4 \boldsymbol{\tau}_1 \times \boldsymbol{\tau}_3 \cdot \boldsymbol{\tau}_2 \vec{q}_1 \times \vec{q}_3 \cdot \vec{\sigma}_2 \left. \right] \\
 & - \frac{g_A D}{8F_\pi^2} \frac{\vec{\sigma}_3 \cdot \vec{q}_3}{q_3^2 + M_\pi^2} \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \vec{\sigma}_1 \cdot \vec{q}_3 + \frac{1}{2} E \boldsymbol{\tau}_2 \cdot \boldsymbol{\tau}_3
 \end{aligned}$$

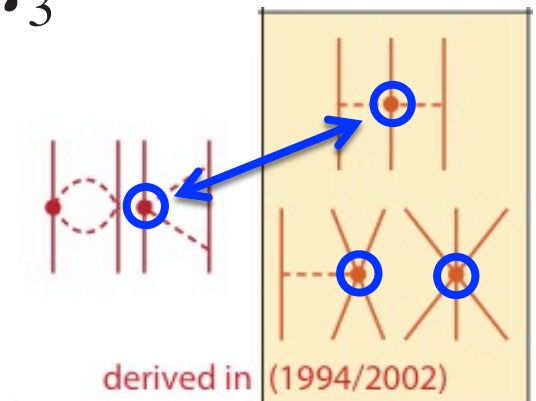
Chiral EFT: N²LO 3N

First non-vanishing 3N contributions: Next-to-next-to-leading order $\nu = 3$



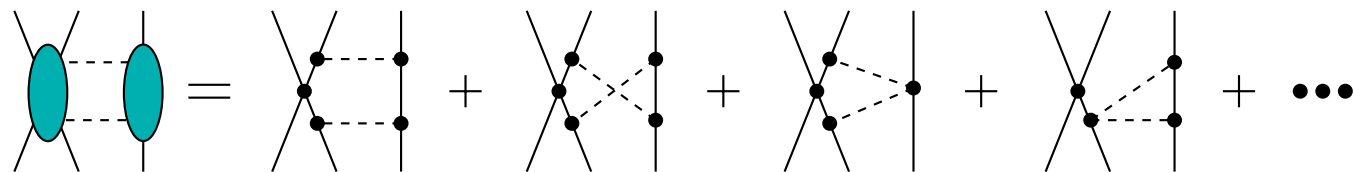
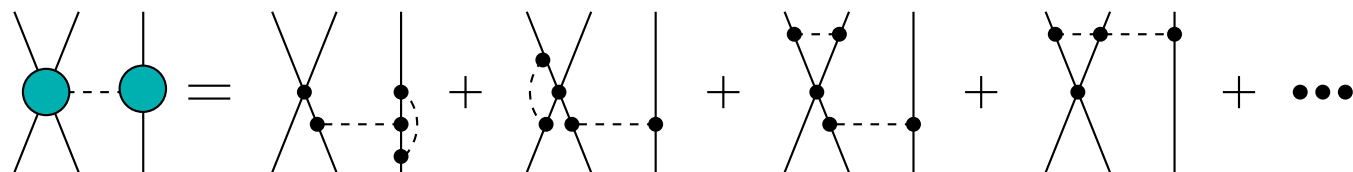
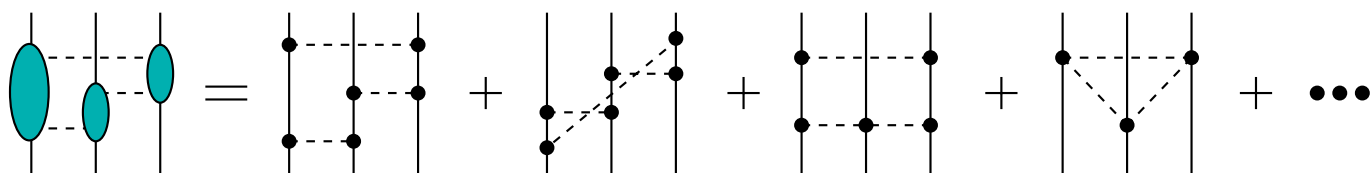
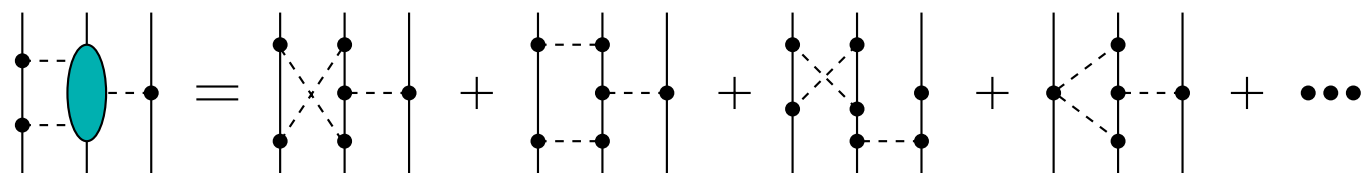
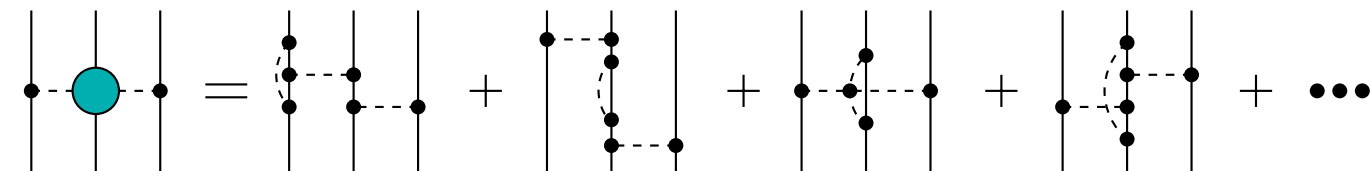
$$\begin{aligned}
 V_{3N}^{(3)} = & \frac{g_A^2}{8F_\pi^4} \frac{\vec{\sigma}_1 \cdot \vec{q}_1 \vec{\sigma}_3 \cdot \vec{q}_3}{[q_1^2 + M_\pi^2][q_3^2 + M_\pi^2]} [\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 (-4c_1 M_\pi^2 \\
 & + 2c_3 \vec{q}_1 \cdot \vec{q}_3) + c_4 \boldsymbol{\tau}_1 \times \boldsymbol{\tau}_3 \cdot \boldsymbol{\tau}_2 \vec{q}_1 \times \vec{q}_3 \cdot \vec{\sigma}_2] \\
 & - \frac{g_A D}{8F_\pi^2} \frac{\vec{\sigma}_3 \cdot \vec{q}_3}{q_3^2 + M_\pi^2} \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \vec{\sigma}_1 \cdot \vec{q}_3 + \frac{1}{2} E \boldsymbol{\tau}_2 \cdot \boldsymbol{\tau}_3
 \end{aligned}$$

Three undetermined π N couplings from NN fit



Chiral EFT: $N^3\text{LO}$ 3N

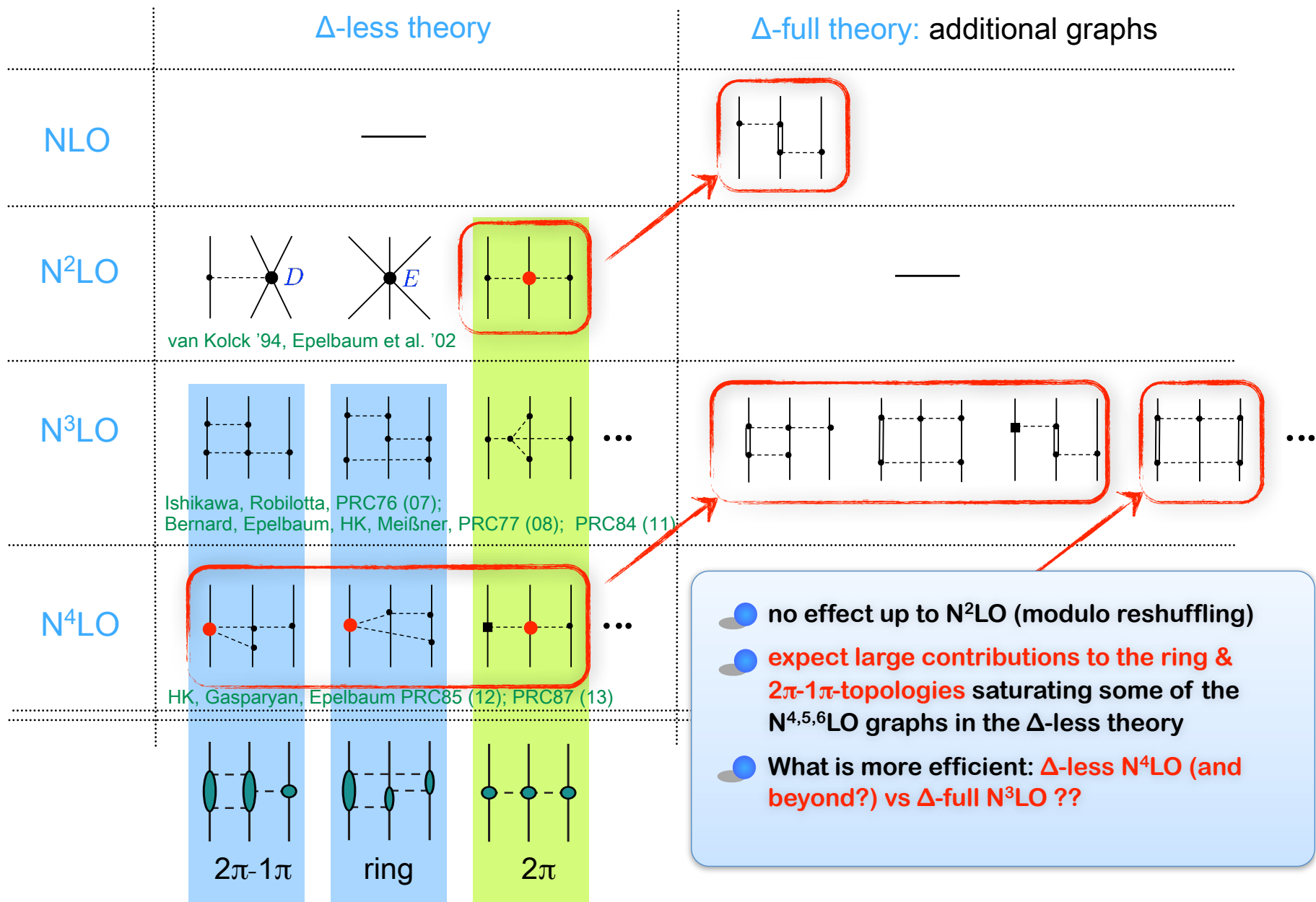
Next-to-next-to-next-to-leading order $\nu = 4$



Good news: **no new constants**

Bad news: well, there's all this

Aside: Effects of Adding Explicit Deltas

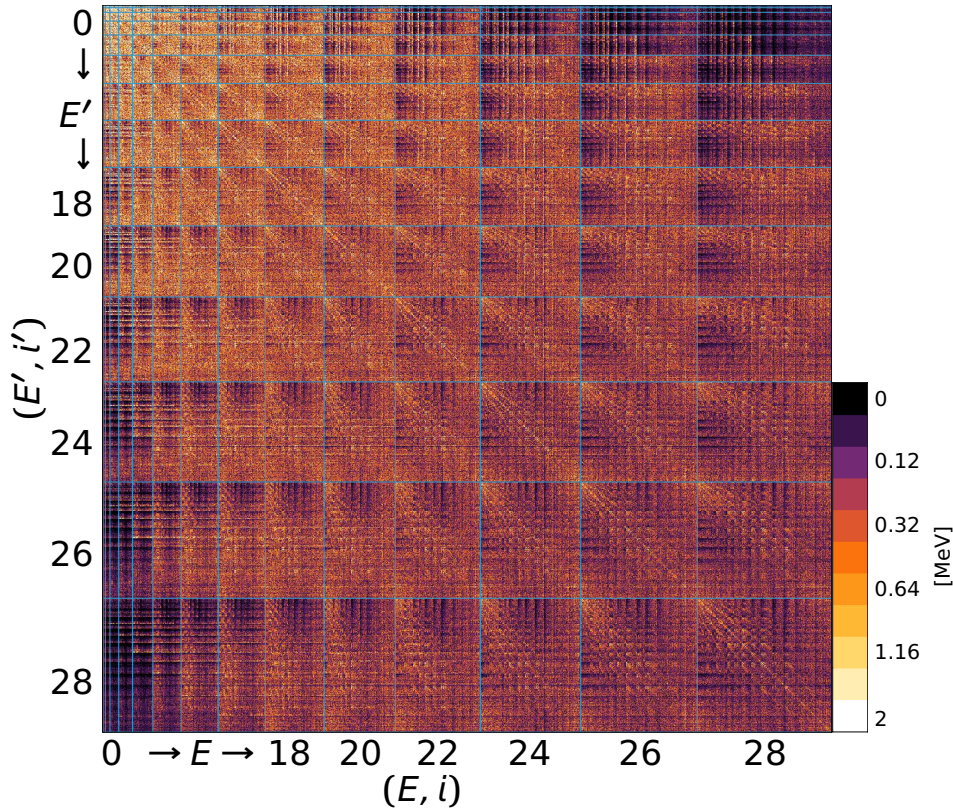


Reshuffles effects to different chiral orders

SRG Evolution in HO Basis

Most common to SRG evolve 3N in HO basis:

3B-Jacobi HO matrix elements



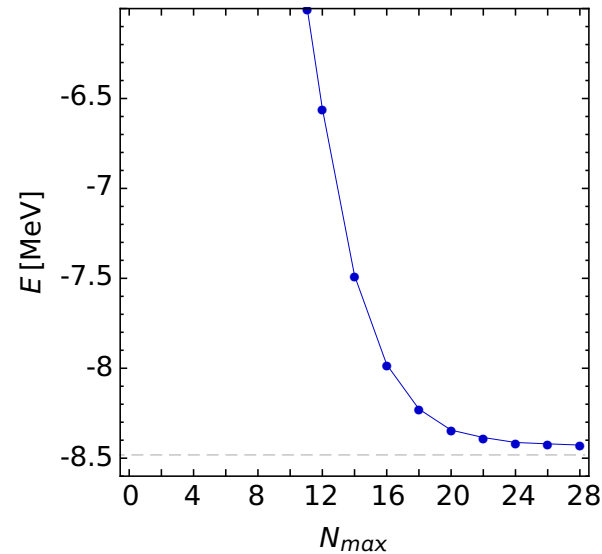
$$\alpha = 0.00 \text{ fm}^4$$

$$\lambda = \infty \text{ fm}^{-1}$$

$$\langle E' i' J T | \tilde{H}_\alpha - T_{\text{int}} | E i J T \rangle$$

$$J^\pi = \frac{1}{2}^+, T = \frac{1}{2}, \hbar\Omega = 24 \text{ MeV}$$

NCSM ground state ${}^3\text{H}$



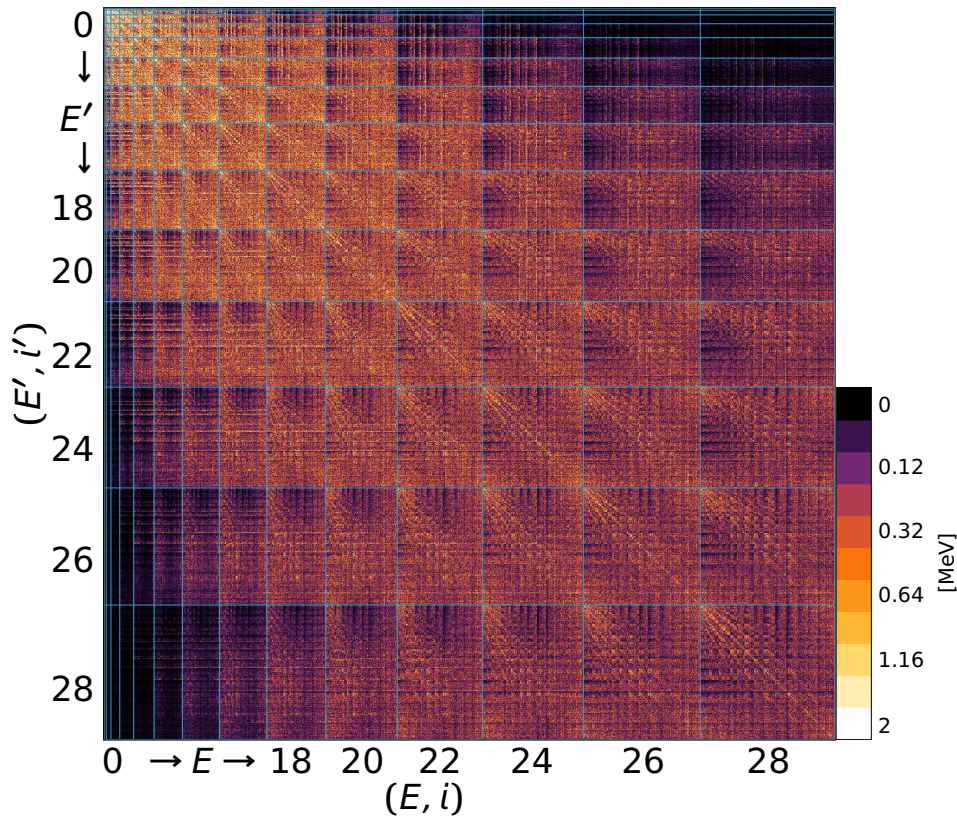
1) SRG-evolve both NN and 3N: NN+3N-full

2) NN Vlowk, refit 3N: NN+3N-fit

SRG Evolution in HO Basis

Most common to SRG evolve 3N in HO basis:

3B-Jacobi HO matrix elements



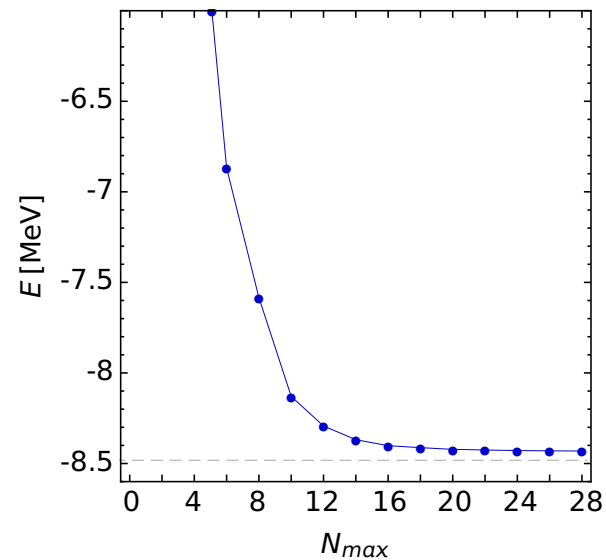
$$\alpha = 0.02 \text{ fm}^4$$

$$\lambda = 2.66 \text{ fm}^{-1}$$

$$\langle E' i' J T | \tilde{H}_\alpha - T_{\text{int}} | E i J T \rangle$$

$$J^\pi = \frac{1}{2}^+, T = \frac{1}{2}, \hbar\Omega = 24 \text{ MeV}$$

NCSM ground state ${}^3\text{H}$



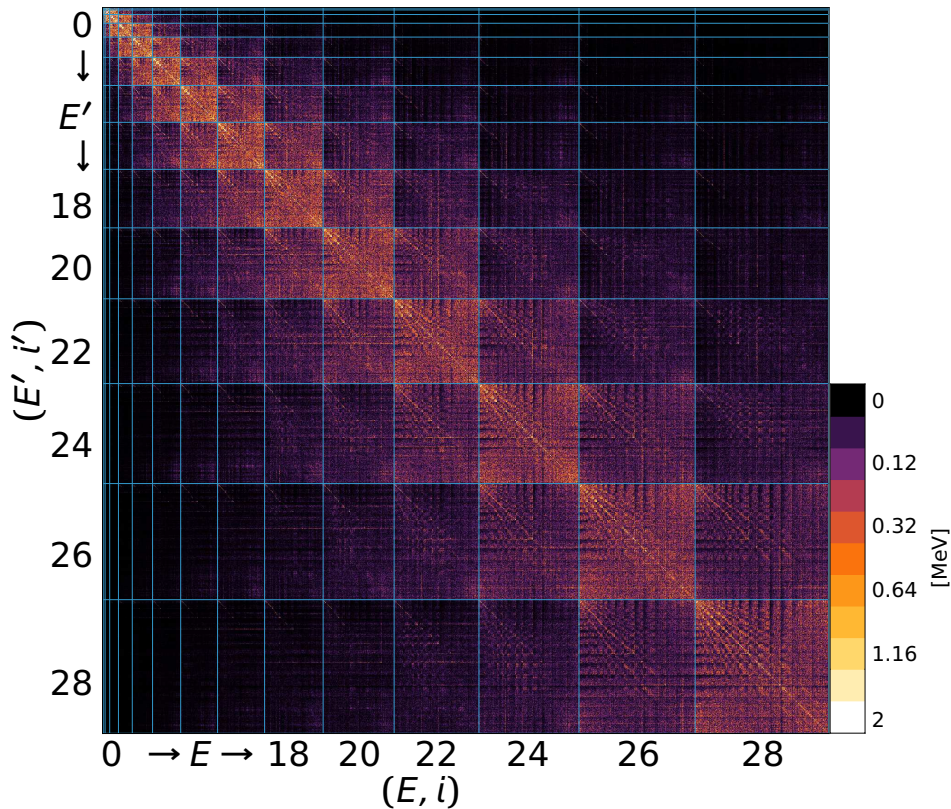
1) SRG-evolve both NN and 3N: NN+3N-full

2) NN Vlowk, refit 3N: NN+3N-fit

SRG Evolution in HO Basis

Most common to SRG evolve 3N in HO basis:

3B-Jacobi HO matrix elements



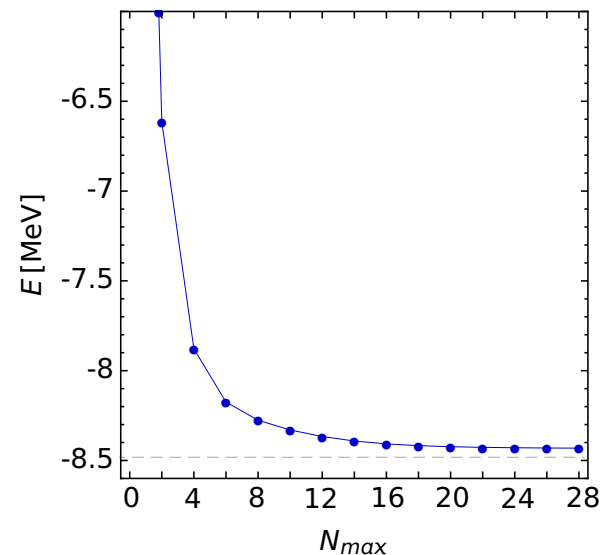
$$\alpha = 1.28 \text{ fm}^4$$

$$\lambda = 0.94 \text{ fm}^{-1}$$

$$\langle E' i' J T | \tilde{H}_\alpha - T_{\text{int}} | E i J T \rangle$$

$$J^\pi = \frac{1}{2}^+, T = \frac{1}{2}, \hbar\Omega = 24 \text{ MeV}$$

NCSM ground state ${}^3\text{H}$

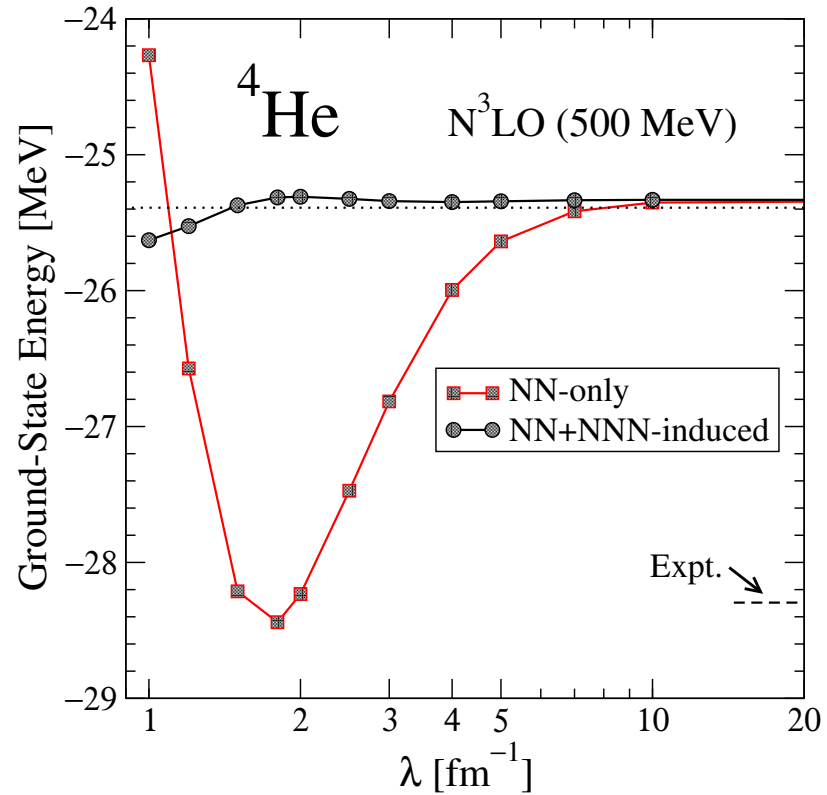
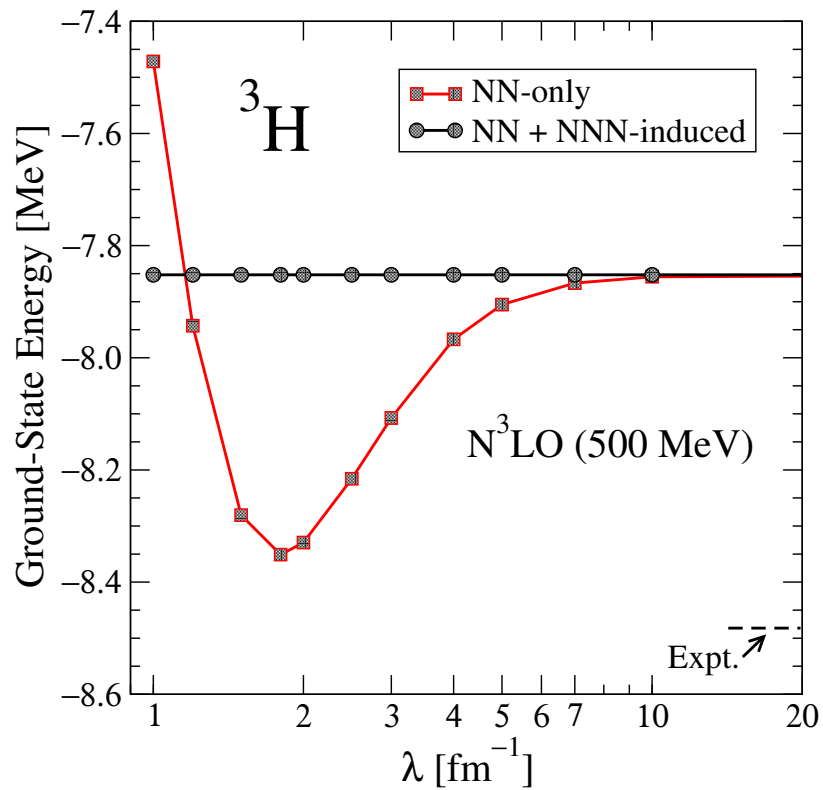


1) SRG-evolve both NN and 3N: NN+3N-full

2) NN Vlowk, refit 3N: NN+3N-fit

Induced 3N Forces

Effect of including 3N-ind? Exactly initial V_{NN} up to neglected 4N-ind

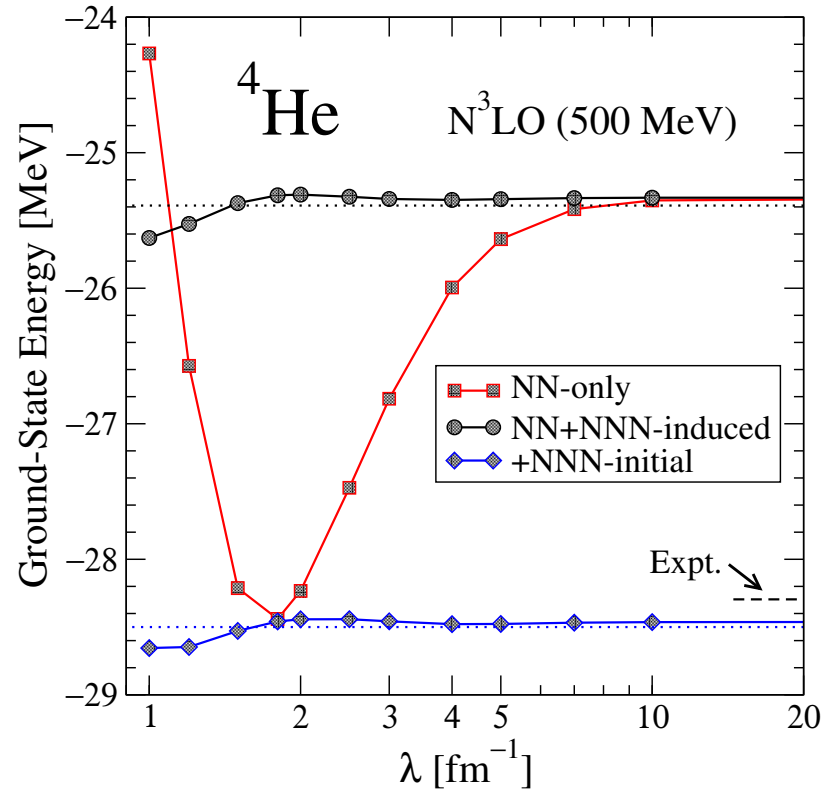
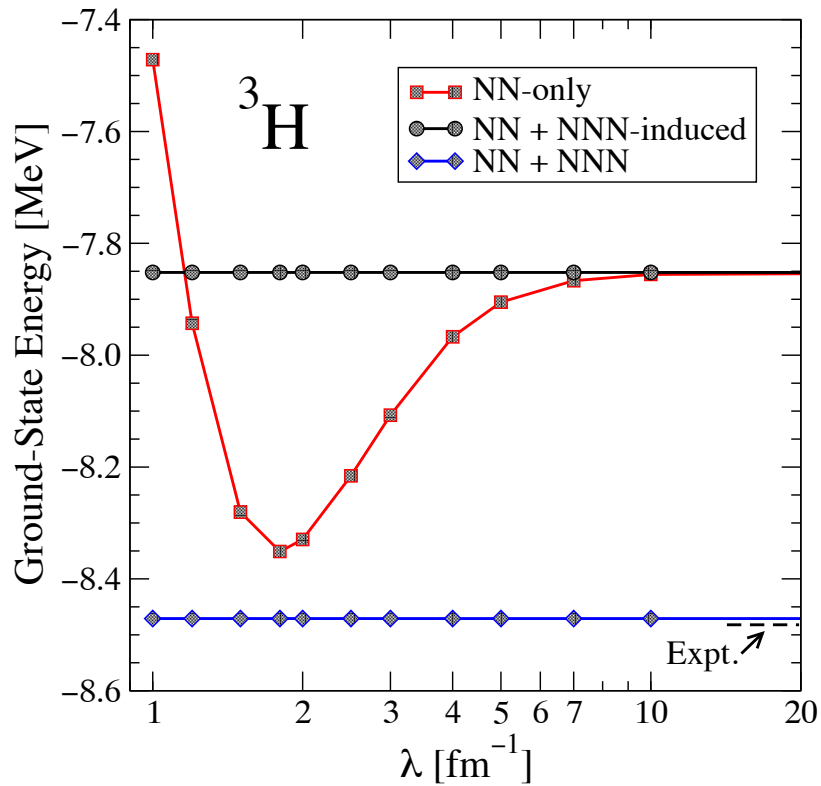


NN-only clear cutoff dependences

3N-ind: dramatic reduction in cutoff dependence, no agreement with experiment

Induced 3N Forces

Effect of including 3N-ind? Exactly initial V_{NN} up to neglected 4N-ind



NN-only clear cutoff dependences

3N-ind: dramatic reduction in cutoff dependence, no agreement with experiment

NN+3N-full retains cutoff independence, reproduces experiment!

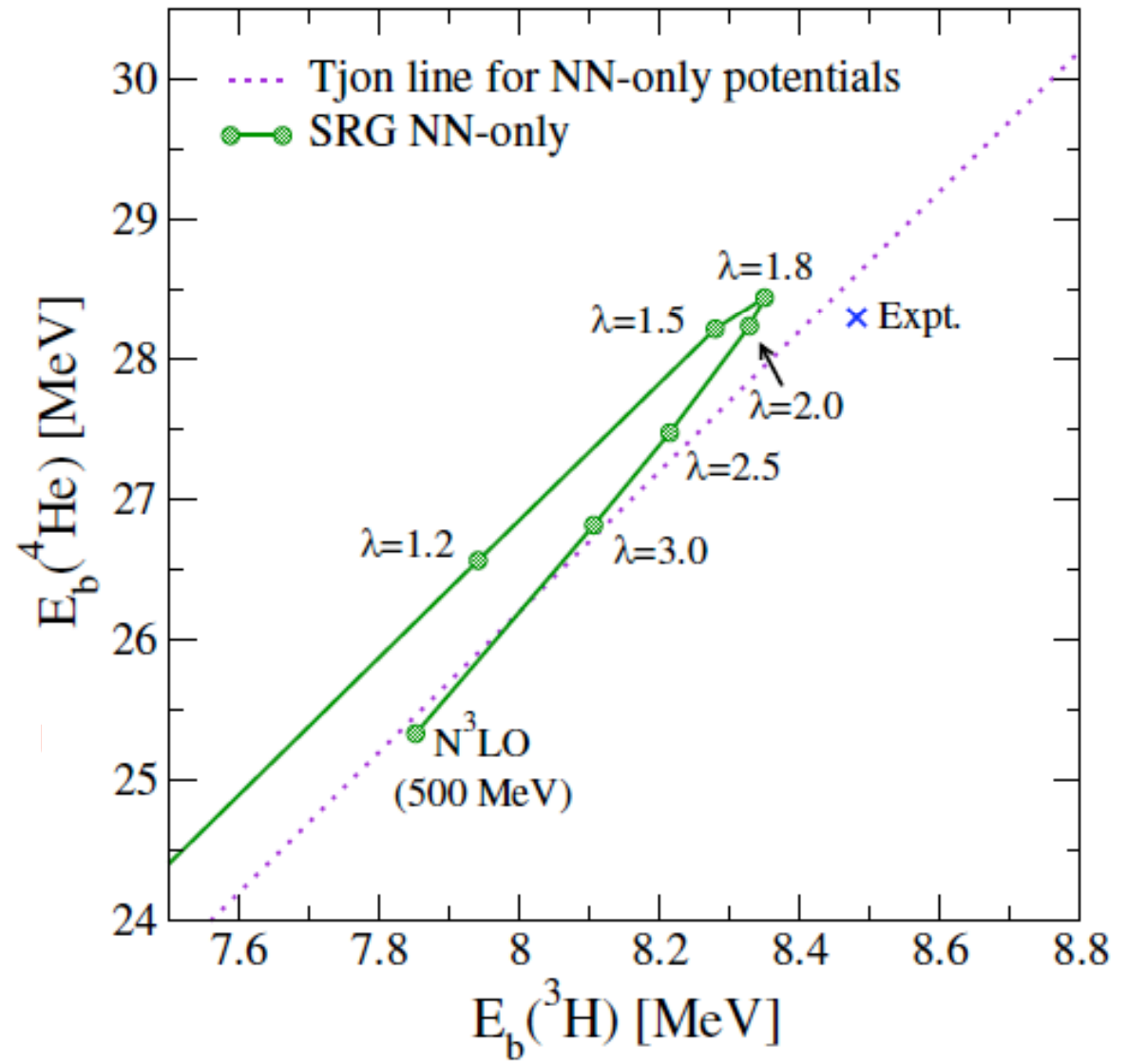
Benefits of Lower Cutoffs

Use cutoff dependence to assess missing physics: return to Tjon line

Varying cutoff moves along line

Still never reaches experiment

Tool, not a parameter!



Benefits of Lower Cutoffs

Use cutoff dependence to assess missing physics: return to Tjon line

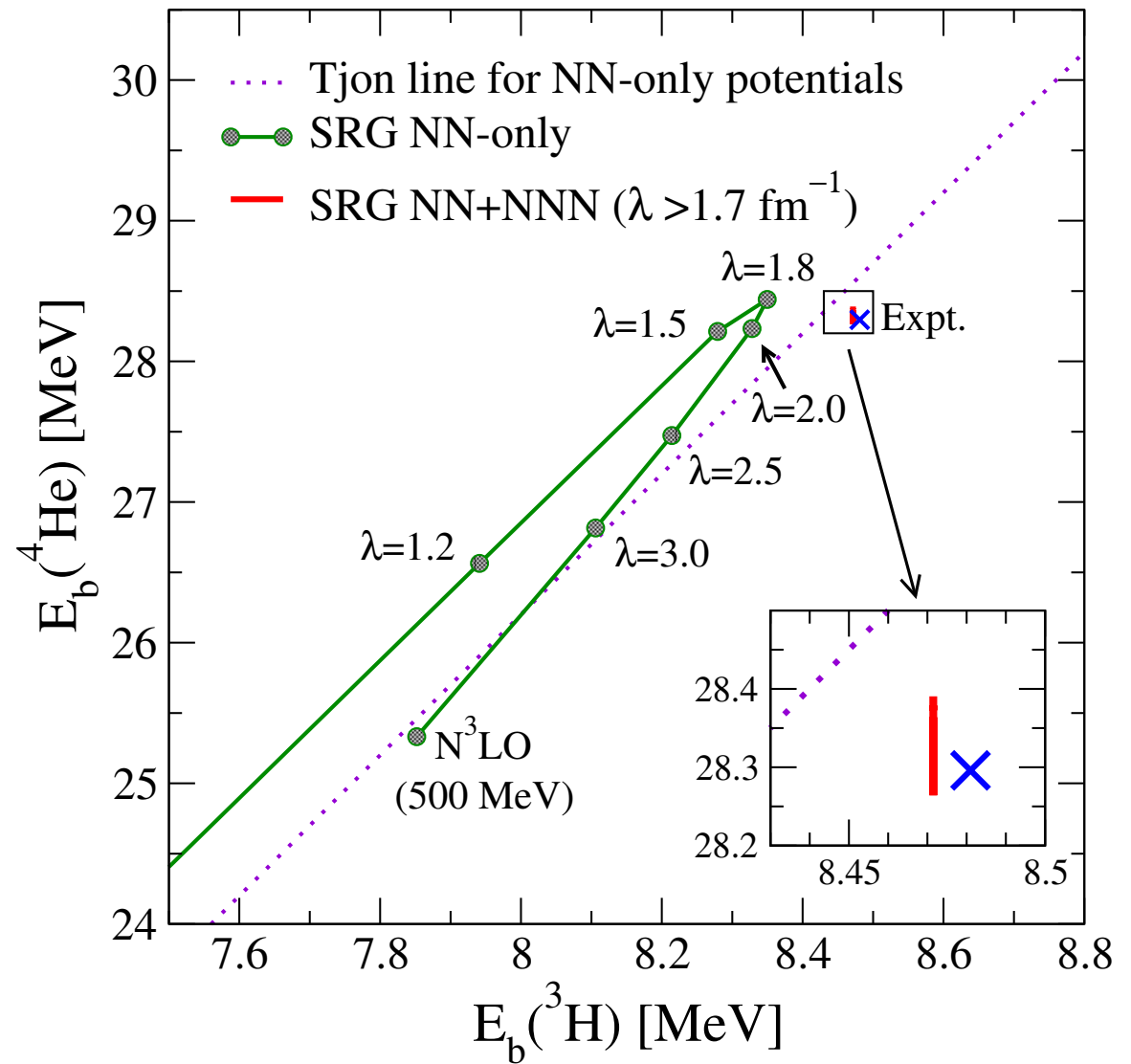
Varying cutoff moves along line

Still never reaches experiment

Tool, not a parameter!

Including 3N reaches expt.

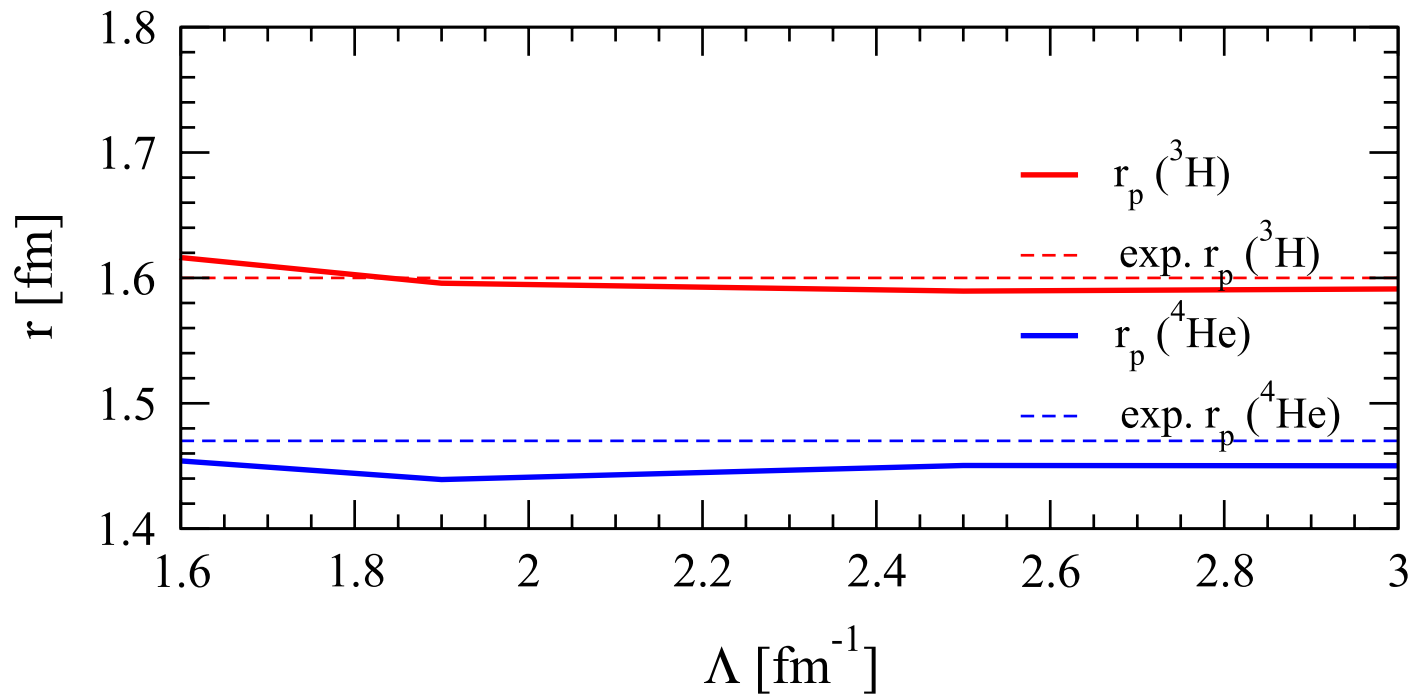
Why not perfect fit?



Cutoff Variation with 3N Forces

Use cutoff variation to assess missing physics in few body systems

Radii of triton and alpha particle calculated from NN+3N forces

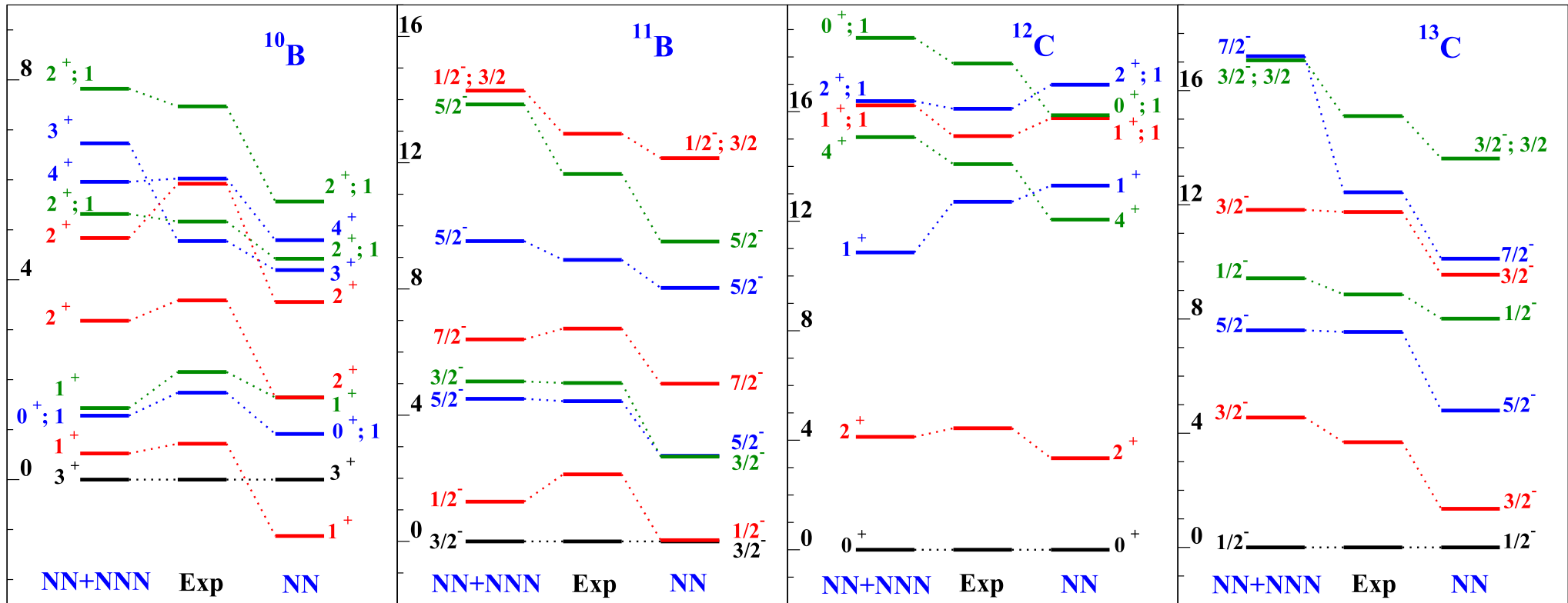


Minimal cutoff variation

Chiral Three-Body Forces in Light Nuclei

Importance of chiral 3N forces established in light nuclei

Converged NCSM (Navratil 2007)

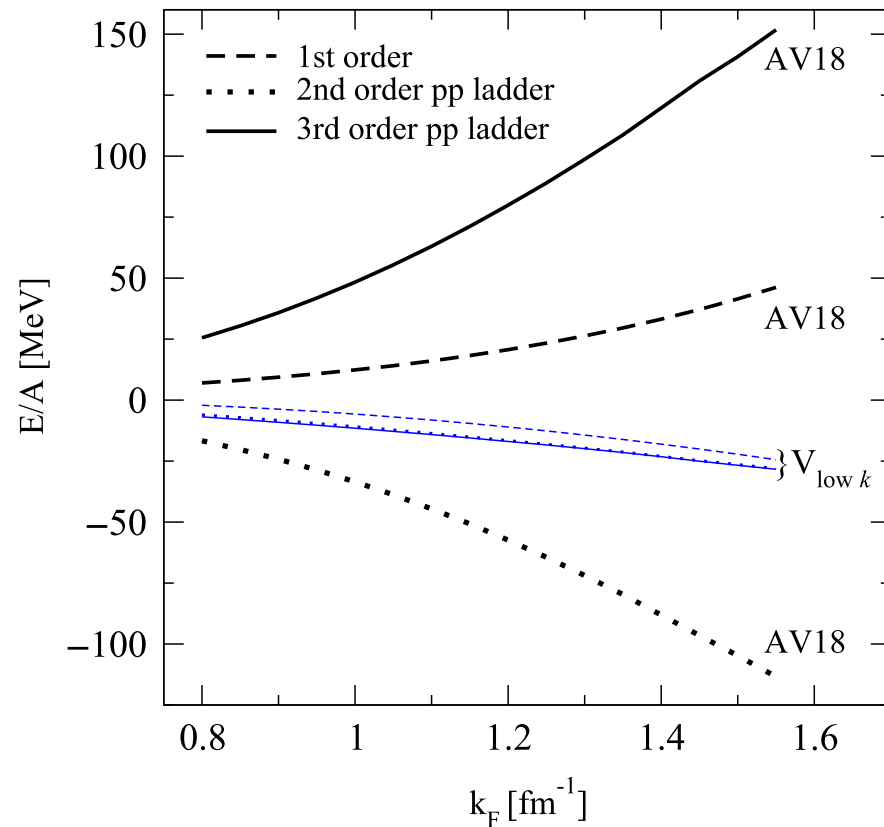
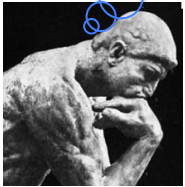


They work! What about nuclear matter?

Perturbative in Symmetric Nuclear Matter?

$$H(\Lambda) = T + V_{\text{NN}}(\Lambda) + V_{\text{3N}}(\Lambda) + V_{\text{4N}}(\Lambda) + \dots$$

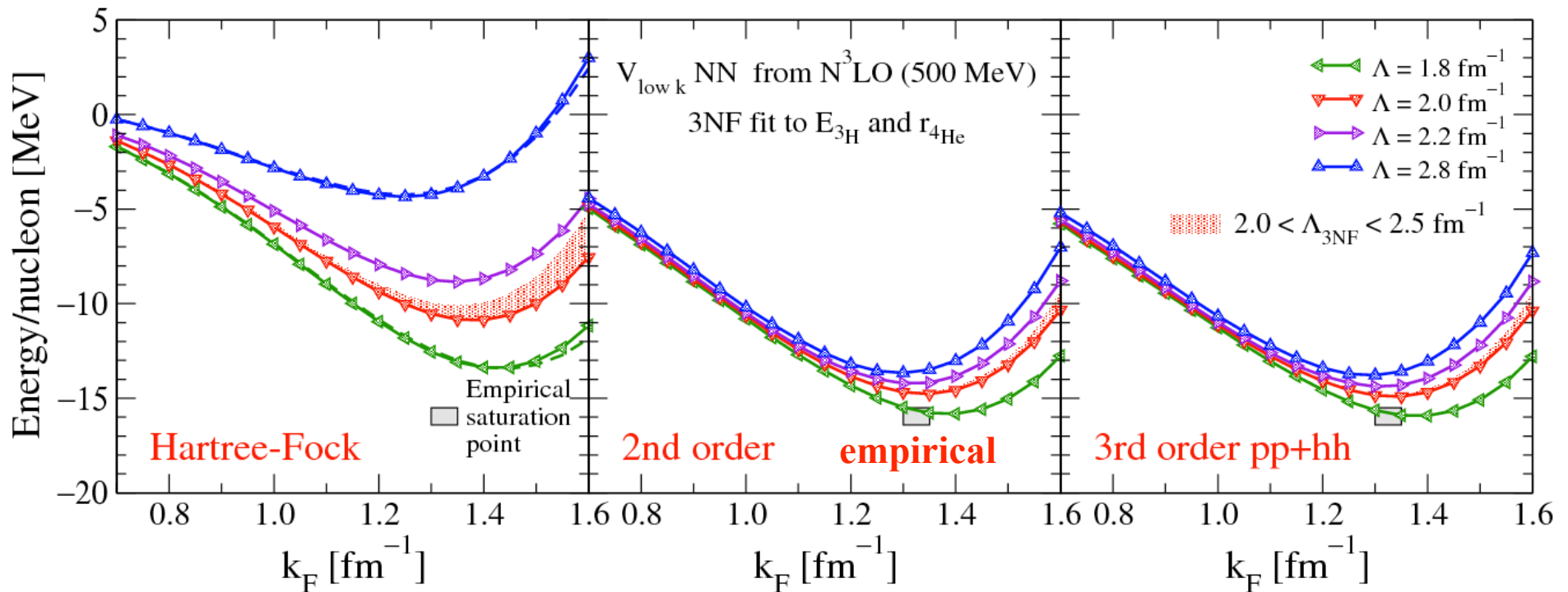
Yes, but if I
remember, saturation
isn't correct



Significant improvement with low-momentum interactions!

Perturbative in Symmetric Nuclear Matter?

$$H(\Lambda) = T + V_{\text{NN}}(\Lambda) + V_{\text{3N}}(\Lambda) + V_{\text{4N}}(\Lambda) + \dots$$



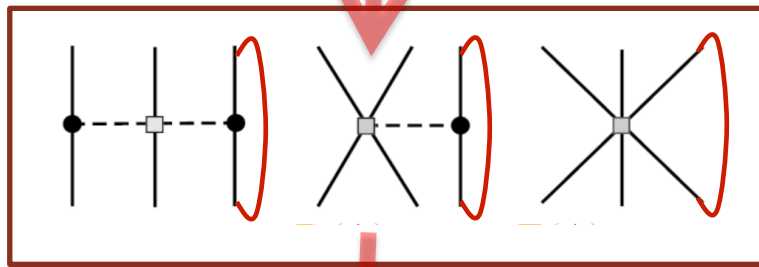
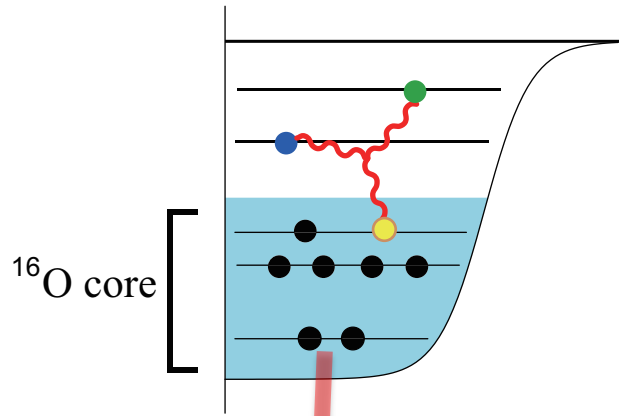
Now NN+3N-fit remain perturbative and reproduce saturation!

Minor but non-negligible cutoff variation

3N Forces for Valence-Shell Theories

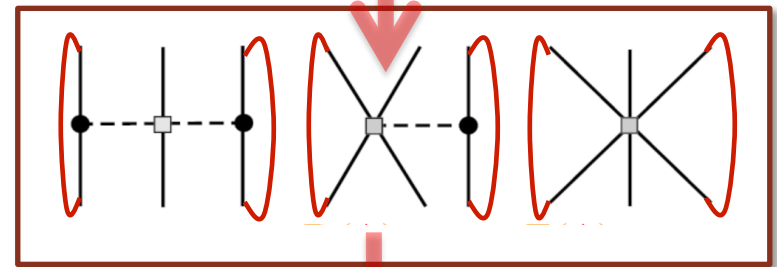
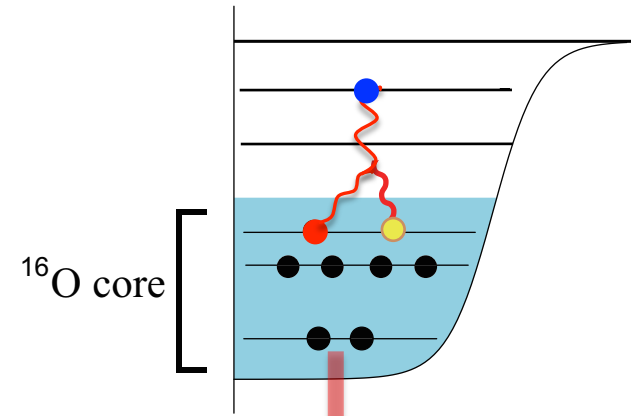
Normal-ordered 3N: contribution to valence neutron interactions

Effective two-body



$$\langle ab | V_{3N,\text{eff}} | a'b' \rangle = \sum_{\alpha=\text{core}} \langle \alpha ab | V_{3N} | \alpha a'b' \rangle$$

Effective one-body



$$\langle a | V_{3N,\text{eff}} | a' \rangle = \frac{1}{2} \sum_{\alpha\beta=\text{core}} \langle \alpha\beta a | V_{3N} | \alpha\beta a' \rangle$$

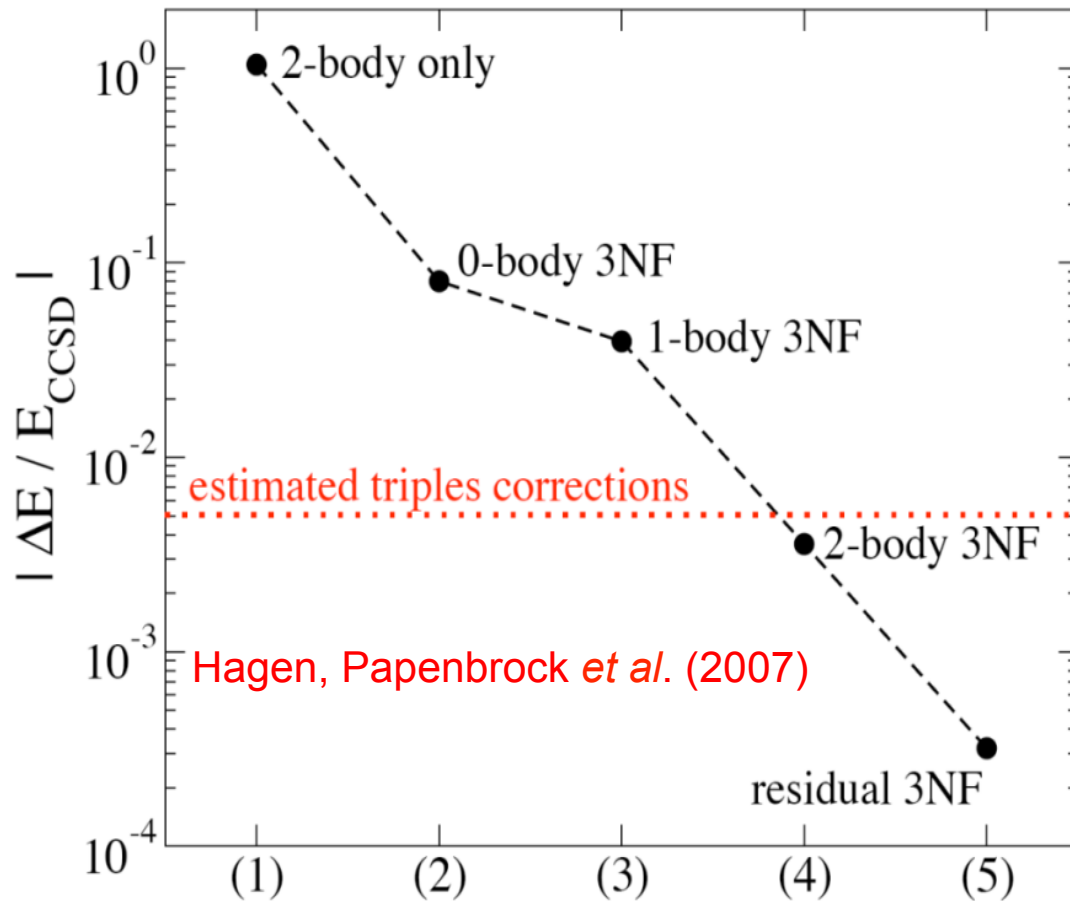
Combine with microscopic NN: eliminate empirical adjustments

3N Forces for Valence-Shell Theories

Effects of residual 3N between 3 valence nucleons?

Normal-ordered 3N: microscopic contributions to inputs for CI Hamiltonian

Effects of residual 3N between 3 valence nucleons?



Coupled-Cluster theory with 3N:
benchmark of ${}^4\text{He}$

0- 1- and 2-body of 3NF dominate

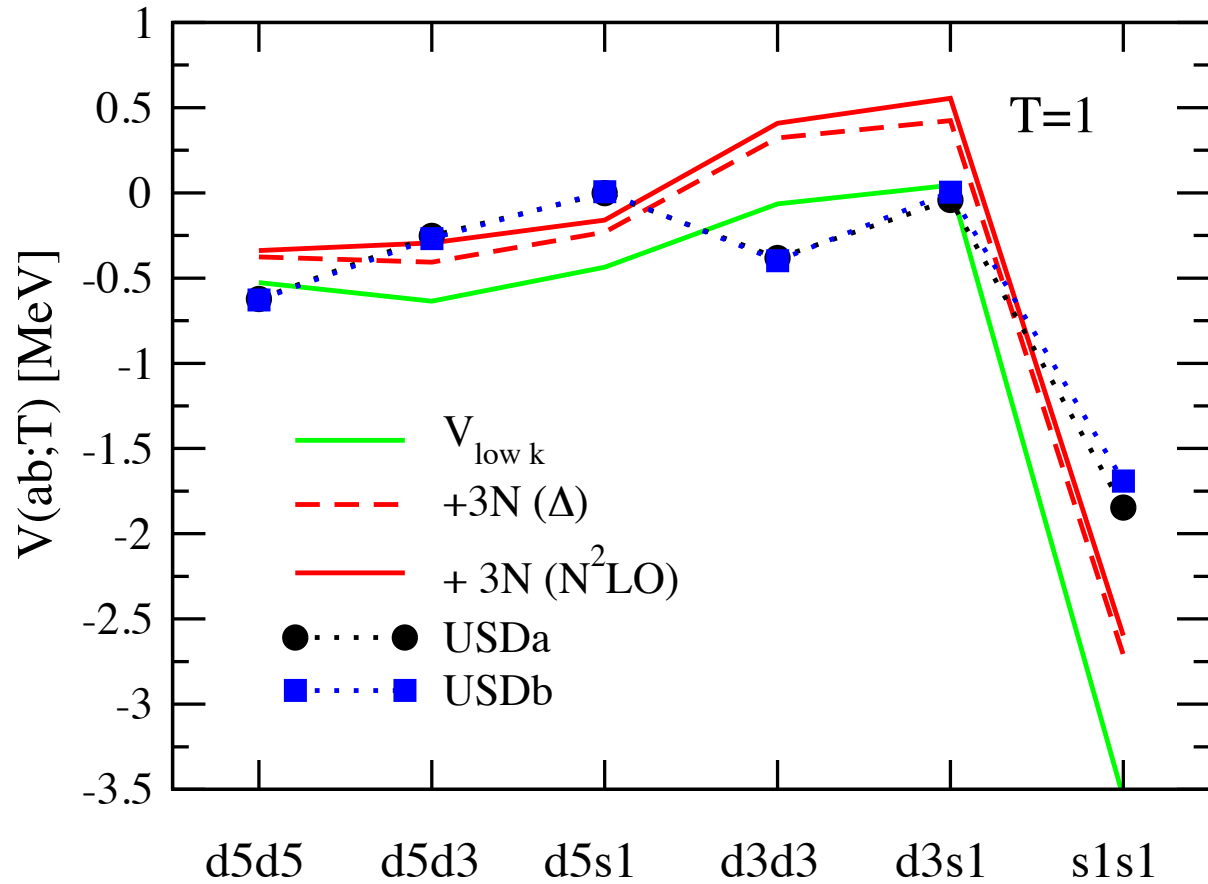
Residual 3N can be neglected

Work on ${}^{16}\text{O}$ in progress

Approximated residual 3N by summing over valence nucleon

– Nucleus-dependent: effect small, not negligible by ${}^{24}\text{O}$

Two-body 3N: Monopoles in *sd*-shell



Dominant effect from
one- Δ – as expected
from cutoff variation

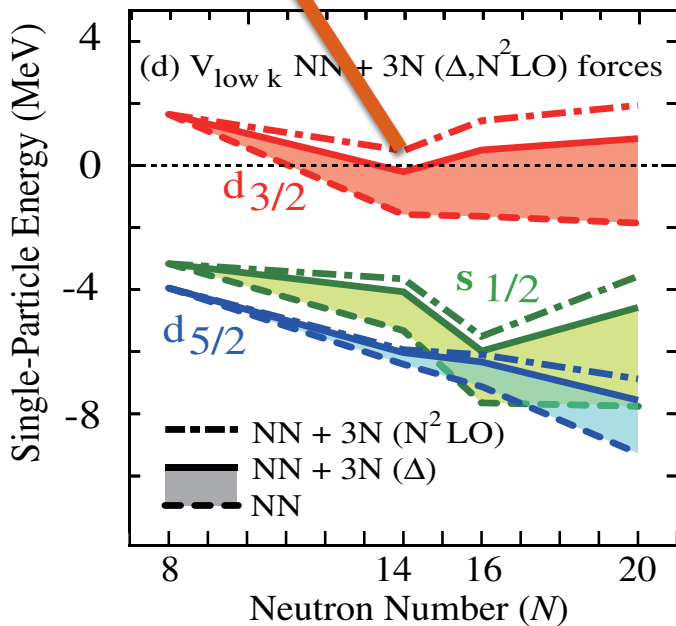
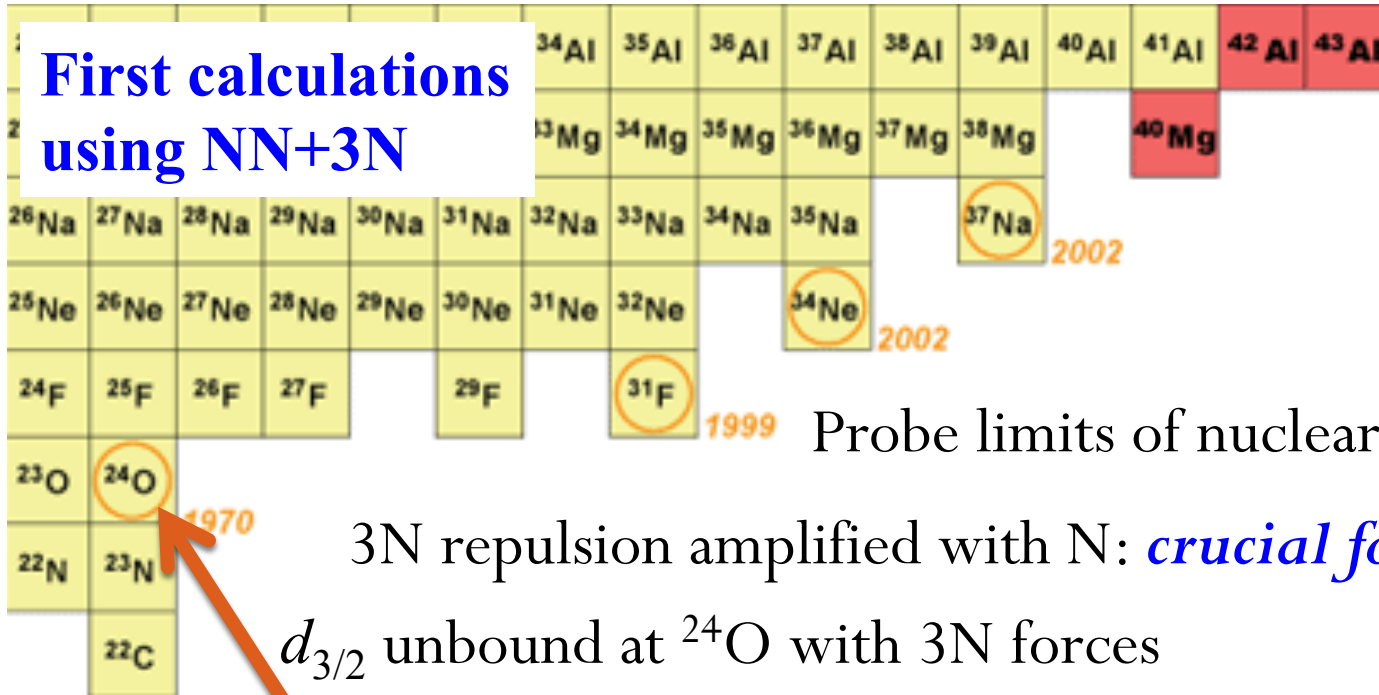
3N forces produce clear
repulsive shift in monopoles

First calculations to show missing monopole strength due to neglected 3N

Future: Improved treatment of high-lying orbits

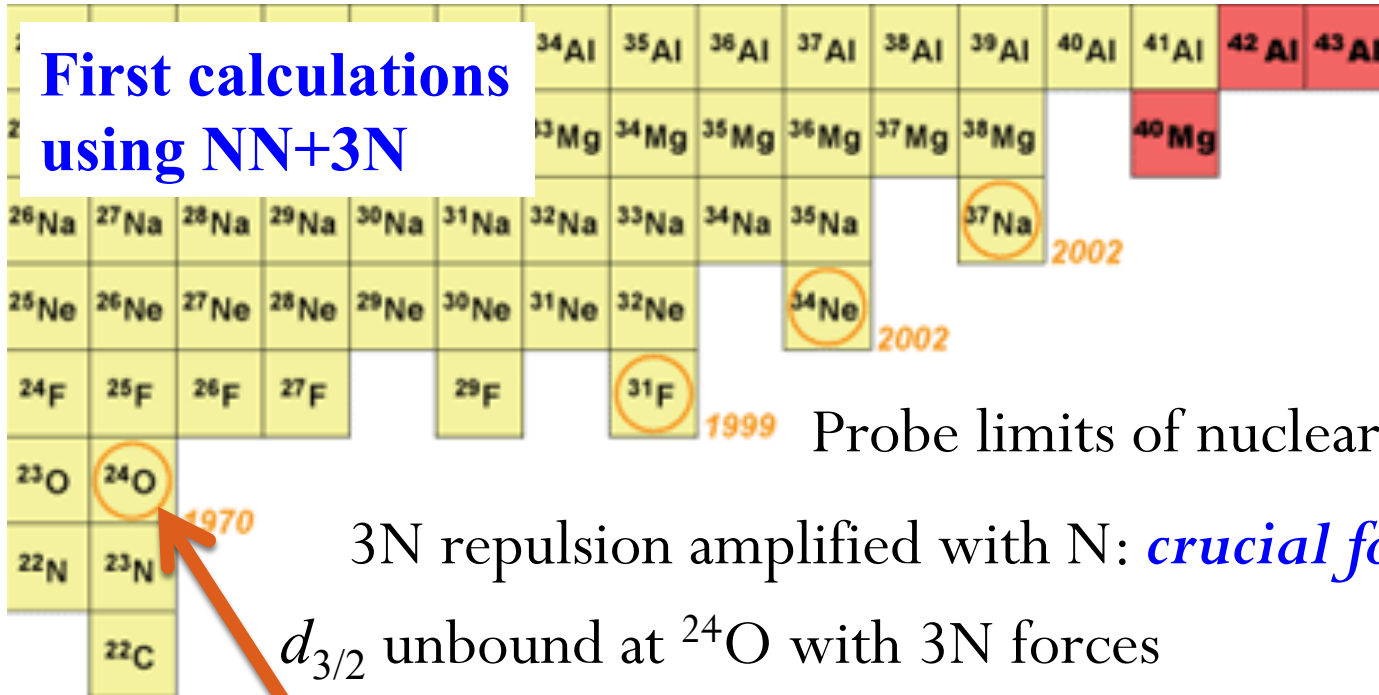
Oxygen Anomaly

First calculations using NN+3N



Oxygen Anomaly

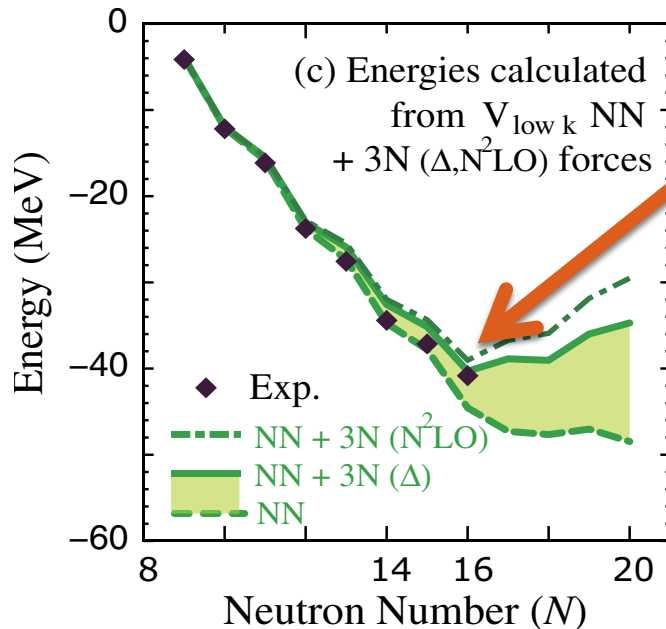
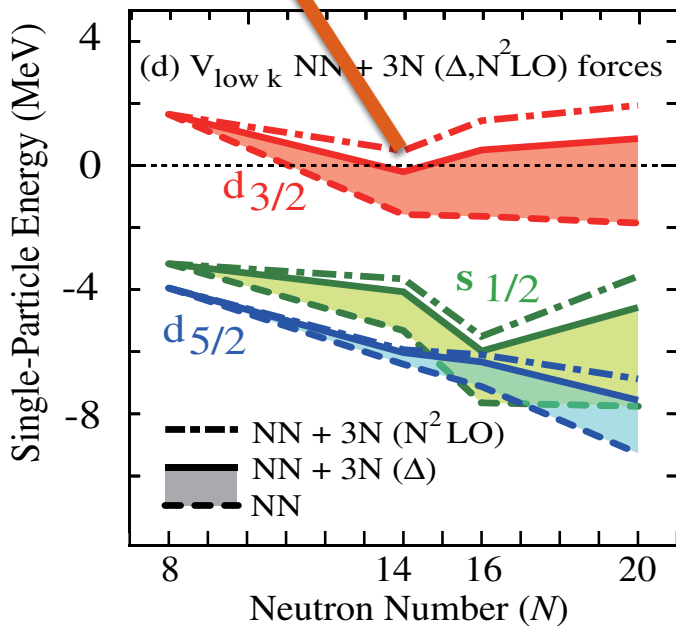
First calculations using NN+3N



Probe limits of nuclear existence with 3N forces

3N repulsion amplified with N: *crucial for neutron-rich nuclei*

$d_{3/2}$ unbound at ^{24}O with 3N forces

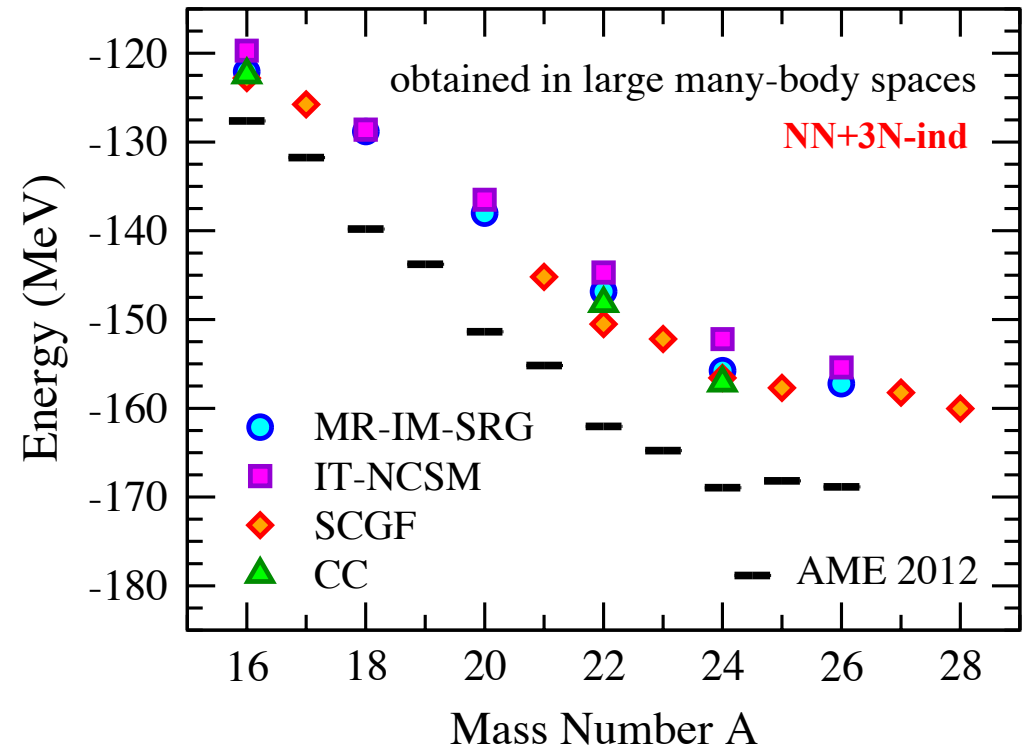
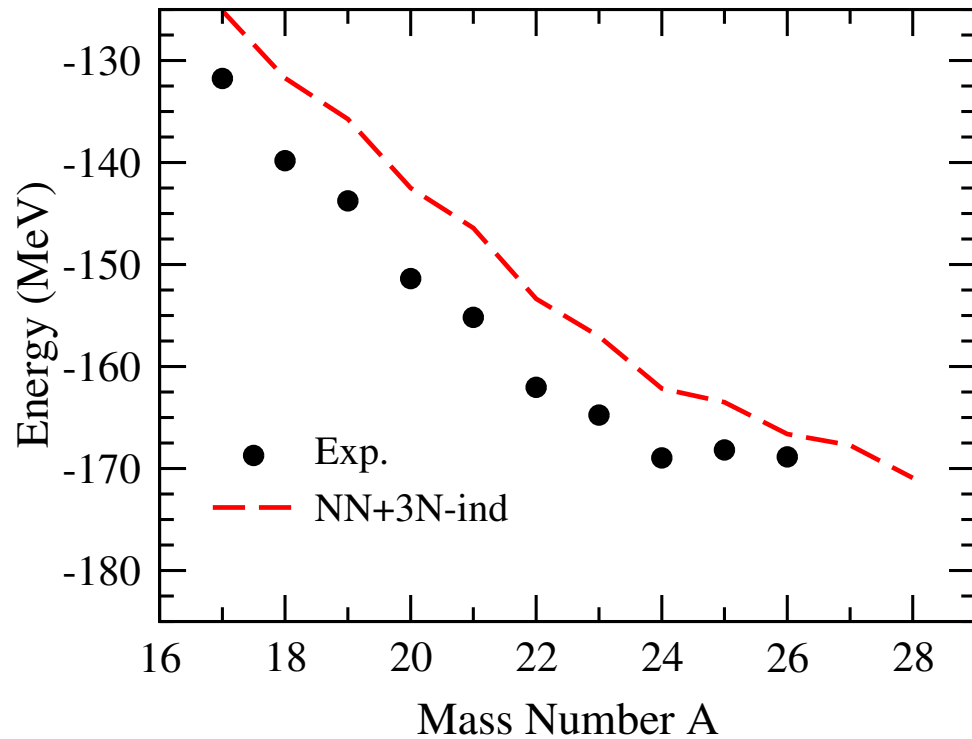


Isotopes unbound beyond ^{24}O

First microscopic explanation of oxygen anomaly

Comparison with Large-Space Methods

Large-space methods with **same SRG-evolved NN+3N-ind forces**



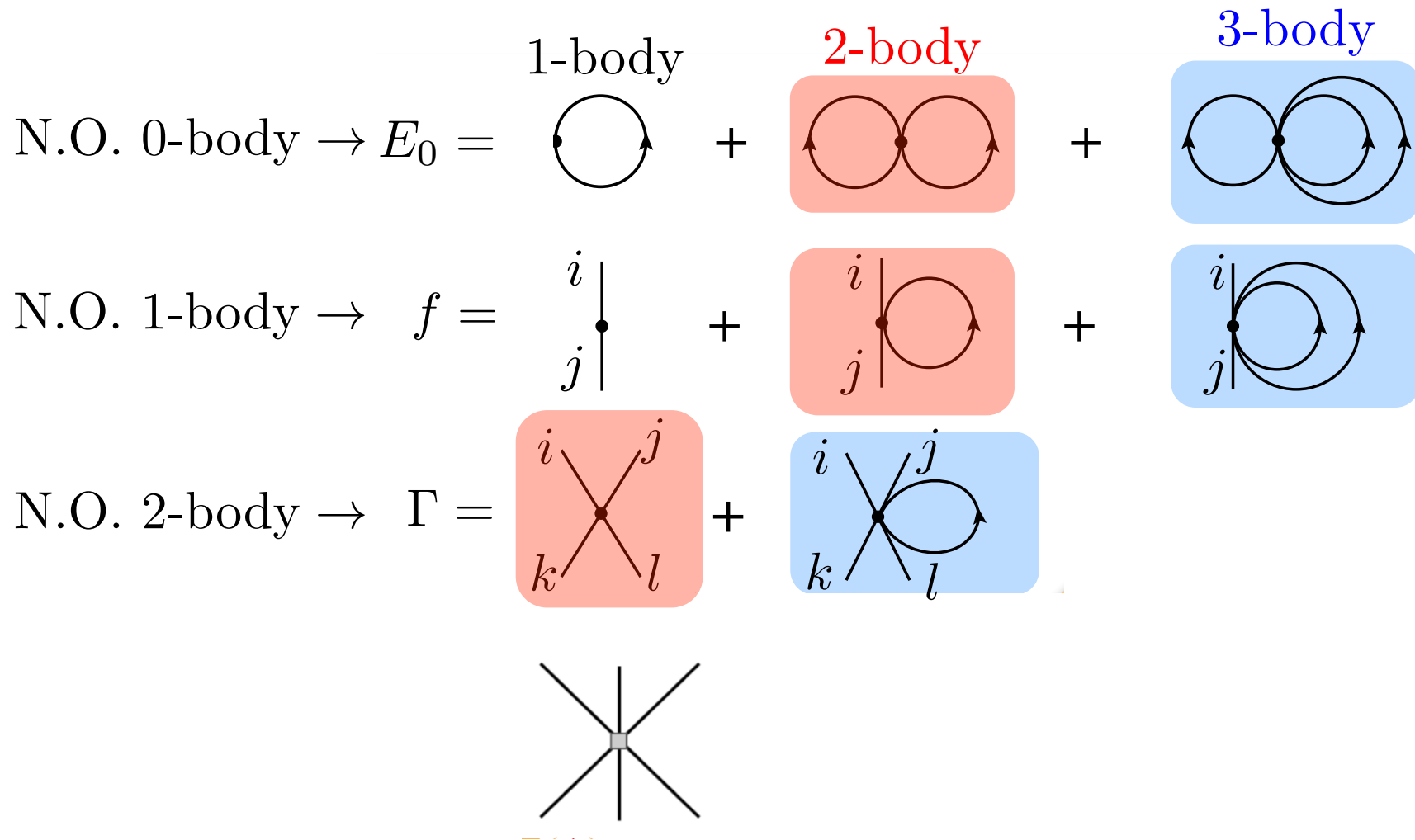
Agreement between all methods with same input forces

No reproduction of dripline in any case

Normal-Ordered Hamiltonian

Now rewrite exactly the initial Hamiltonian in normal-ordered form

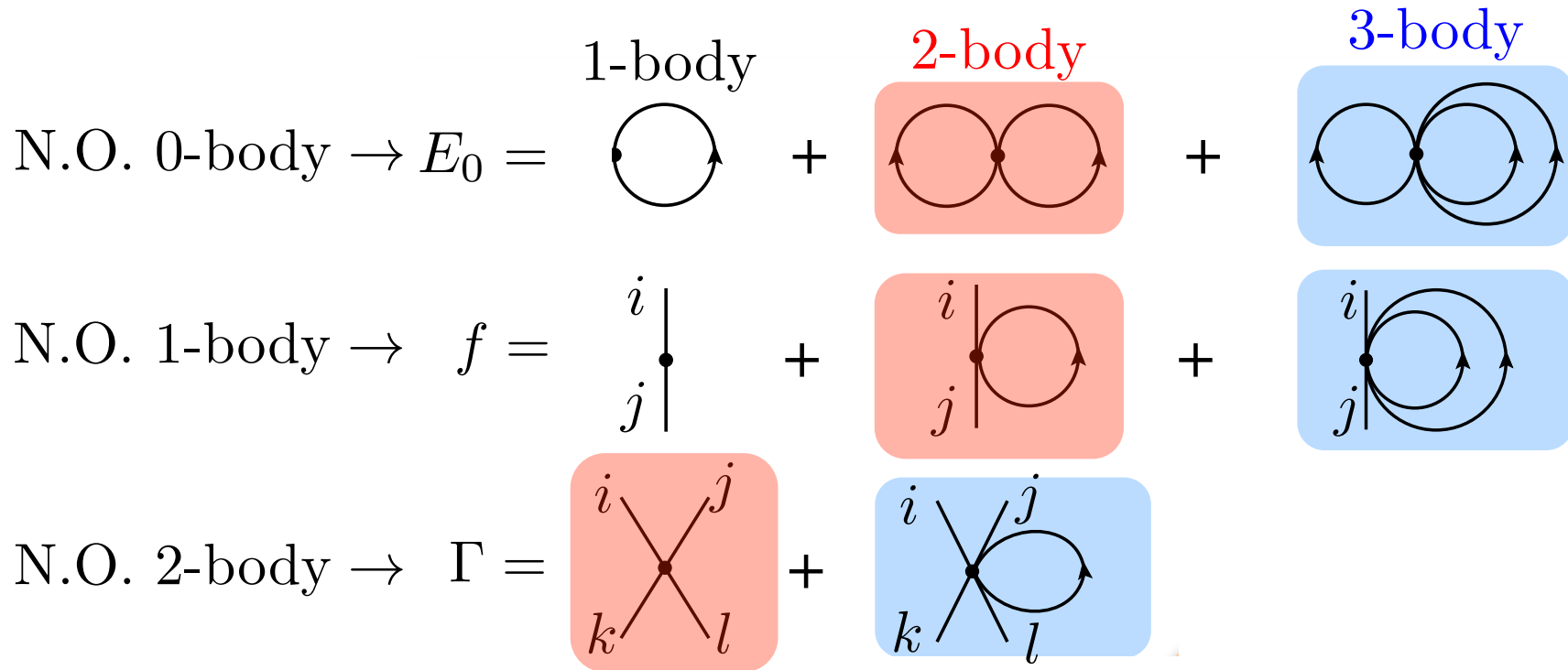
$$H_{\text{N.O.}} = E_0 + \sum_{ij} f_{ij} \{a_i^\dagger a_j\} + \frac{1}{4} \sum_{ijkl} \Gamma_{ijkl} \{a_i^\dagger a_j^\dagger a_l a_k\} + \frac{1}{36} \sum_{ijklmn} W_{ijklmn} \{a_i^\dagger a_j^\dagger a_k^\dagger a_l a_m a_n\}$$



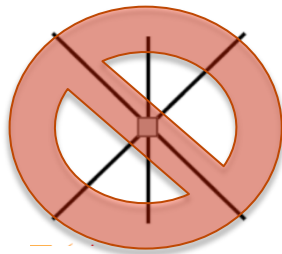
Normal-Ordered Hamiltonian

Now rewrite exactly the initial Hamiltonian in normal-ordered form

$$H_{\text{N.O.}} = E_0 + \sum_{ij} f_{ij} \{a_i^\dagger a_j\} + \frac{1}{4} \sum_{ijkl} \Gamma_{ijkl} \{a_i^\dagger a_j^\dagger a_l a_k\} + \frac{1}{36} \sum_{ijklmn} W_{ijklmn} \{a_i^\dagger a_j^\dagger a_k^\dagger a_l a_m a_n\}$$

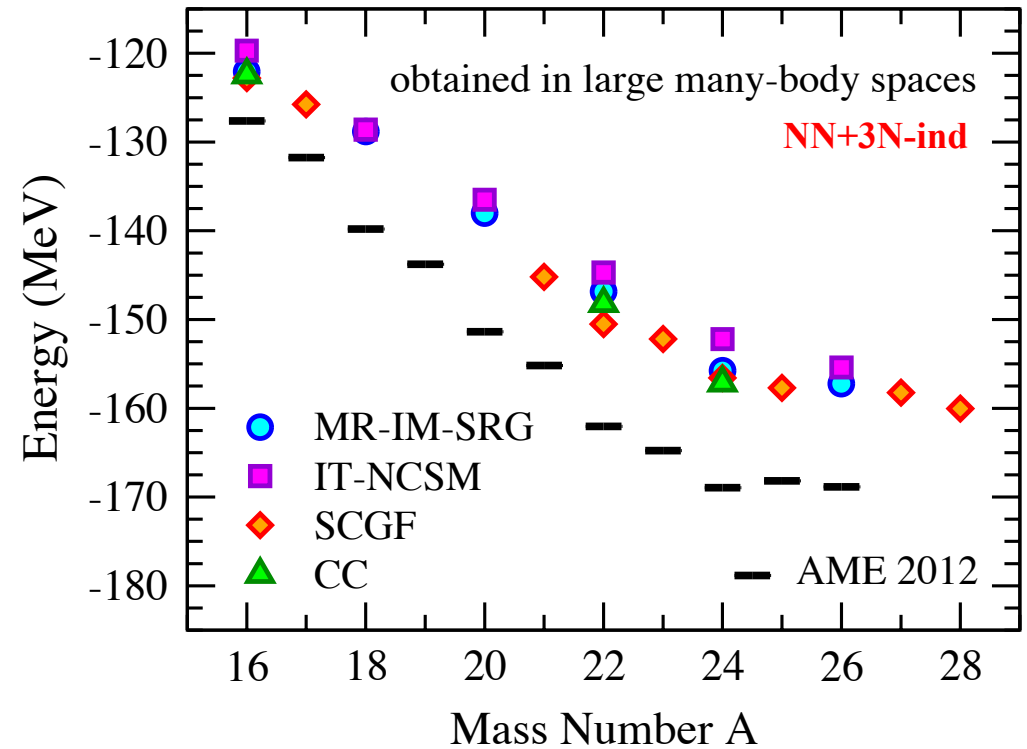
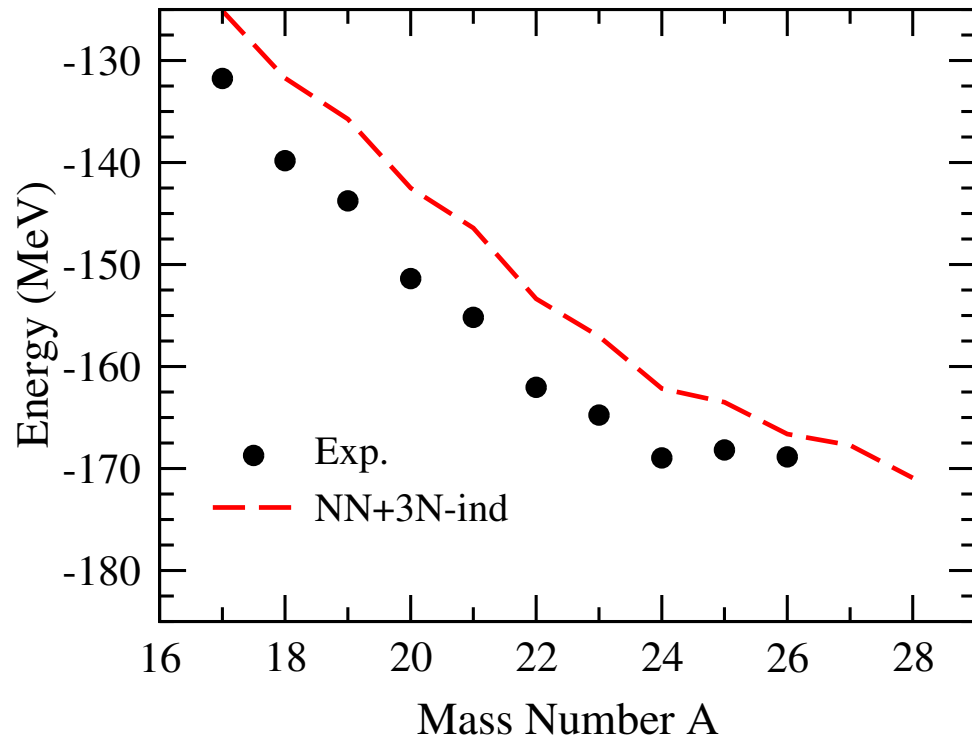


Neglect residual 3N



Comparison with Large-Space Methods

Large-space methods with **same SRG-evolved NN+3N-ind forces**

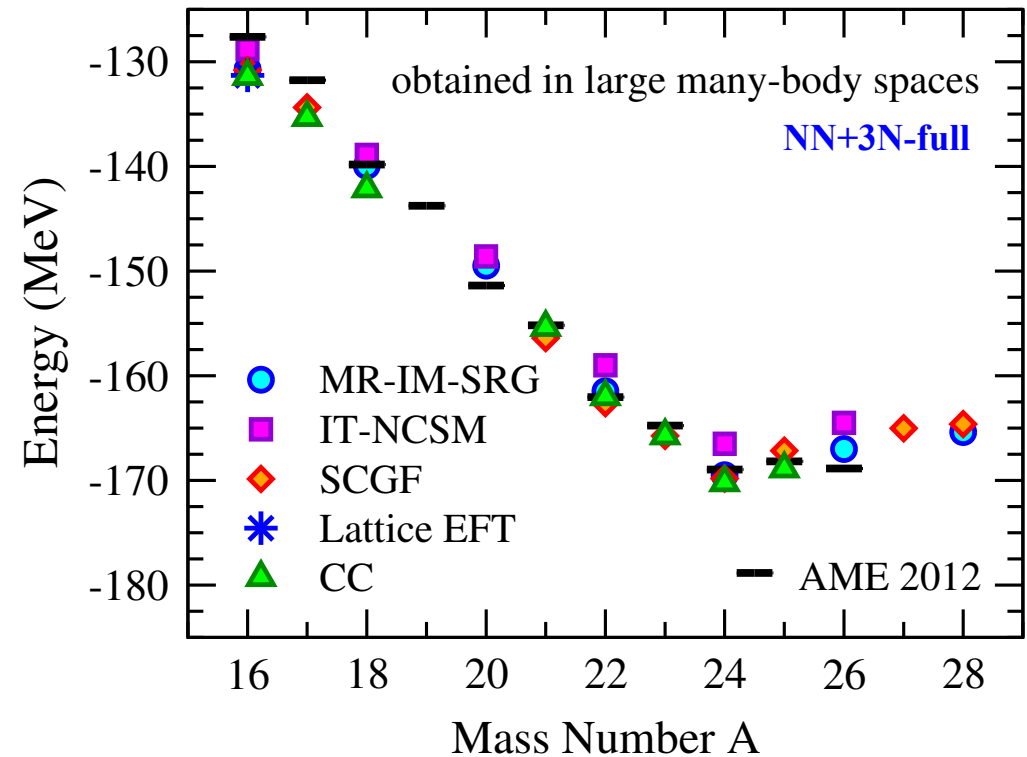
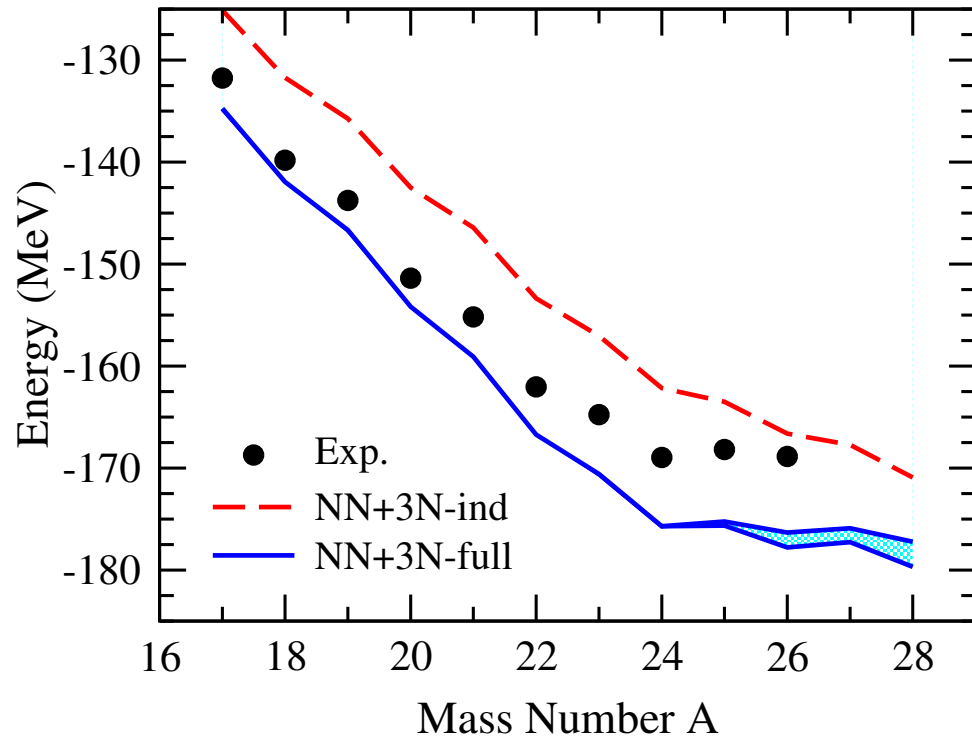


Agreement between all methods with same input forces

No reproduction of dripline in any case

Comparison with Large-Space Methods

Large-space methods with **same SRG-evolved NN+3N-full forces**



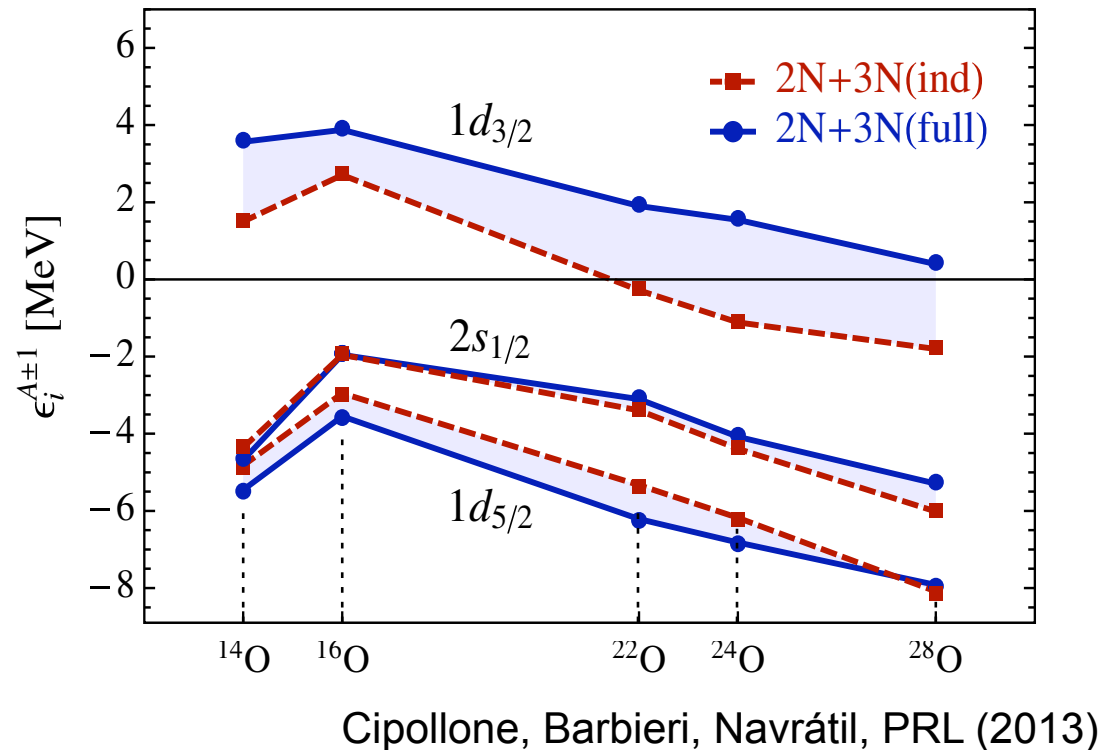
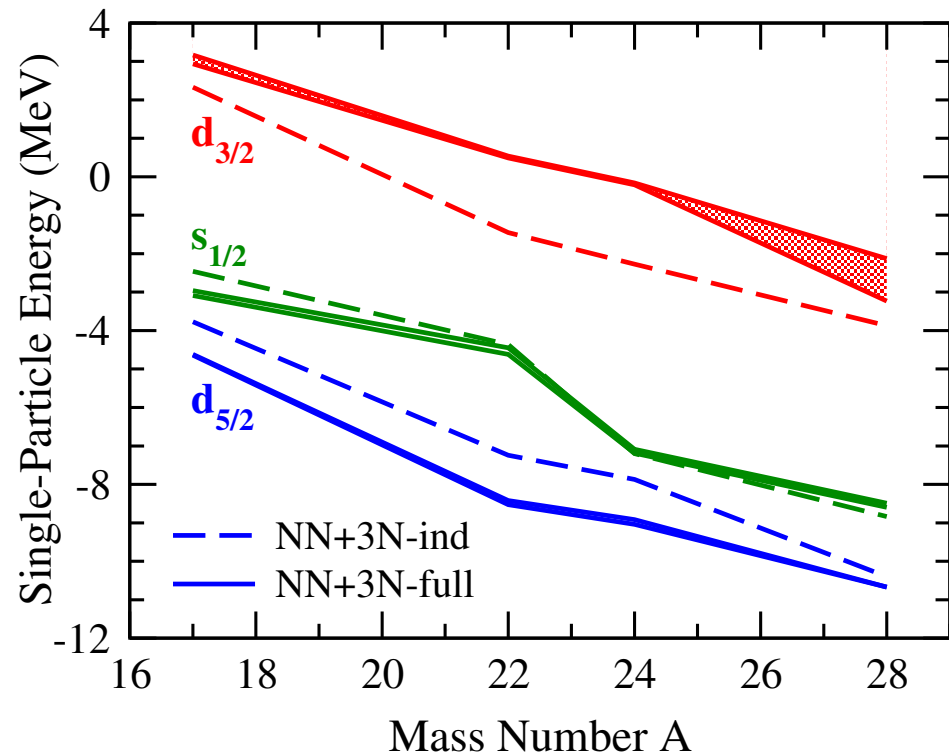
Agreement between all methods with same input forces

Clear improvement with NN+3N-full

Validates valence-space results

Oxygen Dripline Mechanism

Self-consistent Green's Function with **same SRG-evolved NN+3N forces**



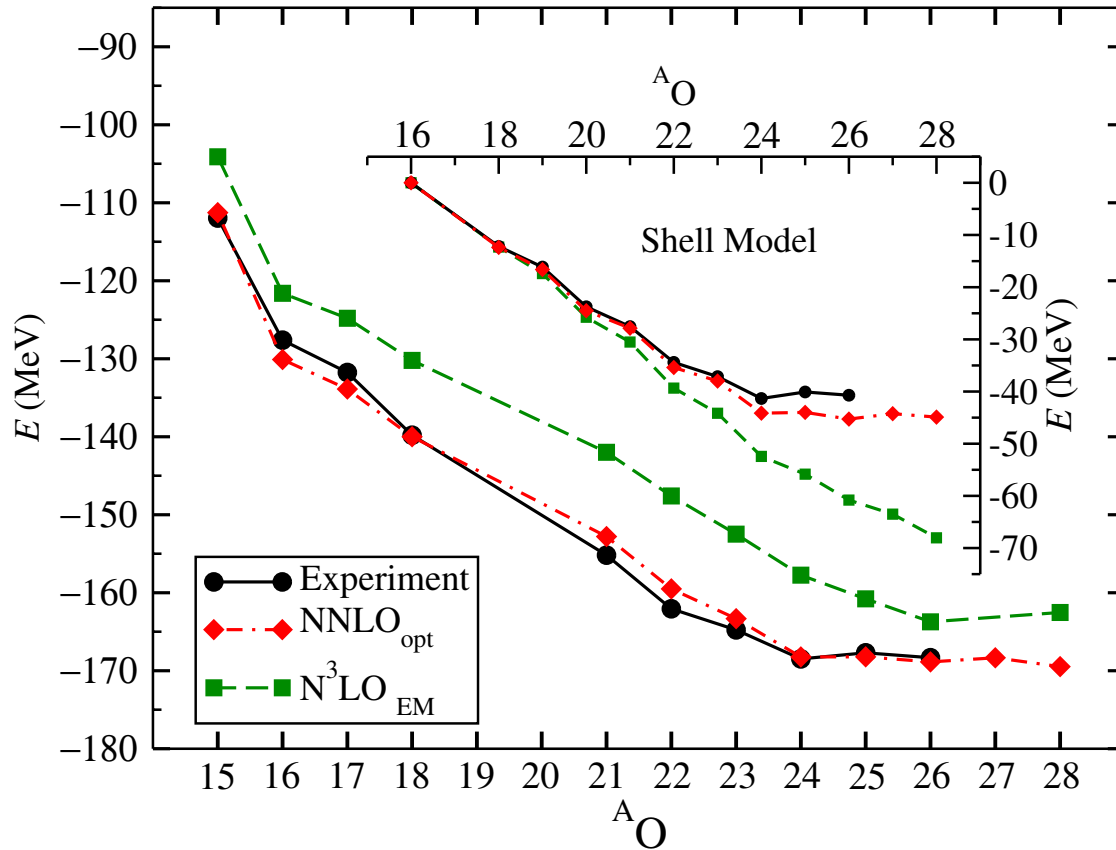
Robust mechanism driving dripline behavior

3N repulsion raises $d_{3/2}$, lessens decrease across shell

Similar to first MBPT NN+3N calculations in oxygen

Optimized Chiral Forces N²LO NN-Only

Recent calculations at N²LO without 3N forces found a remarkable result



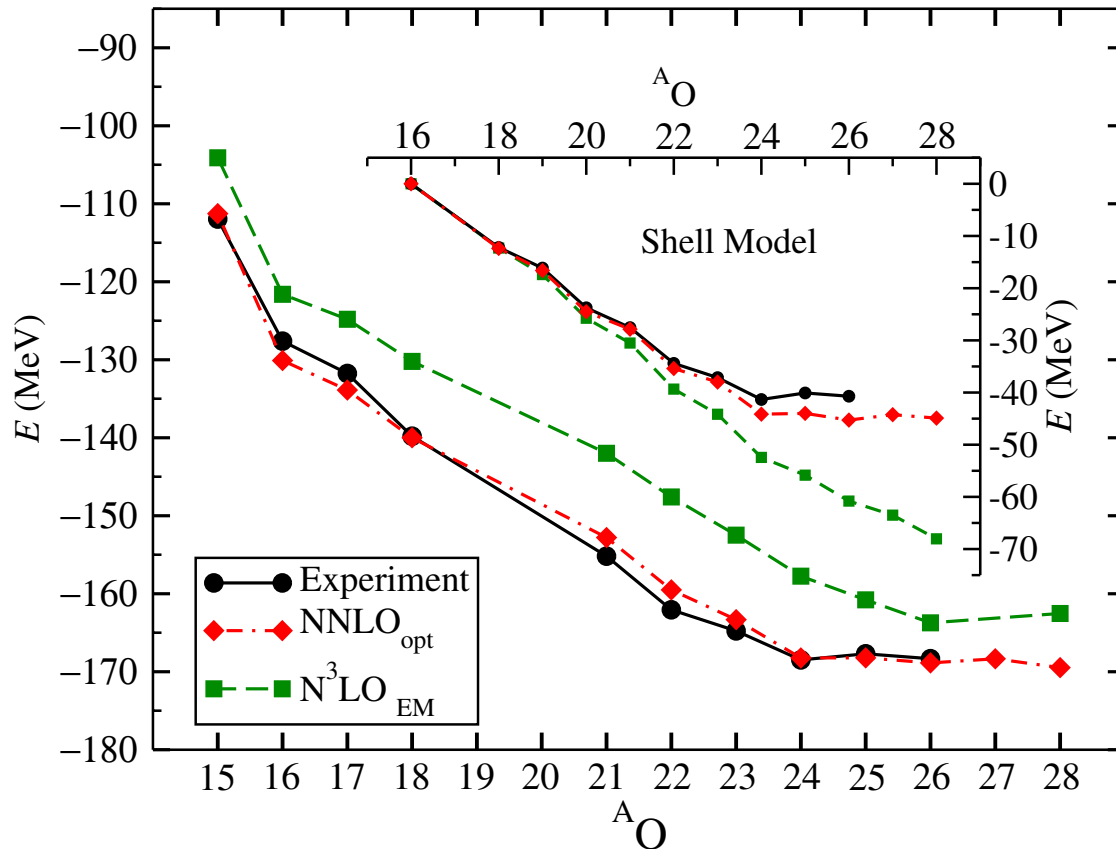
Ekström et al (PRL 2013)

Oxygen dripline reproduced with NN forces only!

What does this mean about 3N?

Optimized Chiral Forces N²LO NN-Only

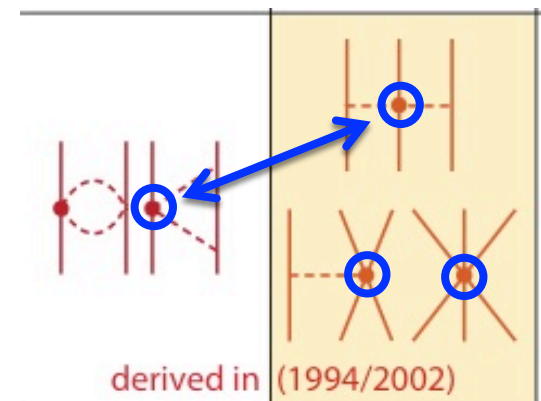
Recent calculations at N²LO without 3N forces found a remarkable result



Ekström et al (PRL 2013)

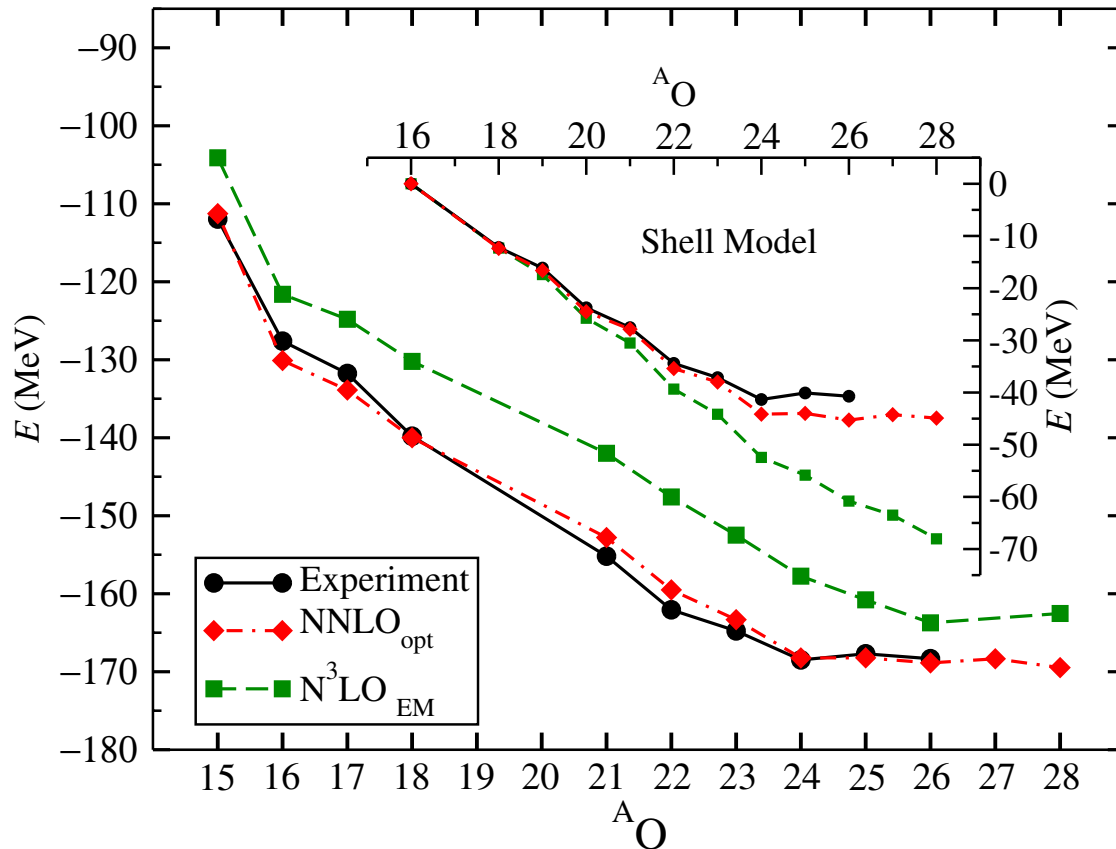
Oxygen dripline reproduced with NN forces only!

Power counting dictates 3N forces be included



Optimized Chiral Forces N²LO NN-Only

Recent calculations at N²LO without 3N forces found a remarkable result



Ekström et al (PRL 2013)

Oxygen dripline reproduced with NN forces only

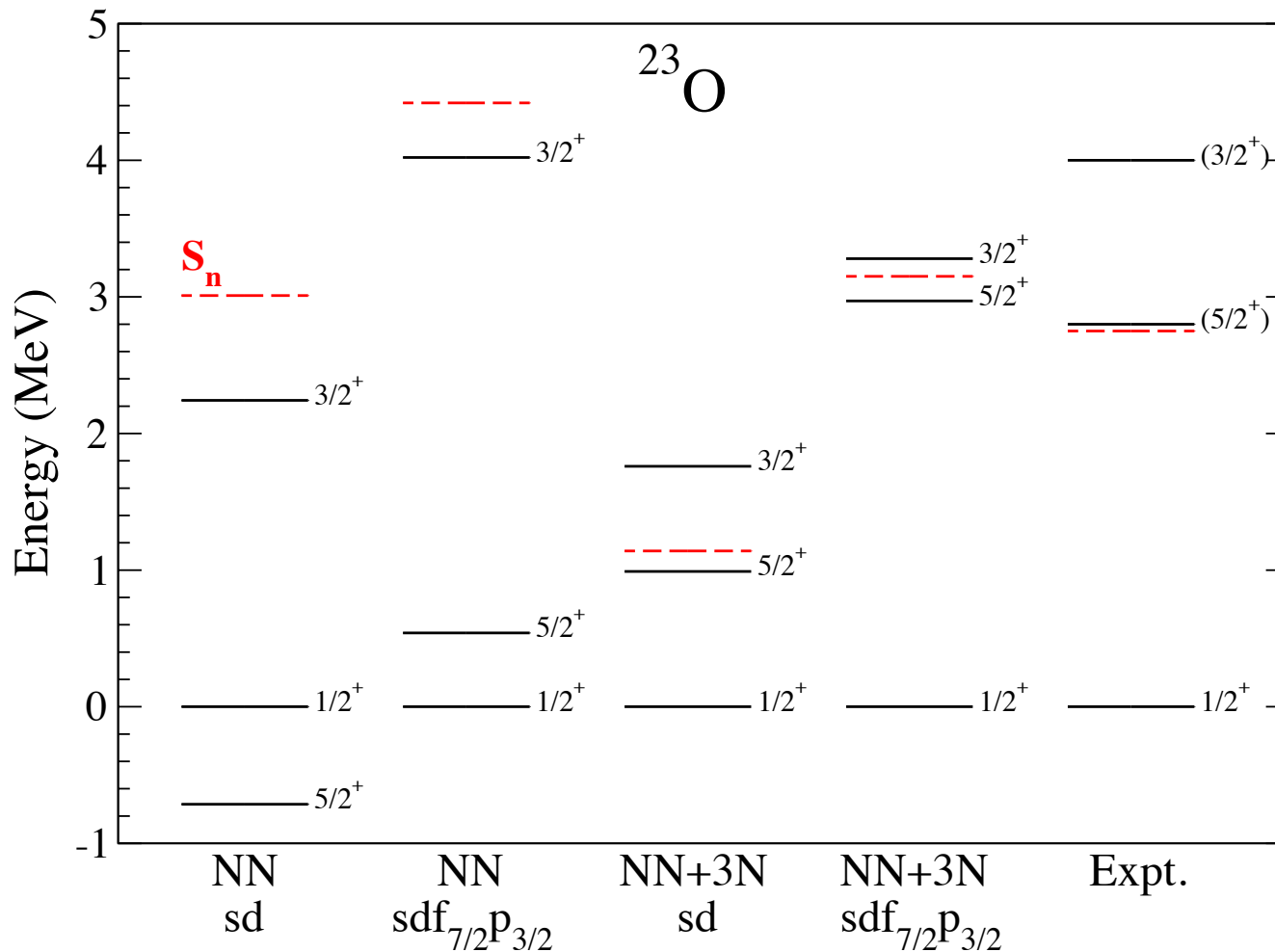
Unnaturally large couplings when 3N fit in ³H(?) – results off the plot!

Lesson: 3N forces unavoidable part of theory – must investigate importance

Impact on Spectra: ^{23}O

Neutron-rich oxygen spectra with NN+3N

$5/2^+$, $3/2^+$ energies reflect $^{22,24}\text{O}$ shell closures



sd-shell NN only

Wrong ground state

$5/2^+$ too low

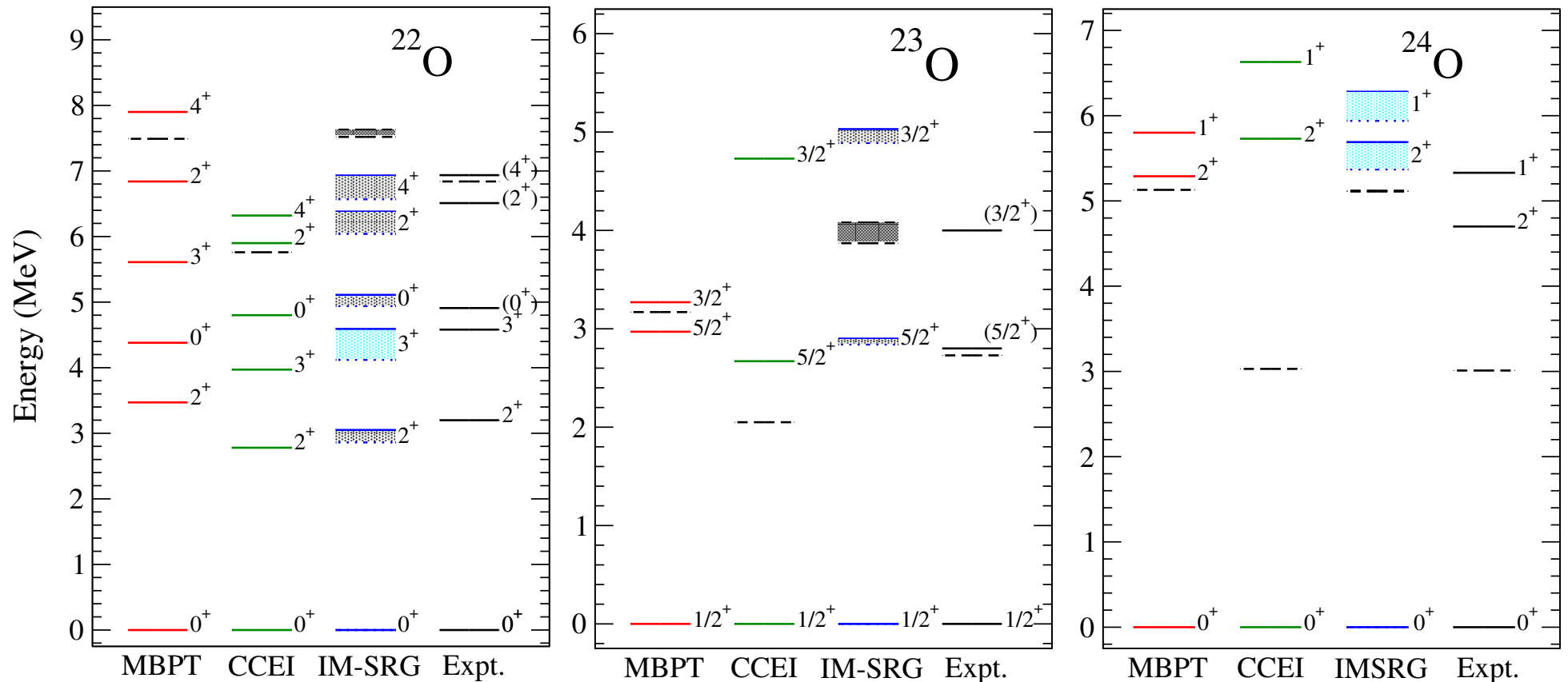
$3/2^+$ bound

NN+3N

Clear improvement in
extended valence space

Comparison with MBPT/CCEI Oxygen Spectra

Oxygen spectra: Effective interactions from **Coupled-Cluster theory**

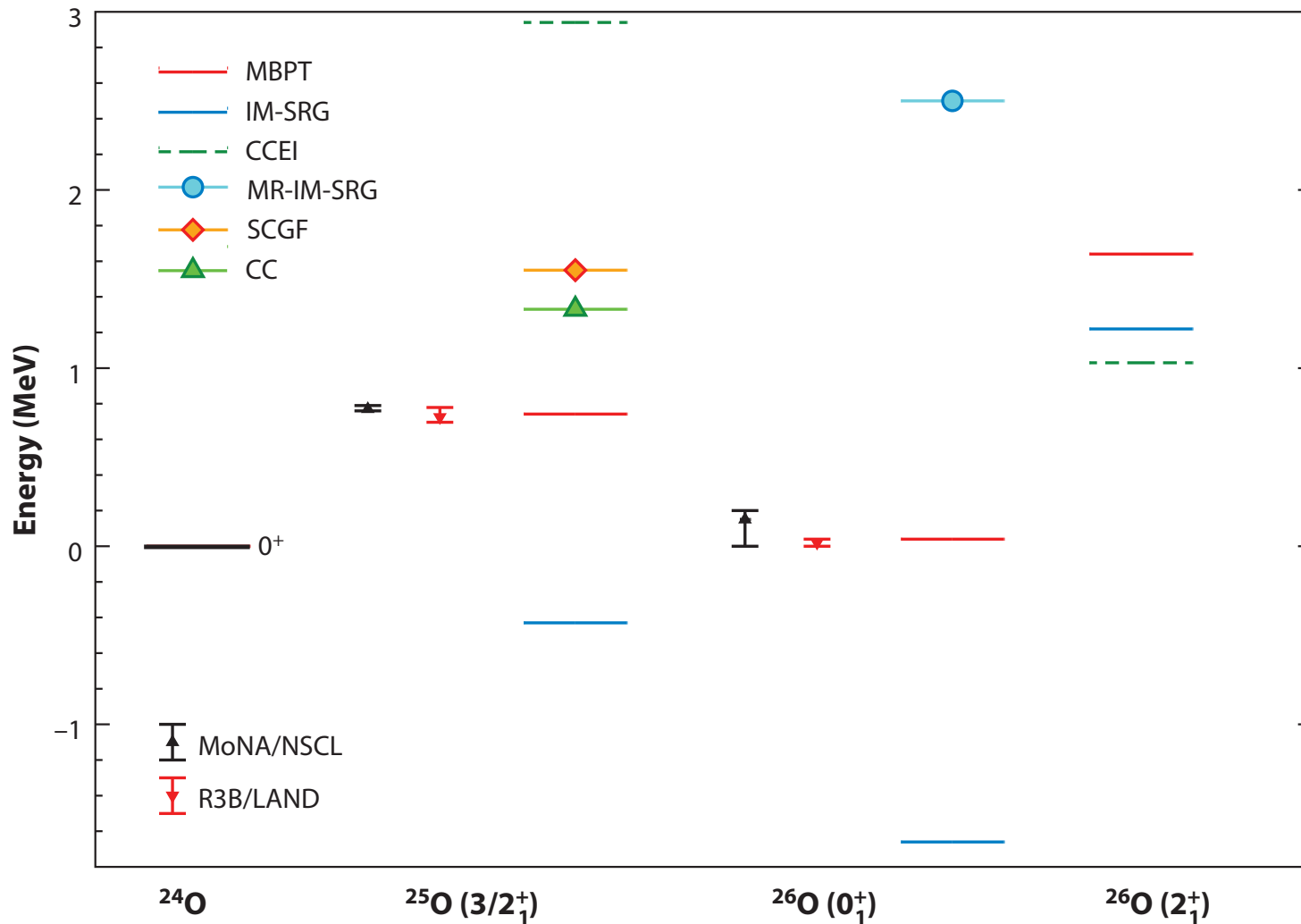


MBPT in extended valence space

IM-SRG/CCEI spectra agree within ~ 300 keV

Beyond the Oxygen Dripline

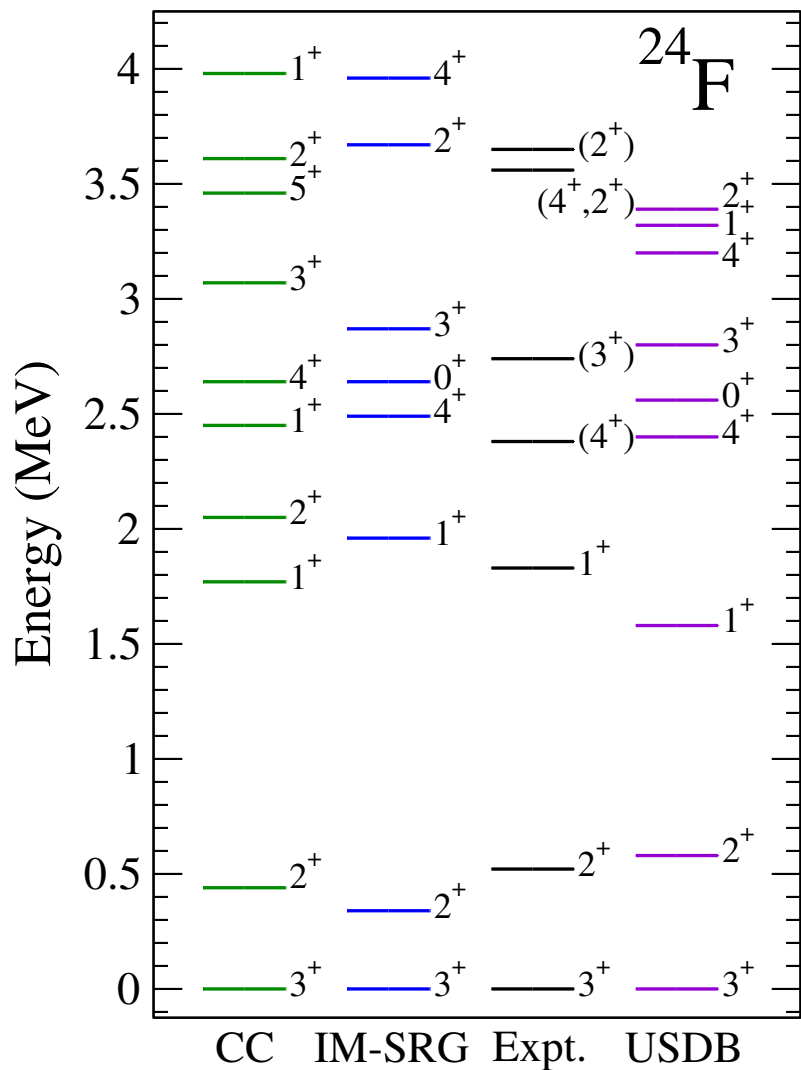
Physics beyond dripline highly sensitive to 3N and continuum effects



Prediction of low-lying 2^+ in ^{26}O (recently measured at RIKEN)

Experimental Connection: ^{24}F Spectrum

^{24}F spectrum: **IM-SRG** (*sd* shell), **full CC**, **USDB**



Ekström et al., PRL (2014)

Cáceres et al., arXiv:1501.01166

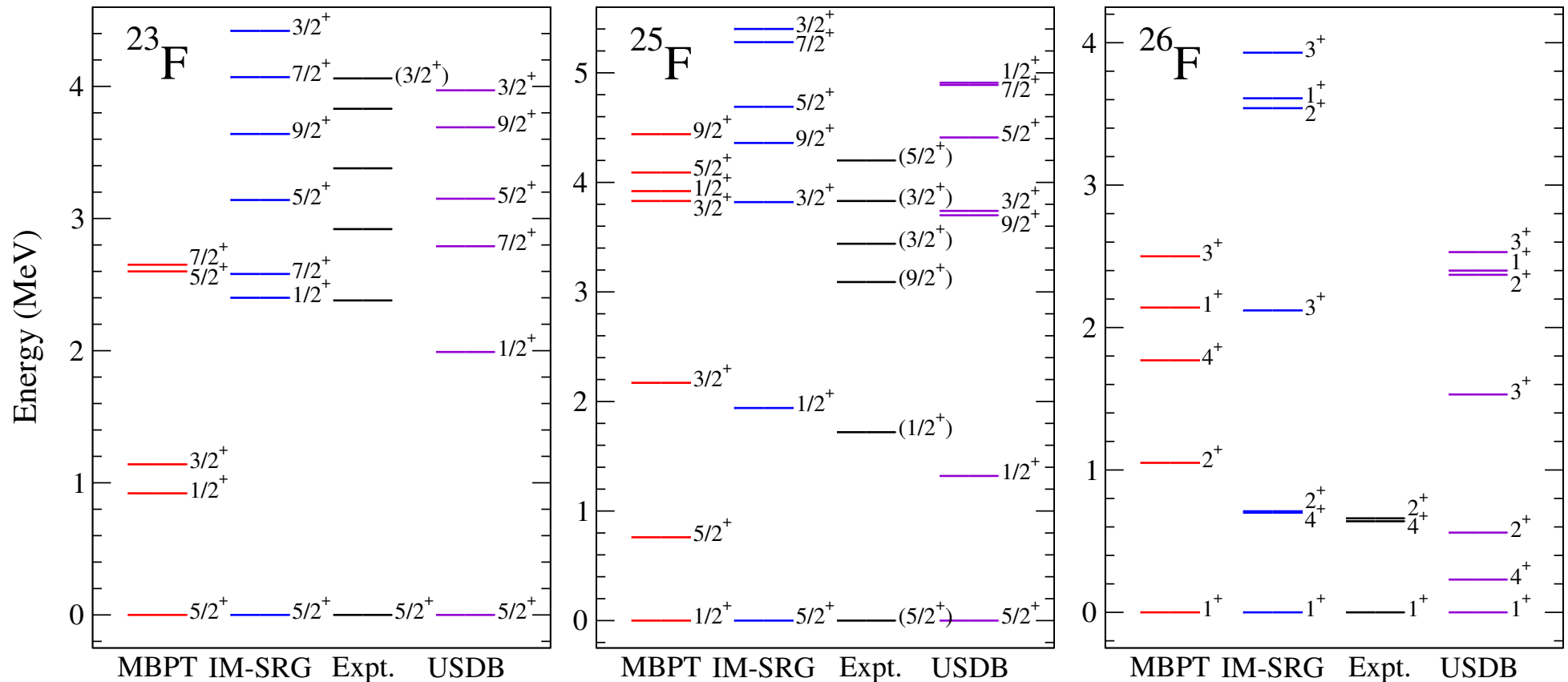
Hebeler, JDH, Menéndez, Schwenk, ARNPS (2015)

New measurements from GANIL

IM-SRG: comparable with phenomenology, good agreement with new data

Fully Open Shell: Neutron-Rich Fluorine Spectra

Fluorine spectroscopy: **MBPT** and **IM-SRG** (*sd* shell) from NN+3N forces



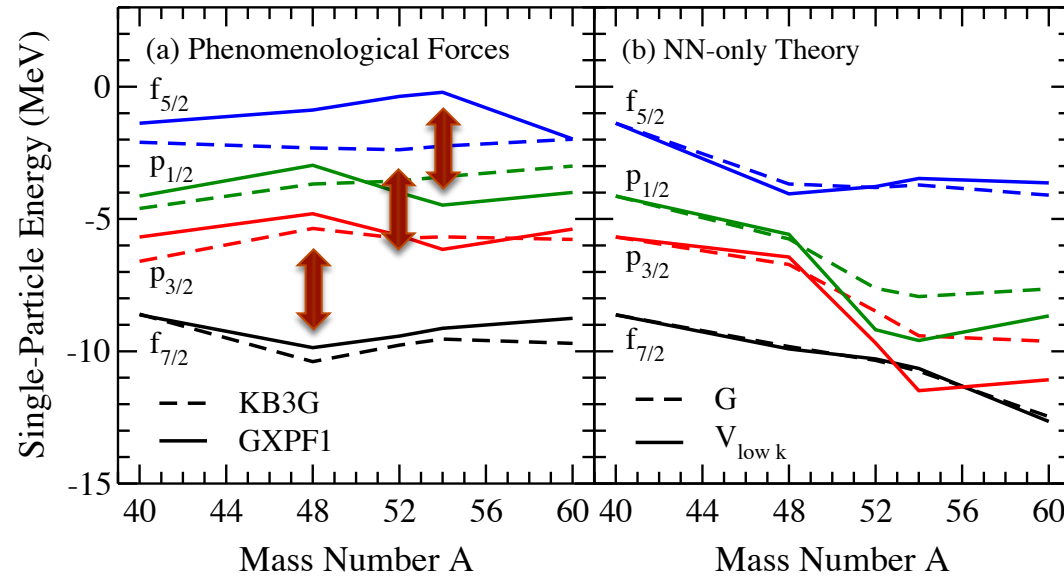
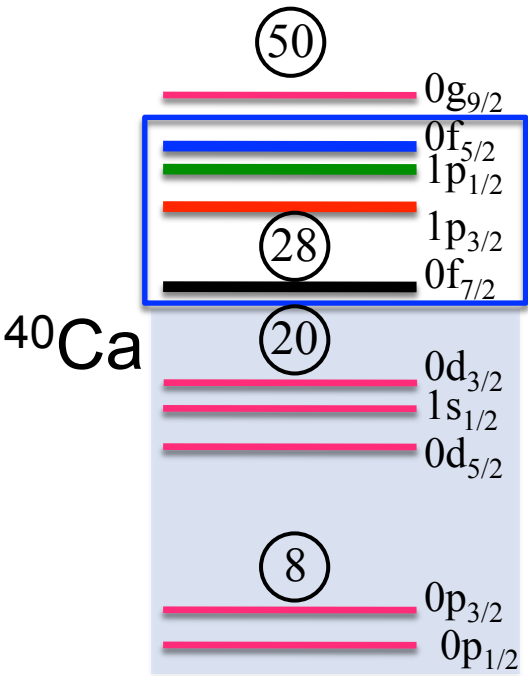
Bogner, Hergert, JDH, Schwenk, in prep.

IM-SRG: **competitive with phenomenology**, good agreement with data

Preliminary results already for scalar operators: charge radii, E0 transitions

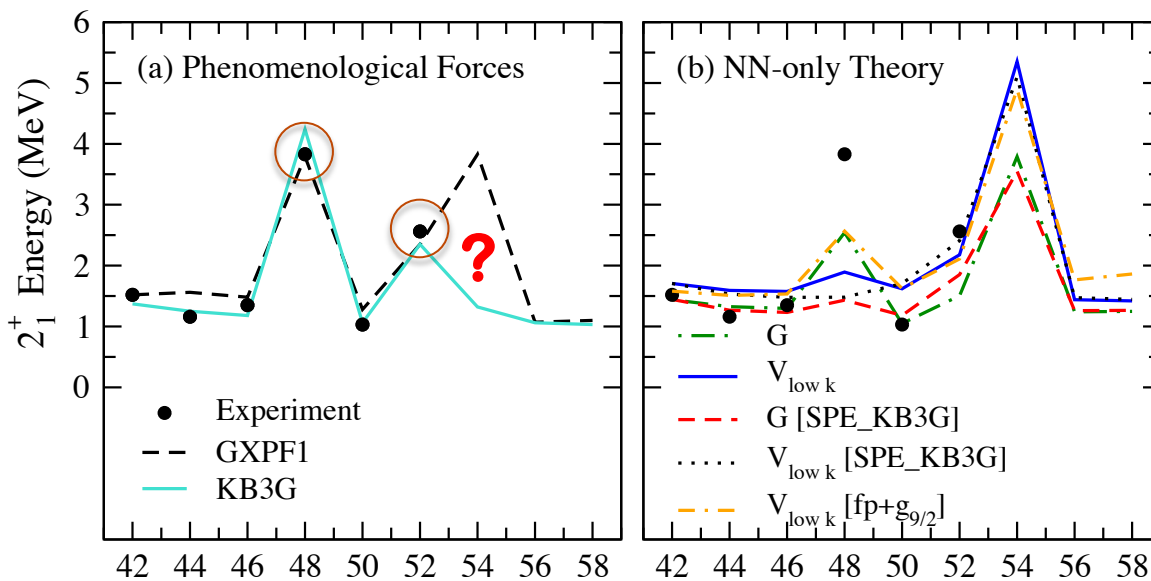
Upcoming: general operators M1, E2, GT, double-beta decay **Stroberg et al.**

Calcium Isotopes: Magic Numbers



GXPF1: Honma, Otsuka, Brown, Mizusaki (2004)

KB3G: Poves, Sanchez-Solano, Caurier, Nowacki (2001)



Phenomenological Forces

Large gap at ^{48}Ca

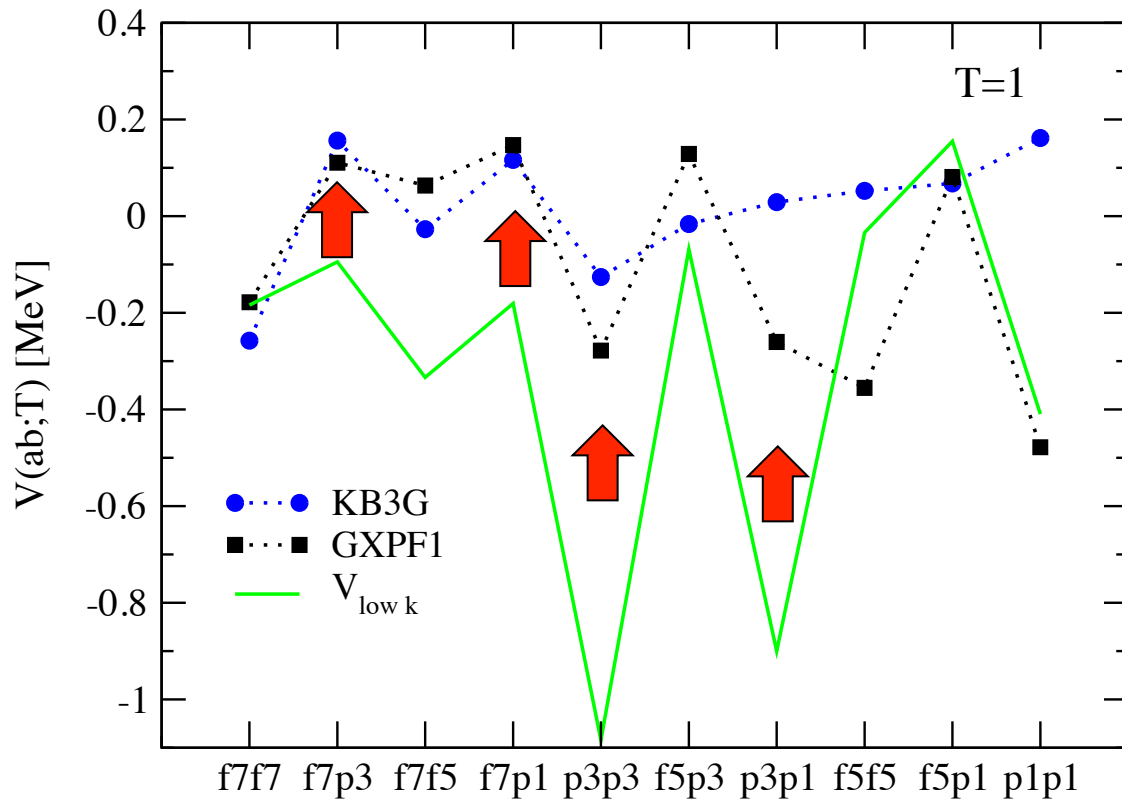
Discrepancy at $N=34$

Microscopic NN Theory

Small gap at ^{48}Ca

N=28: first standard magic number not reproduced in microscopic NN theories

Phenomenological vs. Microscopic



Compare monopoles from:

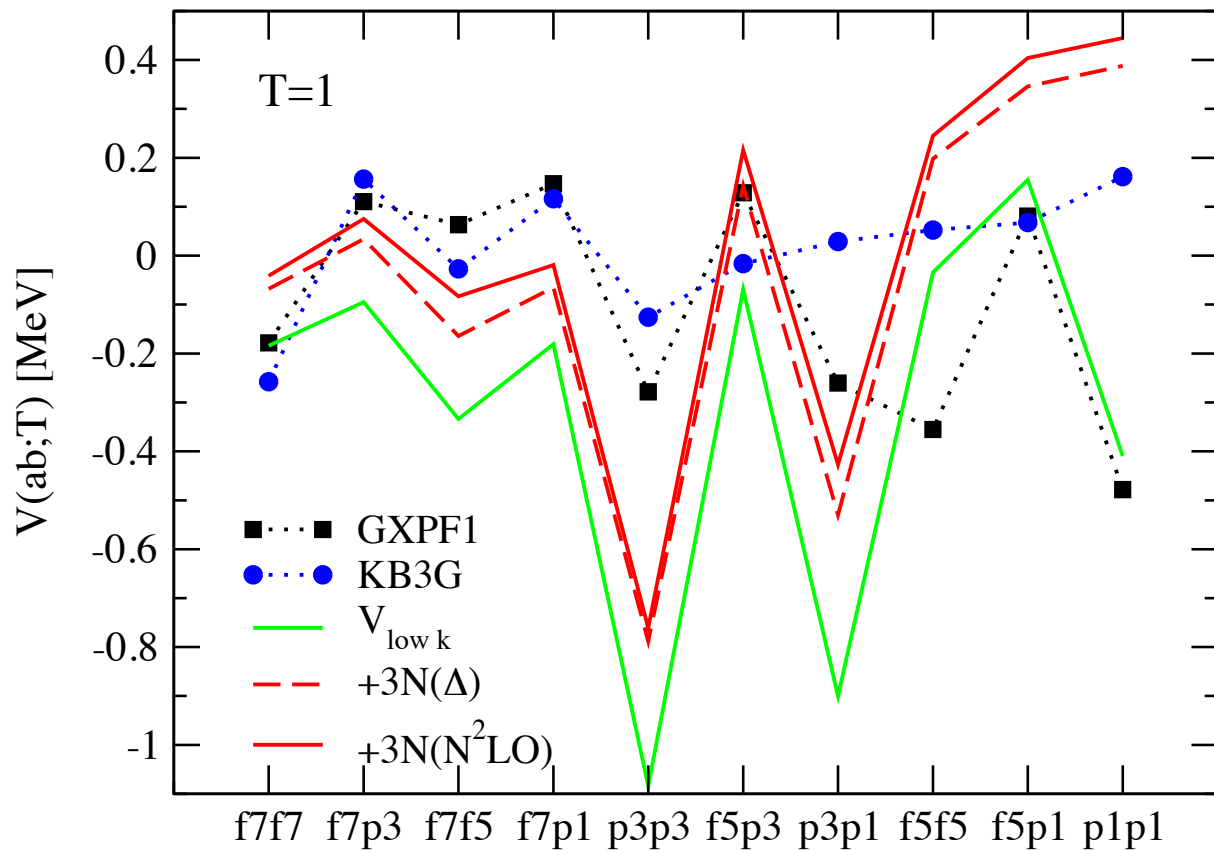
Microscopic **low-momentum** interactions

Phenomenological **KB3G, GXPF1** interactions

Shifts in **low-lying orbitals**:

- $T=1$ repulsive shift

Two-body 3N: Monopoles in pf -shell



Dominant effect from
one- Δ – as expected
from cutoff variation

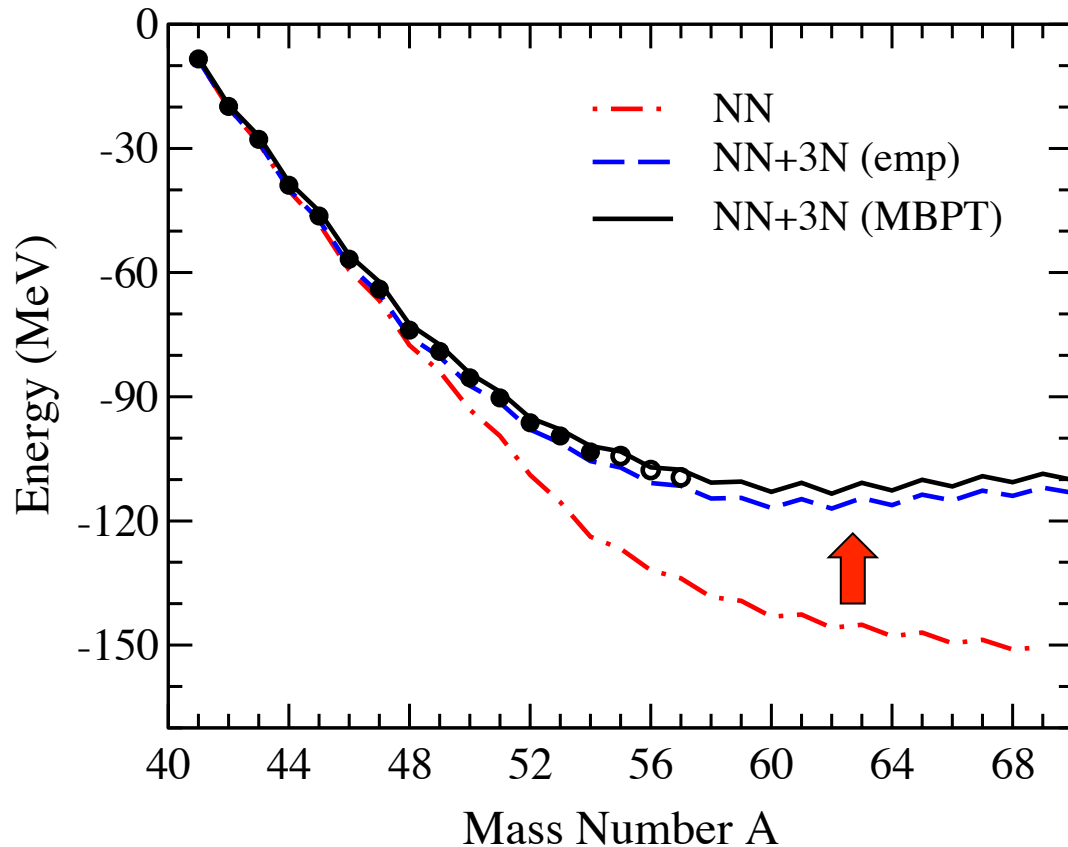
3N forces produce clear
repulsive shift in monopoles

Similar to sd -shell

First calculations to show missing monopole strength due to neglected 3N

Calcium Ground State Energies and Dripline

Signatures of shell evolution from ground-state energies?



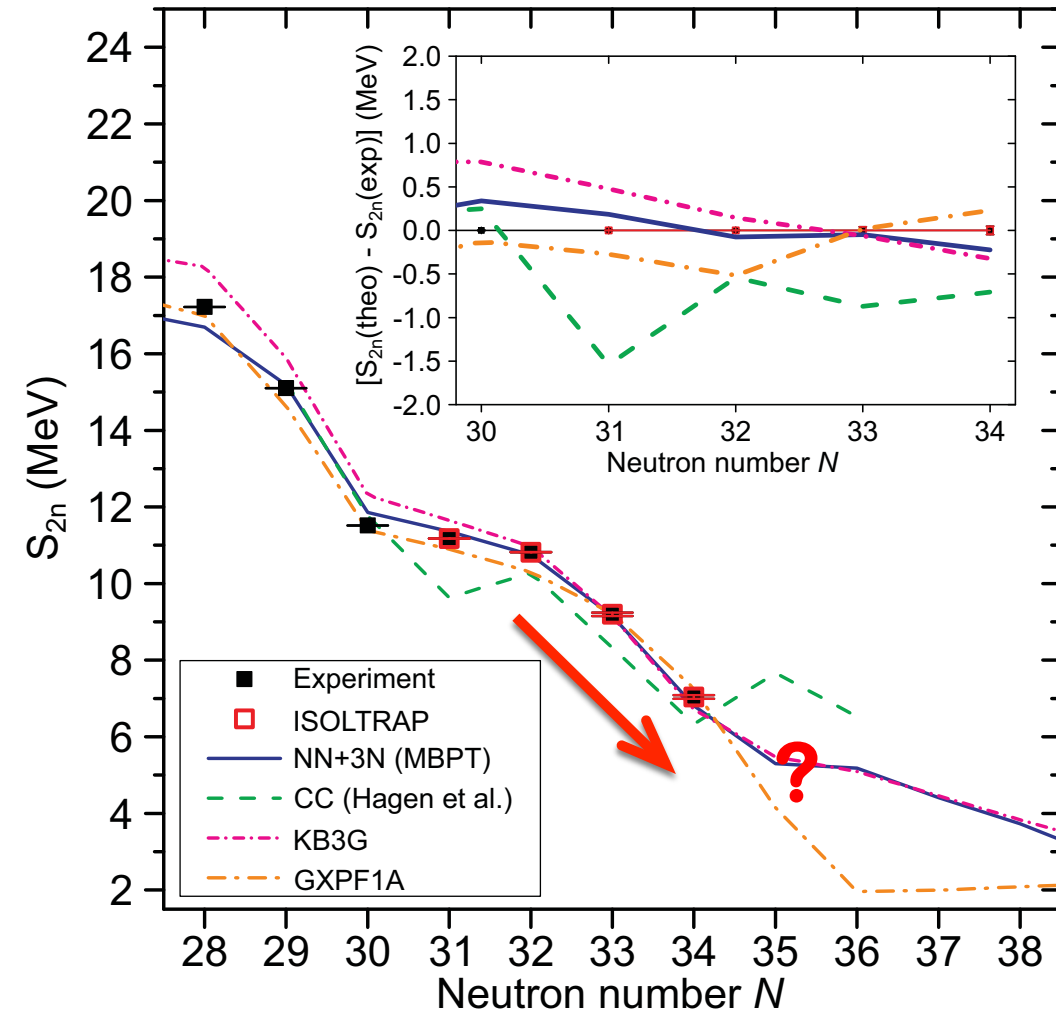
Holt, Otsuka, Schwenk, Suzuki, JPG (2012)

No clear dripline; flat behavior past ^{54}Ca – **Halos beyond ^{60}Ca ?**

$S_{2n} = -[BE(N, Z) - BE(N - 2, Z)]$ **sharp decrease indicates shell closure**

Experimental Connection: Mass of ^{54}Ca

New precision mass measurement of $^{53,54}\text{Ca}$ at **ISOLTRAP**: multi-reflection ToF



Wienholtz et al., Nature (2013)

TITAN Measurement

Flat trend from $^{50-52}\text{Ca}$

Mass ^{52}Ca 1.74 MeV from AME

ISOLTRAP Measurement

Sharp decrease past ^{52}Ca

Unambiguous closed-shell ^{52}Ca

Test predictions of various models

MBPT NN+3N

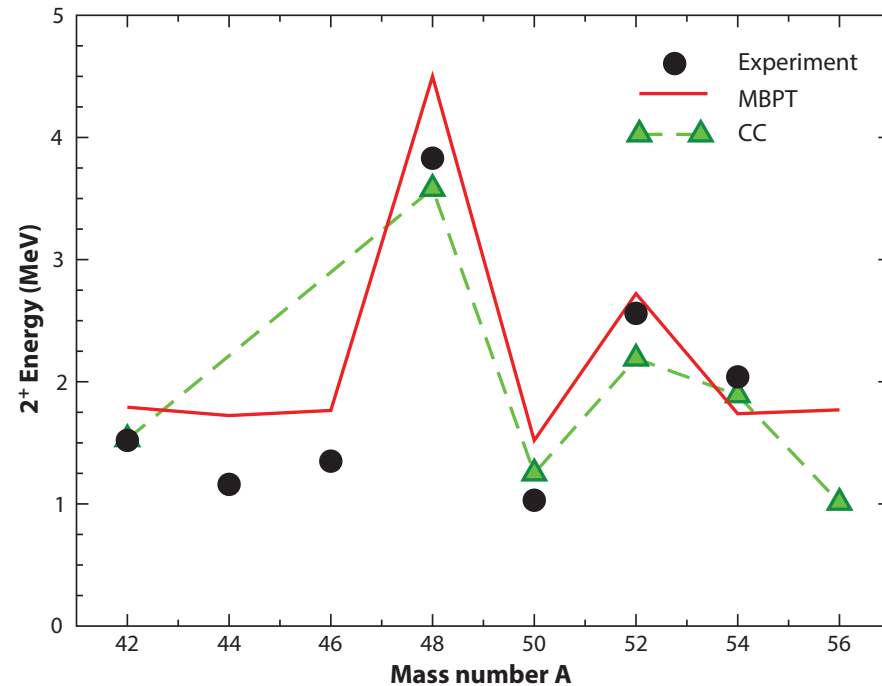
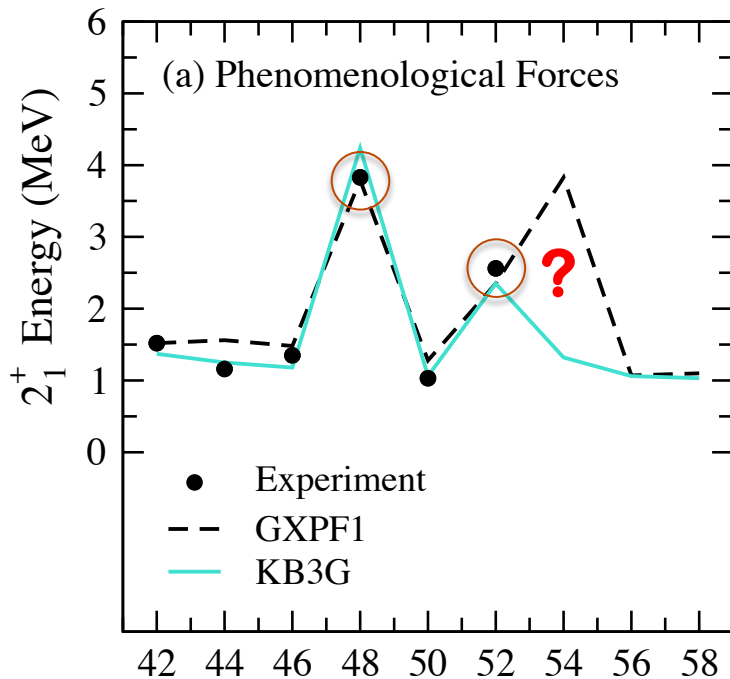
Excellent agreement with new data

Reproduces closed-shell $^{48,52}\text{Ca}$

Weak closed shell signature past ^{54}Ca

N=34 magic number in calcium?

Calcium Isotopes: Magic Numbers



GXPF1: Honma, Otsuka, Brown, Mizusaki (2004)

KB3G: Poves, Sanchez-Solano, Caurier, Nowacki (2001)

Phenomenological Models

Large gap at ^{48}Ca , discrepancy at $N=34$

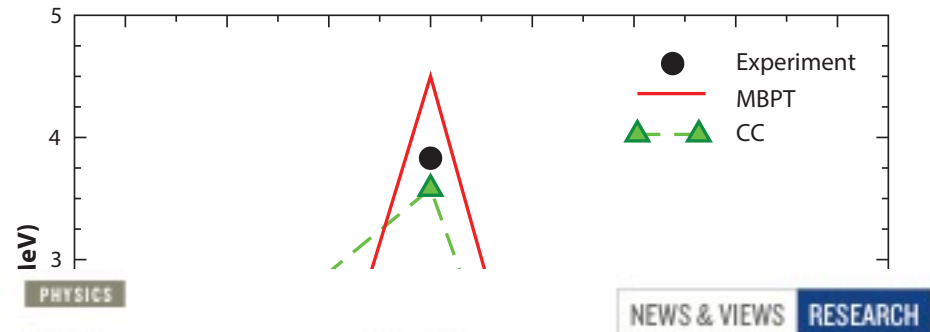
Ab initio theories

Reproduce all new magic numbers, **consistent predictions**

Calcium Isotopes: Magic Numbers



LETTER



Heavy calcium nuclei weigh in

The configurations of calcium nuclei make them good test cases for studies of nuclear properties. The measurement of the masses of two heavy calcium nuclei provides benchmarks for models of atomic nuclei. [SEE LETTER P.346](#)

ALEXANDRA GADE

quarks and gluons, which interact to form

LETTER

doi:10.1038/nature12226

Masses of exotic calcium isotopes pin down nuclear forces

F. Wienholtz¹, D. Beck², K. Blaum³, Ch. Borgmann³, M. Breitenfeldt⁴, R. B. Cakirli^{3,5}, S. George¹, F. Herfurth², J. D. Holt^{6,7}, M. Kowalska⁸, S. Kreim^{3,8}, D. Lunney⁹, V. Manea⁹, J. Menéndez^{6,7}, D. Neidherr², M. Rosenbusch¹, L. Schweikhard¹, A. Schwenk^{7,6}, J. Simonis^{6,7}, J. Stanja¹⁰, R. N. Wolf¹ & K. Zuber¹⁰

doi:10.1038/nature12522

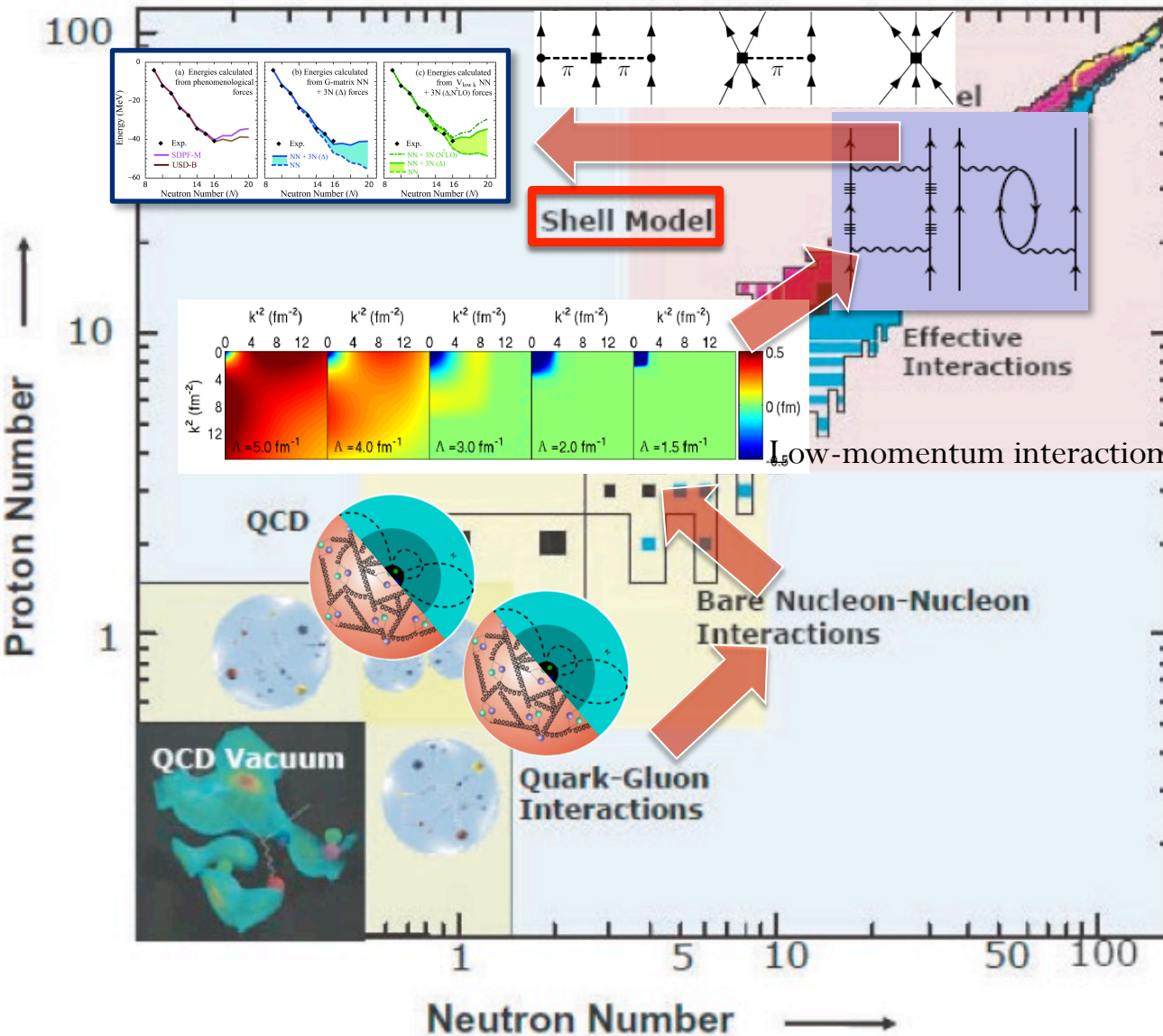
t predictions

Evidence for a new nuclear ‘magic number’ from the level structure of ⁵⁴Ca

D. Steppenbeck¹, S. Takeuchi², N. Aoi³, P. Doornenbal², M. Matsushita¹, H. Wang², H. Baba², N. Fukuda², S. Go¹, M. Honma⁴, J. Lee², K. Matsui⁵, S. Michimasa¹, T. Motobayashi², D. Nishimura⁶, T. Otsuka^{1,5}, H. Sakurai^{2,5}, Y. Shiga⁷, P.-A. Söderström², T. Sumikama⁸, H. Suzuki², R. Taniuchi², Y. Utsuno⁹, J. J. Valiente-Dobón¹⁰ & K. Yoneda²

The Challenge of Microscopic Nuclear Theory

To understand the properties of complex nuclei from elementary interactions



Three-Nucleon Forces

Clear path from symmetries of QCD to shell model

Ideas of:

Effective field theories

Renormalization group

Advances in many-body

Advances in computing

All essential for this progress

Still much to do!!

How will we approach this problem:

QCD → NN (3N) forces → Renormalize → Solve many-body problem → Predictions

New Directions and Outlook

Heavier semi-magic chains: MBPT as guide

Ab initio valence-shell Hamiltonians

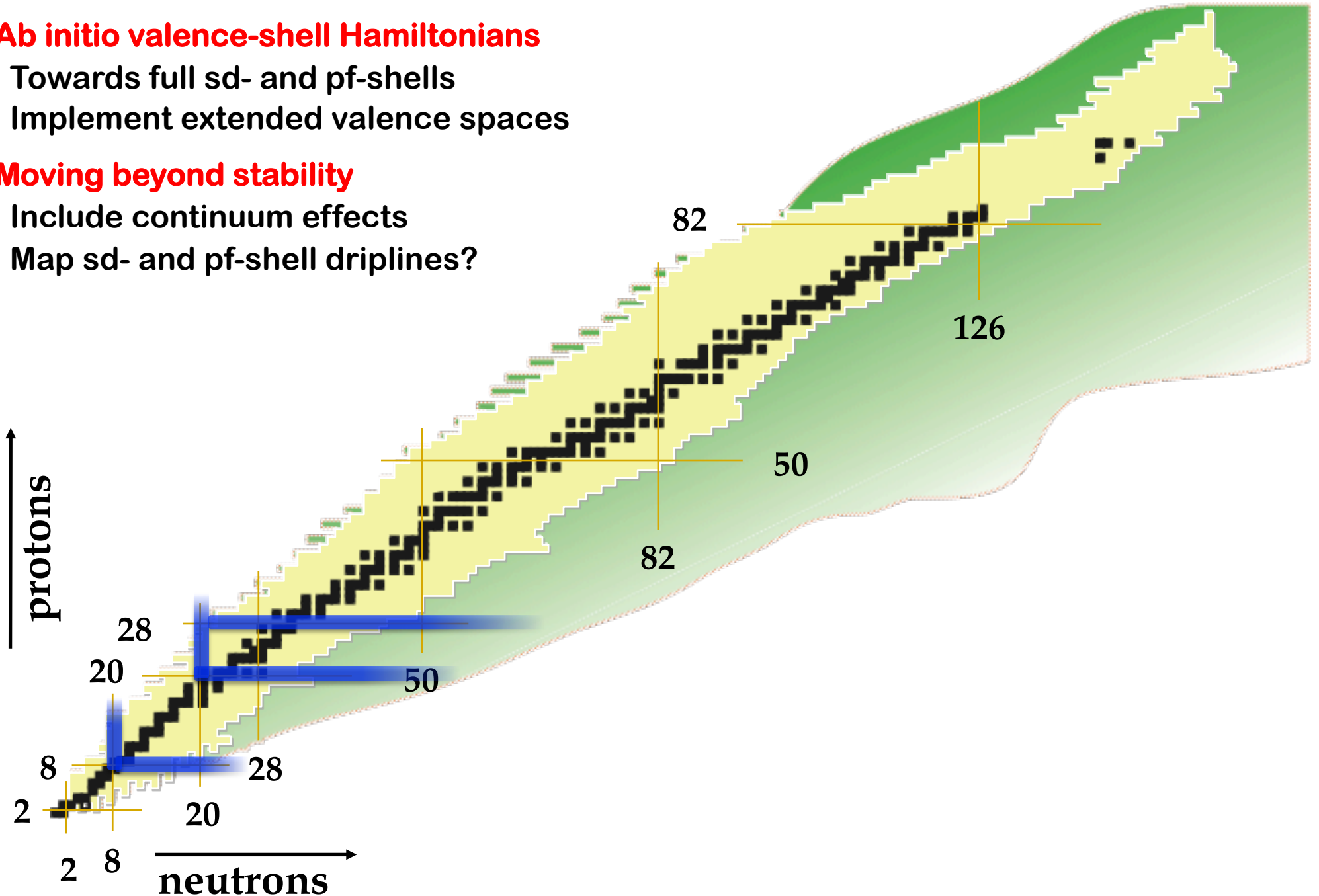
Towards full sd- and pf-shells

Implement extended valence spaces

Moving beyond stability

Include continuum effects

Map sd- and pf-shell driplines?



New Directions and Outlook

Heavier semi-magic chains: MBPT as guide

Ab initio valence-shell Hamiltonians

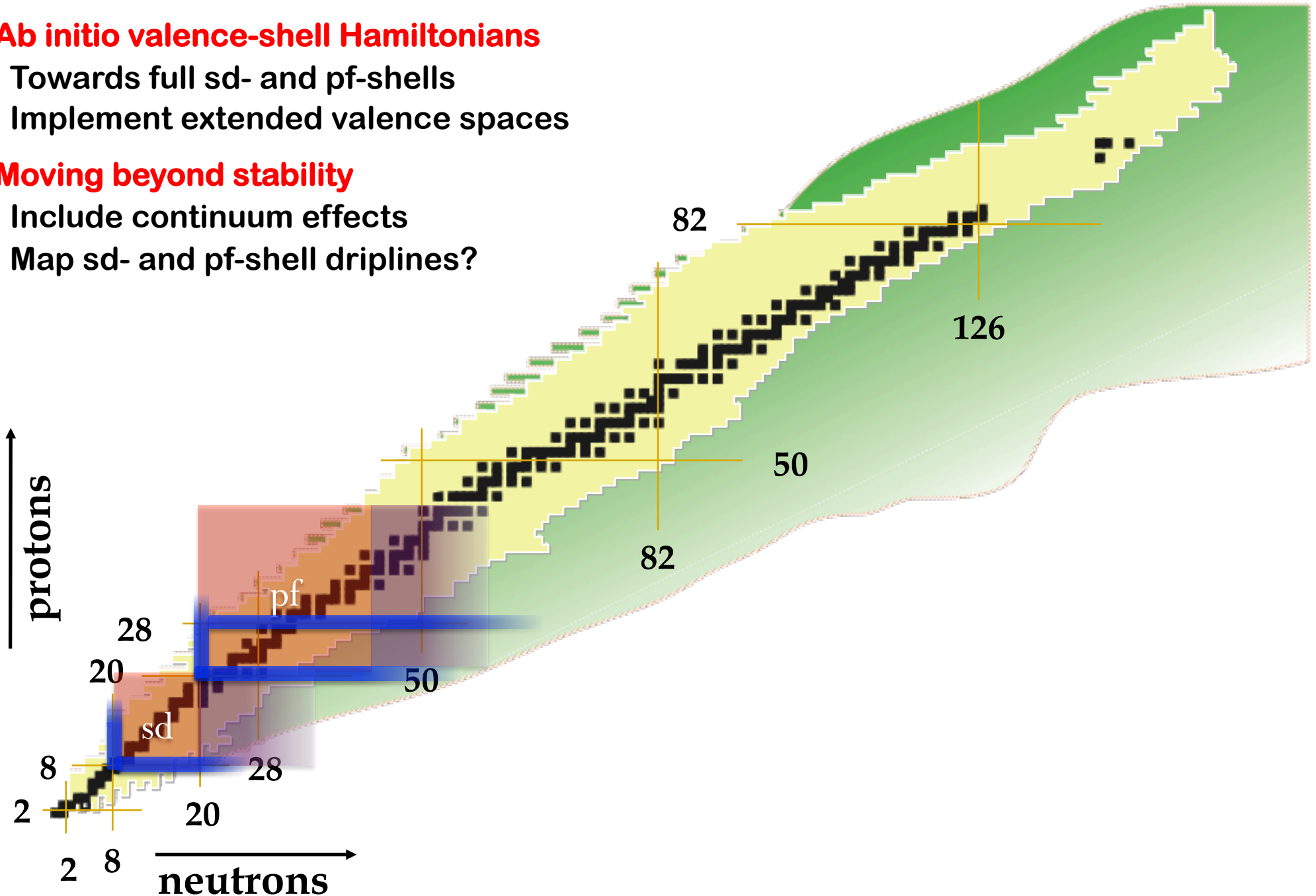
Towards full sd- and pf-shells

Implement extended valence spaces

Moving beyond stability

Include continuum effects

Map sd- and pf-shell driplines?



New Directions and Outlook

Heavier semi-magic chains: MBPT as guide

Fundamental symmetries

Ab initio valence-shell Hamiltonians

Towards full sd- and pf-shells
Implement extended valence spaces

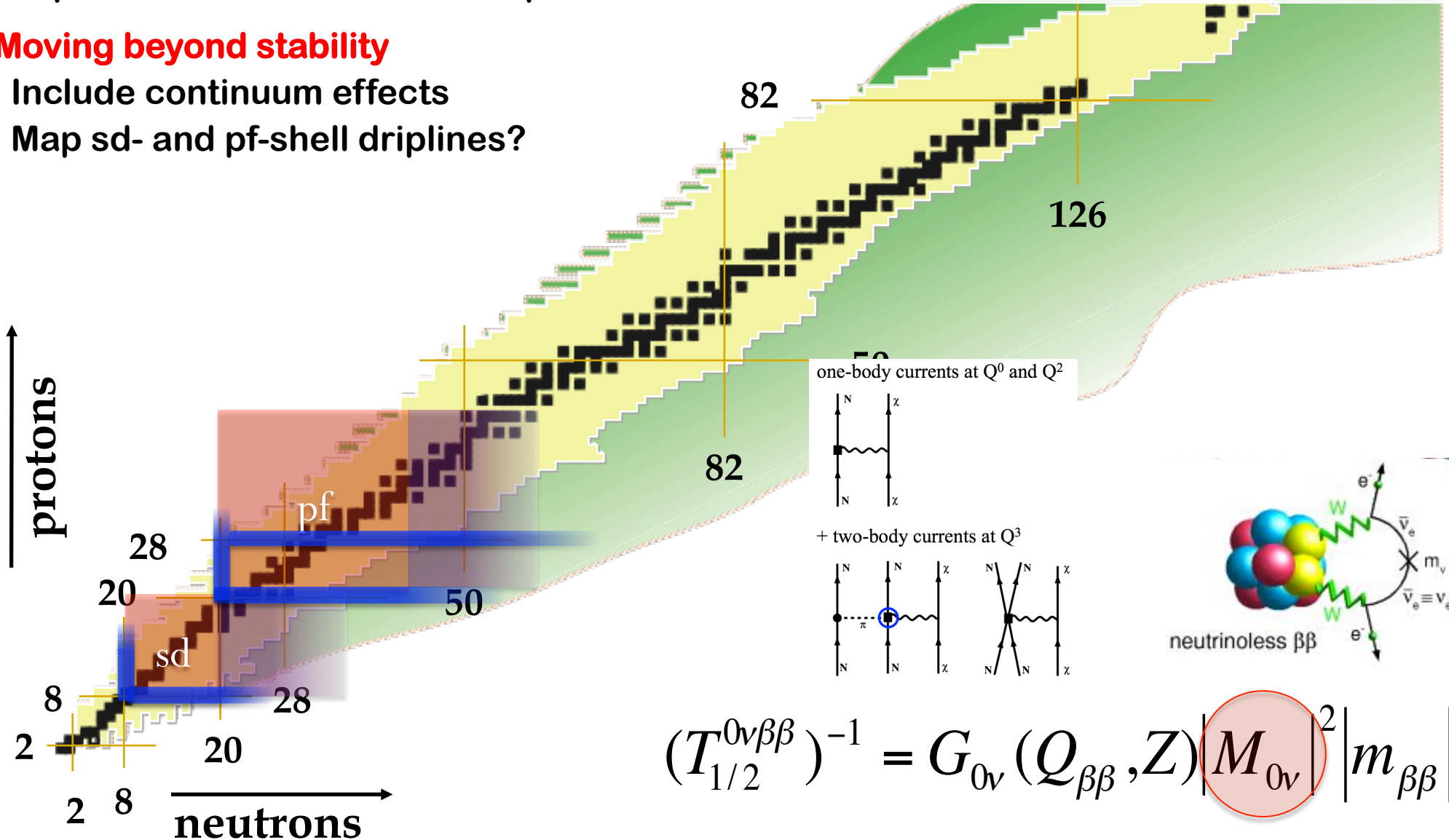
Effective electroweak operators

ab initio calculation of $0\nu\beta\beta$ decay

WIMP-nucleus scattering

Moving beyond stability

Include continuum effects
Map sd- and pf-shell driplines?

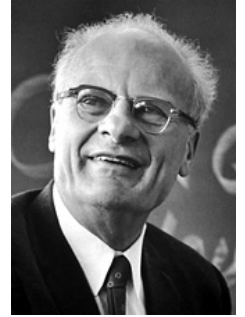


$$(T_{1/2}^{0\nu\beta\beta})^{-1} = G_{0\nu}(Q_{\beta\beta}, Z) \left| M_{0\nu} \right|^2 \left| m_{\beta\beta} \right|^2$$

Final Thought

“Very soft (NN) potentials must be excluded because they do not give saturation; they give too much binding and too high density.”

- *H. Bethe*

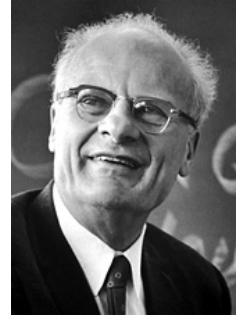


How might you respond?

Final Thought

“Very soft (NN) potentials must be excluded because they do not give saturation; they give too much binding and too high density.”

- *H. Bethe*



How might you respond?

Further Reading

Lepage, nucl-th/9706029 (1997)

Epelbaum, Hammer, Meißner, Rev. Mod. Phys. (2009)

Machleidt, Entem, Phys. Rep. (2011)

Bogner, Furnstahl, Schwenk, Prog. Part. Nucl. Phys. (2010)

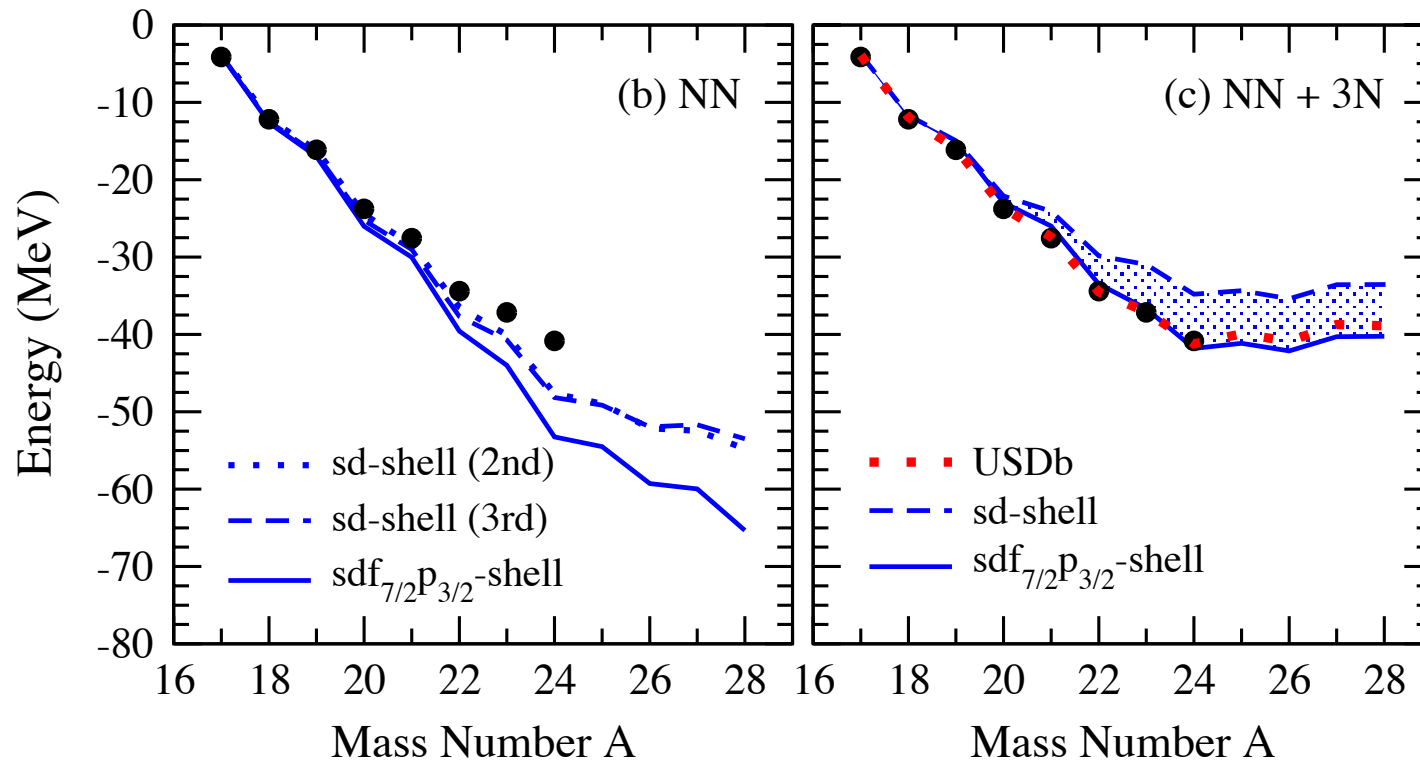
Hebeler, Holt, Menendez, Schwenk, Ann. Rev. Nucl. Part. Sci. (2015)

Thanks to (ie, results, plots, ideas, entire slides, jokes etc., used without citation from):

Scott Bogner, Angelo Calci, Thomas Duguet, Dick Furnstahl, Alex Gezerlis, Gaute Hagen, Kai Hebeler, Heiko Hergert, Herman Krebs, Javier Menendez, Petr Navratil, Achim Schwenk, Johannes Simonis, Ragnar Stroberg

Ground-State Energies of Oxygen Isotopes

Valence-space interaction and SPEs from NN+3N-fit



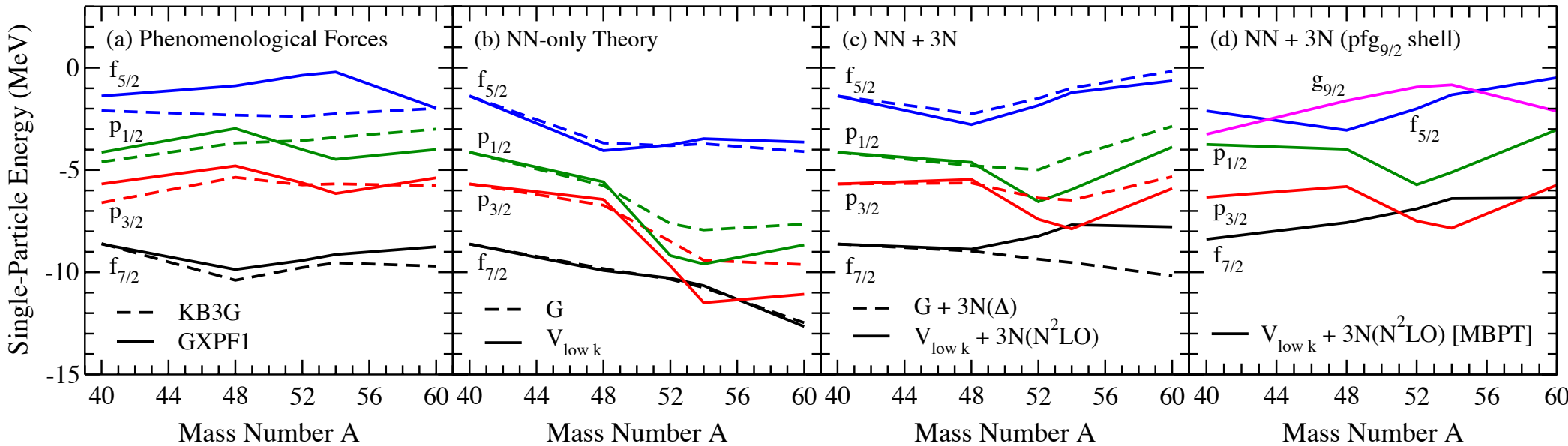
JDH, Menendez, Schwenk, EPJA (2013)

Repulsive character improves agreement with experiment

sd -shell results underbound; improved in **extended space** $sdf_{7/2} p_{3/2}$

Evolution of Shell Structure

SPE evolution with 3N forces in pf and $pf g_{9/2}$ spaces:



NN+3N pf -shell:

JDH, Otsuka, Schwenk, Suzuki JPG (2012)

Trend across: improved binding energies

Increased gap at ^{48}Ca : enhanced closed-shell features

Include $g_{9/2}$ orbit, calculated SPEs

Different behavior of ESPEs (not observable, model dependent)

Small gap can give large 2^+ energy: due to many-body correlations

Duguet, Hagen, PRC (2012)